Long-Range Propagation of the Semidiurnal Internal Tide from the Hawaiian Ridge

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ABSTRACT

The northeastward progression of the semidiurnal internal tide from French Frigate Shoals (FFS), Hawaii, is studied with an array of six simultaneous profiling moorings spanning 25.5°–37.1°N (~1400 km) and 13-yr-long Ocean Topography Experiment (TOPEX)/Poseidon (T/P) altimeter data processed by a new technique. The moorings have excellent temporal and vertical resolutions, while the altimeter offers broad spatial coverage of the surface manifestation of the internal tide’s coherent portion. Together these two approaches provide a unique view of the internal tide’s long-range propagation in a complex ocean environment. The moored observations reveal a rich, time-variable, and multimodal internal tide field, with higher-mode motions contributing significantly to velocity, displacement, and energy. In spite of these contributions, the coherent mode-1 internal tide dominates the northeastward energy flux, and is detectable in both moored and altimetric data over the entire array. Phase and group propagation measured independently from moorings and altimetry agree well with theoretical values. Sea surface height anomalies (SSHAs) measured from moorings and altimetry agree well in amplitude and phase until the northern end of the array, where phase differences arise presumably from refraction by mesoscale flows. Observed variations in SSHA, energy flux, and kinetic-to-potential energy ratio indicate an interference pattern resulting from superposed northeastward radiation from Hawaii and southeastward from the Aleutian Ridge. A simple model of two plane waves explains most of these features.

1. Introduction

Internal tides have been observed for many decades and are the subject of much recent study owing to their potential to mix the deep ocean (e.g., Wunsch 1975; Hendry 1977; St. Laurent and Garrett 2002; Garrett and Kunze 2007). Egbert and Ray (2000, 2001) estimate that about 1 TW of barotropic tidal energy dissipates in the deep ocean over rough topography such as ridges and island chains. The 2500-km-long Hawaiian Ridge is one intensive barotropic-to-baroclinic tide conversion region, over which the energy loss rate of the $M_2$ barotropic tide is estimated to be $20 \pm 6$ GW (e.g., Egbert and Ray 2000, 2001; Rudnick et al. 2003; Lee et al. 2006). Most of this energy is transported away and is associated with low-mode internal tides (e.g., Merrifield and Holloway 2002; Alford 2003; Klymak et al. 2006). Mode-1 internal tides have been observed to propagate thousands of kilometers from their generation sites (e.g., Dushaw et al. 1995; Ray and Cartwright 2001; Alford and Zhao 2007a; Zhao and Alford 2009). Thus, how far the low-mode internal tides propagate (and where and how they dissipate) is a focus of ongoing efforts (e.g., Alford 2003; MacKinnon and Winters 2005; Rainville and Pinkel 2006b; Alford et al. 2007).

In this paper, we examine the long-range propagation of the internal tide northeast of Hawaii as observed using an array of profiling moorings and satellite altimeter measurements. These two techniques are complementary, measuring quite different aspects of the
motions. The moorings give a rich depth–time view of the internal tide at a few discrete locations. By contrast, satellite altimeters offer broad spatial coverage, but measure only the surface manifestation of the coherent, mode-1 component, and that at only \( \approx 250 \) km horizontal resolution. Since the moored array was deployed directly beneath a Ocean Topography Experiment (TOPEX)/Poseidon (T/P) track, we conduct a direct comparison. In spite of the differences between the techniques, moored and altimetric sea surface height anomalies (SSHAs) agree remarkably well, facilitating their combined use to construct a synthesized view.

The altimeter observations are of particular value in interpreting our measurements owing to the known difficulty in interpreting point measurements in the presence of multidirectional wave fields (Nash et al. 2006; Martini et al. 2007; Rainville et al. 2010). That is, in contrast to a single propagating wave, which shows constant energy flux, velocity ellipses elongated in the direction of wave propagation, and a constant ratio of horizontal kinetic to available potential energy (HKE/APE), superposed wave trains exhibit interference patterns where all of these quantities change depending on their spatial locations. Our measurements indicate a superposition of northeastward radiation from multiple sources at the Hawaiian Ridge and southeastward radiation from the Aleutian Ridge. A simple two-wave model reproduces many of the observed aspects of the observed interference pattern.

After giving a detailed description of the moored and altimetric data and their processing (section 2), our goals in this paper are threefold: first, to present a qualitative picture of the internal tide in the region (section 3); second, to conduct a detailed comparison of the moored and altimetric techniques (section 4); and finally to demonstrate the manifestation of the Hawaii–Aleutian interference pattern in the moored and altimetric measurements (section 5). A summary (section 6) and a discussion (section 7) follow.

2. Data and methods

From the moored data, we seek to characterize the modal content and the degree of coherence with astronomical forcing of the semidiurnal internal tide versus range from the generation site. To do this, the semidiurnal signals are first isolated via bandpass filtering. The bandpassed velocity and isopycnal displacement signals are then decomposed into the five lowest baroclinic modes. The semidiurnal signals are further decomposed into coherent and incoherent constituents. These methods, mostly quite standard, are described in this section, followed by a description of the T/P altimeter SSHA data and processing methods.

a. The IWAP experiment

The Internal Waves Across the Pacific (IWAP) experiment (Fig. 1; Alford et al. 2007) was designed to follow the northeastward internal tidal beam originating from French Frigate Shoals (FFS), Hawaii, identified in numerical model results from the Primitive Equations Z-coordinate Harmonic Analysis of Tides (PEZHAT) (Fig. 1, yellow arrows, data courtesy of E. Zaron) and Princeton Ocean Model (POM) (data courtesy of S. Johnston, not shown) datasets. IWAP’s focus was on the internal tide’s long-range propagation, complementing the Hawaii Ocean Mixing Experiment (HOME), which focused on the internal tide’s generation and local dissipation at the Hawaiian Ridge. Measurements were conducted spanning moored profiler 1 (MP1) at 25.5°N, about 200 km from the ridge, and MP6 at 37.1°N, over 1600 km away. The experiment consisted of two cruises on board the R/V Revelle: the first was from 18 April to 25 May 2006, and the second from 1 to 18 June 2006. Its primary components were 1) an array of six profiling moorings (black dots showing MP1–MP6, described in Tables 1 and 2), 2) a spatial survey (white line), and 3) three high-resolution time series (blue dots). Only the moored data are discussed here.

b. Profiling mooring measurements

The six moorings were identical in their instrument arrangements (see Fig. 2a). From the surface to 40 m, an upward-looking 300-kHz RDI ADCP on the subsurface float measured velocity profiles every 2 min. Between 85 and 1400 m, a McLane MP climbed up and down through the water column along the mooring wire, completing one up or down profile each 1.5 h (left panels of Fig. 3, green line). Each carried a Falmouth Scientific CTD and acoustic current meter, delivering profiles of density and velocity with a vertical resolution of about 2 m (Doherty et al. 1999). At \( \approx 3000 \) m, an Aanderaa RCM8 current meter measured the horizontal velocity every 10 min, and a Sea-Bird temperature logger measured the temperature twice every minute. High wire tension (\( \approx 500 \) kg), together with generally weak currents, ensured mooring knockdown was less than 20 m at all sites.

The instruments generally functioned well, with the exception of MP4, where the MP stopped profiling after 23 days because of a broken drive axle, and MP5, where the MP sampled its full range for 4 days and afterward only at 700–1400 m owing to heavy ballasting. In addition, the upward-looking ADCP data at MP5 were lost because of a corrupt flash memory card. Thus, the durations of full-depth time series at MP1, MP2, MP3, and MP6 range between 40 and 50 days, but only 23 and 4 days for MP4 and MP5, respectively (Table 1).
The measured horizontal velocity is decomposed into along-beam ($u$) and cross-beam ($v$) components, with the along-beam direction defined toward 60° north of east, the mean direction of the FFS internal tidal beam. Samples of raw data at MP1, MP3, and MP6 are shown in the left panels of Fig. 3. Colors indicate the along-beam velocities obtained from the ADCP (0–40 m), MP (85–1400 m), and Aanderaa (≈3000 m). The Aanderaa-measured velocities are plotted in a 200-m-thick layer on the same color scale. A rich field of low-frequency and near-inertial flows is seen. These flows largely obscure tidal signals, which will be isolated below with harmonic analysis and bandpass filtering.

The sawtooth pattern traced by MP leads to a variable temporal resolution pattern ranging from 3 h at 85 and 1400 m to 1.5 h at middepths (left panels of Fig. 3, green line). The ADCP, Sea-Bird, and Aanderaa data have much higher temporal resolutions, ranging between 0.5 and 10 min. All data are gridded onto uniform 1.5-h temporal and 2-m vertical grids using linear interpolation.

c. Bandpass filtering

The semidiurnal components of the mooring data are extracted by bandpassing data at each depth with a fourth-order Butterworth filter centered at the $M_2$ tidal frequency ($2.23 \times 10^{-5}$ s$^{-1}$) with zero-phase response.
and quarter-power points at \(2.01, 2.47 \times 10^{-5} \text{ s}^{-1}\), that is, \(1.73 – 2.13 \text{ cpd}\). This passband includes the \(M_2\) and \(S_2\) tidal constituents but rejects inertial motion, whose frequency ranges between 0.99 and \(1.39 \times 10^{-5} \text{ s}^{-1}\) in this latitudinal range.

Semidiurnal signals obtained in this manner are plotted in the right panels of Fig. 3. Note that barotropic currents are also removed as described below (section 2e). Baroclinic semidiurnal currents with amplitudes \(O(10 \text{ cm} \text{ s}^{-1})\) are now seen. Opposing shallow–deep flows at MP1 (Fig. 3b) indicate the dominance of mode-1 motions, with higher-mode contributions becoming more evident to the north (Figs. 3d and 3f). These signals will be considered in detail in section 3.

d. Vertical displacement

From the surface to 1400 m, displacement of a set of isopycnals evenly spaced in depth is computed by linear interpolation of the MP-measured density profiles. Signals associated with long-term density changes are small, but are removed by means of a sliding 5-day window. At \(\approx 3000 \text{ m}\), isothermal displacement (assumed equal to the isopycnal displacement) is calculated from the bandpassed Sea-Bird temperature \(T(t)\) by \(\eta(t) = T(t)/T_z\), where \(T_z = 2 \times 10^{-4} \text{ °C m}^{-1}\) is the vertical temperature gradient at \(\approx 3000 \text{ m}\) from the annual-mean hydrography in the 2005 \textit{World Ocean Atlas} (WOA05; Antonov et al. 2006; Locarnini et al. 2006). Raw and semidiurnally bandpassed displacements are shown in Fig. 3. Semidiurnal signals (right panels) are more visible in Fig. 3 in the raw displacement than in the raw velocity (left panels), owing to the smaller near-inertial contributions to the displacement.

e. Modal decomposition

The bandpassed semidiurnal signals are next projected onto baroclinic modes. In an ocean of depth \(H\), internal tides can be represented by a superposition of discrete baroclinic modes that depend on the buoyancy frequency profile \(N(z)\). The baroclinic modes for vertical displacement, \(\Phi(z)\), can be determined by the eigenvalue equation,

\[
d^2\Phi(z)/dz^2 + N^2(z)/c_n^2 \Phi(z) = 0 \quad (1)
\]

subject to the boundary conditions \(\Phi(0) = \Phi(-H) = 0\), where \(n\) is the mode number and \(c_n\) is the eigenspeed (Gill 1982).

The corresponding baroclinic modes for pressure and horizontal velocity \(\Pi(z)\) are related to \(\Phi(z)\) via

\[
\Pi(z) = \rho_0 c_n^2 d\Phi(z)/dz. \quad (2)
\]
\[ \Phi(z) = -\frac{1}{N^2(z) \rho_0} \frac{d\Pi(z)}{dz}, \]  

where \( \rho_0 \) is the water density. Here, \( N(z) \) is taken from the annual-mean hydrography from WOA05. The climatological stratification profile (Fig. 2b, gray) is consistent with the MP measurement (dashed) except in the upper 400 m. This difference has little impact on the modal shapes (Figs. 2c and 2d).

Phase velocity \( c_p \) and group velocity \( c_g \) can be derived from the eigenspeed \( c_n \) following the dispersion relation (Rainville and Pinkel 2006b),

\[ c_p = \frac{\omega}{(\omega^2 - f^2)^{1/2}} c_n, \]  

and

\[ c_g = \frac{(\omega^2 - f^2)^{1/2}}{\omega} c_n, \]  

where \( \omega \) is the \( M_2 \) or \( S_2 \) tidal frequency and \( f \) is the inertial frequency. We will show in section 4b that the mooring-observed propagation of the semidiurnal internal tide agrees very well with the theoretical values.

The mooring data are highly resolved in the upper 1400 m, equivalent to 62%–68% of the Wentzel–Kramers–Brillouin (WKB) stretched water column (depending on water depth), allowing resolution of the first five modes (Nash et al. 2005). The gaps above and below the \( \sim 3000 \)-m measurements render higher-mode fits unstable in the modal decomposition. Velocity at 3000 m is given a higher weight owing to the longer vertical scales at depth. Error estimates in the modal decomposition are discussed in the appendix.

Five-mode solutions were computed for all moorings except for MP5. At MP5, the failure of ADCP and MP to sample the upper 700 m after 4 days precludes velocity solutions owing to the sensitivity of the velocity fits to the near-surface gaps (Nash et al. 2005). Hence, velocity, energy, and flux results at MP5 are based on the first 4 days alone (preventing separate resolutions of \( M_2 \) and \( S_2 \)). However, mode-1 displacement solutions using deep data alone are stable for the whole time series (due to the location of the mode-1 maximum near 1300 m; Fig. 2d), allowing comparison of moored SSHA with altimetric observations at MP5 as described in section 4. In support of their validity, mode-1 displacements

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**Fig. 2.** Mooring instrumentation and ocean stratification profiles at MP1: (a) mooring instruments and measured ocean parameters and (b) buoyancy frequency profiles (in cph). The black line is from the MP measurements. The gray line is from the WOA05 annual-mean hydrographic data. (c) Normalized vertical structures of the first five baroclinic modes for pressure and horizontal velocity. (d) As in (c), but for vertical displacement and vertical velocity.
computed during the first 4 days using all data and using data deeper than 700 m alone agree well with each other (not shown).

The baroclinic velocity and displacement are expressed as

\[
\mathbf{u}'(z, t) = \sum_{n=0}^{5} \mathbf{u}_n(t) \Phi_n(z),
\]  

(6)

Fig. 3. Samples of (left) raw measurements and (right) bandpassed semidiurnal signals at (a),(b) MP1, (c),(d) MP3, and (e),(f) MP6. Colors indicate along-beam velocities from ADCP (0–40 m), MP (85–1400 m), and Aanderaa (~3000 m), with the scale indicated at the upper right. Pink lines are the Sea-Bird temperature data at ~3000 m depth. Green lines in the left panels show the MP sampling patterns. Isopycnal displacements with mean spacing 100 m are overplotted in black.
where $\Phi_n(z)$ and $\Pi_n(z)$ are the vertical structures of the baroclinic modes (Figs. 2c and 2d) and $u_n'(t)$ and $\eta_n'(t)$ represent the time-varying magnitudes of the baroclinic modes. At each time, $u_n'(t)$ and $\eta_n'(t)$ are extracted from the velocity and displacement profiles by least squares modal fitting (Alford 2003; Nash et al. 2005). For the velocity, the zeroth-mode, or barotropic, solution is also allowed ($n = 0$).

f. Pressure, SSHA, energy, and flux

These quantities can be calculated for each baroclinic mode independently following the same procedure. For clarity, we neglect the subscript $n$, which denotes the mode number, in the following equations. Following Nash et al. (2005), the baroclinic pressure perturbation, $p'(z, t)$, is calculated from the modal fit displacement $\eta'(z, t)$ by

$$
p'(z, t) = \rho_0 \int_{-z}^{0} N^2(\hat{z}) \eta'(\hat{z}, t) d\hat{z} - p_{surf}'(t). \quad (8)
$$

The definition of baroclinic motions requires the depth average of the baroclinic pressure perturbation to be zero; therefore, the sea surface pressure, $p_{surf}'(t)$, can be calculated as

$$
p_{surf}'(t) = \rho_0 \int_{-H}^{0} N^2(z) \eta'(z, t) dz. \quad (9)
$$

SSHA caused by the internal tide can be calculated from the sea surface pressure $p_{surf}'$.

$$
SSHA = \frac{p_{surf}'}{\rho_0 g}, \quad (10)
$$

where $g$ is the acceleration due to gravity.

Depth-integrated horizontal kinetic energy and available potential energy are computed from the above modal fit baroclinic perturbations, $u'(z, t)$ and $\eta'(z, t)$, by

$$
HKE = \frac{1}{2} \rho_0 \int_{-H}^{0} \langle |u'(z, t)|^2 \rangle dz, \quad (11)
$$

and

$$
APE = \frac{1}{2} \rho_0 \int_{-H}^{0} \langle N^2(z) \eta'^2(z, t) \rangle dz, \quad (12)
$$

where the angle brackets indicate the average over one tidal cycle. The total energy $E$ is calculated by

$$
E = HKE + APE. \quad (13)
$$

The depth-integrated energy flux is computed as the covariance of the modal fit velocity and pressure perturbations by

$$
F = \int_{-H}^{0} \langle u'(z, t) p'(z, t) \rangle dz. \quad (14)
$$

g. Coherent portion

For each mode, the bandpassed semidiurnal internal tide is further decomposed into coherent and incoherent constituents. Neglecting the much weaker $N_2$ and $K_2$ constituents, coherent $M_2$ and $S_2$ signals are harmonically fit using tide analysis software (Pawlowicz et al. 2002). Because both the modal fitting and harmonic fitting operations are linear, their order makes no difference. We chose to conduct harmonic analyses on the modal fits.

The baroclinic velocity and pressure signals are thus expressed as

$$
(u', p')_{semi} = (u', p')_{M_2} + (u', p')_{S_2} = (u', p')_{in}, \quad (15)
$$

where $(u', p')_{in}$ represent the incoherent constituents.

For each constituent, the energy and flux are calculated following the procedure described in section 2f. According to Eq. (14), the coherent flux is calculated as

$$
\langle u'_M M_2, p'_M M_2 \rangle + \langle u'_S S_2, p'_S S_2 \rangle + \langle u'_M S_2, p'_S M_2 \rangle. \quad (16)
$$

The first and second terms are the $M_2$ and $S_2$ fluxes, respectively. The two cross terms between $M_2$ and $S_2$ lead to the spring–neap cycles. If averaged over an integer number of spring–neap cycles ($\approx 14.7$ days), the cross terms will cancel.

The coherent fractions (CFs) comprising the energy and flux magnitude are then computed for each mode as

$$
CF_E = \frac{E_{(M_2+S_2)}}{E_{total}}, \quad (17)
$$

and

$$
CF_F = \frac{|F_{(M_2+S_2)}}{|F_{total}|}. \quad (18)
$$

This definition is straightforward for energy. However, because flux is a vector quantity, the definition for flux overestimates the coherent fraction in its exclusion of variations in direction. That is, signals with constant magnitude but wandering angle would give $CF_F = 1$. 
It is known that record length affects the partitioning between incoherent and coherent signals. The record lengths of MP1, MP2, MP3, and MP6 range between 40 and 50 days, while the MP4 time series is only 23 days. The definition of a coherent fraction in this manner may cause confusion. That is, incoherent signals owing to fluctuations in amplitude and phase on time scales longer than the deployment time are not distinguishable as such. This will be kept in mind when comparing coherent signals from the mooring data (~2 months) and the T/P altimeter data (~10 yr).

h. Altimeter data and methods

The IWAP mooring array lies directly under T/P track 249 (Fig. 1, black dots and line). In section 4a, the altimetric results along track 249 will be compared with moored results. Altimetric SSHA is computed following the work of Zhao and Alford (2009), which employs data from both the T/P original tracks and tandem tracks, resulting in higher spatial resolution (~250 km). Briefly, along-track SSHA data spanning 1993–2005 are processed with standard corrections for instrumental bias, environmental effects, and the barotropic tide (Ray 1999; Berwin 2003b). Along each track, data are resampled onto a 6.2-km spatial grid (Berwin 2003a), and bandpass filtered with spatial cutoffs of (120, 400) km to isolate the mode-1 internal tide (~160 km in our region).

The $M_2$ and $S_2$ signals in the T/P SSHA time series have alias periods of 62.1 and 58.7 days, respectively. They can be extracted by harmonic analysis at each location (e.g., Ray and Mitchum 1996; Cummins et al. 2001). In our technique, harmonic analysis is used to solve for traveling solutions over a range of spatial locations, enabling separate resolutions of the northward and southward components. Multiple internal tide wave trains can then be resolved by plane-wave fits using data from all tracks in a fitting region, at the expense of the spatial resolution (Zhao and Alford 2009). Here, we choose the fitting region to be 250 km × 250 km on an overlapping grid of 0.1° × 0.1°. Internal tidal beams from the Aleutian Ridge and Hawaiian Ridge are successfully separated (section 5).

3. Mooring-observed internal tides

We here present the internal tides as observed in the mooring array using these techniques. We begin with the spatial pattern of the time-mean flux (section 3a), then present time series of velocity, displacement, energy, and flux at several sites north of the Hawaiian Ridge (section 3b). These measurements show the relative contributions of higher-mode and incoherent signals.

a. Time-mean flux

As seen in Fig. 1 (red arrows) and noted by Alford et al. (2007), observed semidiurnal energy flux is generally aligned along a beam toward ~60°, apparently emanating from FFS, as indicated by numerical simulation (yellow arrows). The flux generally decreases moving northeastward, ~3.5 kW m⁻¹ at MP1 and ~1.2 kW m⁻¹ at MP6 (Table 1), but not monotonically. An example is MP4, whose flux is about twice that of MP3, and more eastward than at other locations. As will be shown below (section 5), the general nonmonotonicity, and the specific magnitude and direction of the MP4 flux, are consistent with an interference pattern formed by the northeastward Hawaiian and southeastward Aleutian beams.

b. Progression through the profiler array

To gain a qualitative sense of the internal tide as a function of distance from the source, time series for MP1, MP3, MP4, and MP6 are presented (Figs. 4–7). The plot format will be presented for MP1, followed by discussion of the other sites.

1) Internal tide at MP1

The internal tide is clearly visible at MP1, only 218 km from its primary generation region near FFS, as three successive periods of more intense velocities (Figs. 4a and 4b, colors) and greater displacements (black lines), separated by more quiescent periods. These bandpassed semidiurnal signals are a combination of $M_2$, $S_2$, and incoherent constituents, as indicated earlier. A variety of vertical wavelengths can be discerned in the velocity (Fig. 4a). Nonetheless, modal decomposition indicates a dominance of mode 1 (Fig. 4c), as seen qualitatively in the top panel of Fig. 4 with deep flows measured at the Aanderaa site generally opposing those at shallower depths. In addition, velocity weakens appreciably approaching the mode-1 node near 1300 m (refer to Fig. 2c). Stronger mode-2 signals are seen in the displacement (Fig. 4d). However, the energy and flux are still dominantly mode 1 (Figs. 4g and 4h).

The consistency among the observed signals with a theoretical free-propagating wave northeastward along the array is examined next via two forms of hodographs, following Alford and Zhao (2007a). Such a wave would be expected to show current ellipses elongated in the along-beam direction (Fig. 4e, black), and along-beam velocity $v'$ in phase with $\eta$, with the slope indicated (Fig. 4f, black). Elongation of observed ellipses, which are computed in evenly spaced windows spanning the time series, varies with time but is not generally aligned with that expected for an along-beam free wave. However, the $v'\eta'$ ellipses are in excellent agreement with a northeastward
FIG. 4. Time series of the semidiurnal internal tide at MP1. (a) Semidiurnally bandpassed along-beam velocity (colors) and isopycnal displacement (black lines) from the ADCP and MP measurements. (b) As in (a), but for the Aanderaa and Sea-Bird measurements at ~3000 m depth. Black lines indicate isothermal displacement rather than isopycnals. (c) Along-beam velocity amplitude for modes 1 (purple) and 2 (green). (d) As in (c), but for vertical displacement. (e) The $u'v'$ hodographs of the mode-1 (purple) and -2 (green) components. The black ellipsis is for a theoretical free wave propagating northeastward along the beam. (f) As in (e), but for $v'\eta'$ hodographs. Normalization is such that a free wave propagating along the beam would fall along the black line. (g) Energy in modes 1–5 (stacked colors) and the coherent $M_2 + S_2$ constituent of the sum of the first five modes, (h) As in (g), but for flux magnitude. In (g) and (h), times of spring and neap barotropic tides at FFS (map refer to Fig. 1), predicted from the TPXO6.2 tidal model, are shown as vertical solid and dashed lines, respectively.
Taken together, these indicate interference between the signals generated at FFS to the east and Brooks Banks (BB) to the west (Fig. 1): the along-beam velocity would still be expected to be in phase with the displacement, but the two sources lead to distorted $u'v'$ ellipses relative to that for a single wave. That is, ellipses are not necessarily aligned with flux vectors in a multidirectional wave field (K. I. Martini et al. 2010, unpublished manuscript). Mode-2 hodographs are more variable (Fig. 4f, green), and even consistent with southward propagation toward the end of the record.

Fig. 5. As in Fig. 4, but for MP3.
The energy (Fig. 4g) and flux magnitude (Fig. 4h), plotted as stacked bands colored by mode, show clear spring–neap cycles dominated by mode 1 (even more so for flux than for energy). The spring–neap cycles in the energy and flux are synchronous, with spring-tide values exceeding neap-tide values by a factor of about 5. For MP1, the signals are strongly coherent and, hence, the semidiurnal flux and energy vary only slightly from the sum of the coherent portion of the first five modes (heavy black). The 0.7-day time lag of springs and neaps relative to the semidiurnal tidal forcing at FFS (24°N, 193.5°E) (Figs. 4g and 4h, gray and dashed lines, respectively) is consistent with
the group propagation expected for the mode-1 tide, as found by Alford and Zhao (2007a) and shown in more detail later (section 4b).

To summarize, the various signals at MP1 are as expected given its proximity to the forcing regions; namely, they indicate a largely coherent mode-1 tide propagating northeastward at the expected rate. Hodographs indicate interfering northeastward radiation from multiple sources at the Hawaiian Ridge. Southward signals are detected in mode 2, but are weak or undetectable in mode-1 flux.

FIG. 7. As in Fig. 4, but for MP6.
2) **ALONG-BEAM VARIATIONS**

Moving northeastward, moored measurements are displayed in the same format for MP3, MP4, and MP6 (Figs. 5–7).

In contrast to the strong visual dominance of low-mode signals at MP1, higher-mode signals grow in importance moving away from the source. Mode-1 and -2 amplitudes are comparable at each site for both velocity and displacement (panels c and d in Figs. 5–7). These lead to greater contributions in the higher-mode energy (panel g in Figs. 5–7). However, the flux (panel h in Figs. 5–7) continues to be dominated by mode 1 at all sites owing to the red vertical wavenumber spectrum for the baroclinic pressure, as found previously (Alford 2003; Nash et al. 2005).

These conclusions are seen in a more quantitative comparison of the modal content for the energy (Fig. 8a) and flux (Fig. 8b). The modal content for the energy changes from a strong mode-1 dominance at MP1 to near equipartition at MP6, consistent with a decaying mode-1 wave superimposed on a white vertical wavenumber background.

MP4 stands out for its strong mode-2 contribution. The lag of the spring-tide peaks for mode 2 is consistent with travel time from Hawaii, but the short record prevents drawing a stronger conclusion. Local generation near the Musicians Seamounts is also a possibility (refer to Fig. 1). Still, the greatest flux is mode 1 as seen in the time series.

The partitioning between HKE and APE (Fig. 8, plain and hatched boxes) also varies along the array. For a free-propagating internal tide, HKE always exceeds APE in a latitude-dependent ratio ranging from 1.5 at MP1 to 2.3 at MP6, following

\[
\frac{\text{HKE}}{\text{APE}} = \frac{\omega^2 + f^2}{\omega^2 - f^2},
\]

where \(\omega\) is the tidal frequency and \(f\) is the inertial frequency. The observed ratio at MP1 is \(\approx 2\), generally consistent with a free progressive wave. However, in contrast to these free-wave values, the observed APE exceeds HKE at MP3, MP5, and MP6, implying standing-wave patterns (Alford and Zhao 2007a; Zhao and Alford 2009). Additionally, the interference between multiple internal tidal beams is evidenced by the more eastward flux seen at MP2–MP4. As will be shown in section 5, these moorings are influenced by the southeastward Aleutian internal tide, as well as by the northeastward Hawaiian beam.

Coherence with astronomical tidal constituents as defined above generally decreases with increasing range (Fig. 9). (The increased coherence at MP4 is a consequence of its shorter record length.) In contrast to the visually apparent spring–neap cycles in the velocity and displacement at MP1, these are less obvious at the more...
northern sites partially owing both to amplitude and phase variations in the low-mode solutions (panels c and d in Figs. 5–7) and to greater contributions from less-coherent higher modes. Spring–neap cycles in energy and flux are still evident at MP3 and MP6 (the record at MP4 is too short for certainty), but the growing incoherence is evident as an increasing difference between the coherent (panels c and d in Figs. 5–7, black line) and total quantities (colors).

In summary, our observations indicate a rich time-variable multimodal internal tide field. Hodographs suggest that the internal tides radiated from multiple Hawaiian sources, rather than a single-point source at FFS, in addition to contributions from the Aleutian Ridge that will be examined in the following sections. Incoherence with astronomical tidal forcing can arise from variations in amplitude at generation (Mitchum and Chriswell 2000; Chiswell 2002) and refraction in eddy field (Rainville and Pinkel 2006b). Still, the bulk of the time-mean energy flux is carried by the coherent mode-1 internal tide, warranting its separate consideration next.

4. Comparisons with altimetric observations

We next conduct a direct comparison of the moored and altimetric SSHAs, taking advantage of the collocation of the IWAP mooring array and T/P track 249 (Fig. 1). We here focus on the mode-1 semidiurnal internal tide, because it carries the bulk of the tide energy, and because it is by far the most easily detectable in the altimetric measurements. The moored mode-1 SSHA is extracted by projecting profiles of mooring measurements onto discrete baroclinic modes as described in section 2e. The altimetric mode-1 SSHA is extracted by harmonic analysis and the new south–north separation technique (section 2h).

a. Moored and altimetric SSHA

For the most basic comparison, at each mooring site, the moored semidiurnal SSHA from the bandpass filter is plotted (Fig. 10, gray line) with the sum of the harmonic-fit altimetric $M_2$ and $S_2$ SSHA signals (black line). The altimetric records show the spring–neap beating of the

![Figure 9](image-url)
$M_2$ and $S_2$ constituents. In contrast, the temporally variable incoherent signal is also seen in the moored time series. In spite of this, a remarkable degree of agreement in magnitude and phasing is seen between the in situ and remotely sensed data as far north as MP5. Amplitude and phase differences finally develop at MP6 (~1600 km from FFS).

For a more direct comparison, the amplitude and phase of the harmonically fit signals are compared for $M_2$ and $S_2$ (Fig. 11). The moored harmonic fits are extracted in overlapping windows of 14.7 days (a spring–neap tidal cycle); then, the mean and standard deviation are plotted as dots and bars, respectively. As described in section 2h, altimetric signals are harmonically extracted for the northward (red) and southward components (green), and the total (black). For $M_2$, the total SSHA shows general (but nonmonotonic) northeastward decay, and half-wavelength modulation of amplitude indicative of a partial standing wave resulting from superposed signals from Hawaii and the Aleutians (Zhao and Alford 2009), as examined in detail below. The northward–southward-separated components (red and green) both show almost linearly increasing phase. For $S_2$, the differences are somewhat larger due in part to its weaker signals; however, interference between northward and southward signals is still discernible, as shown in the separated components (Figs. 11c and 11d, red and green).

Moored and altimetric $M_2$ signals are in good agreement in both amplitude and phase, in spite of the greatly different bandwidths of the two coherent estimates (~2 months versus ~10 yr). Some of the differences are explainable in terms of the different time scales affecting the two estimates. That is, the altimetric estimates are the signals that are time invariant over a 10-yr record. In contrast, processes such as mesoscale eddies and annual cycles lead to differences in both generation (Mitchum and Chriswell 2000) and propagation time (Rainville and Pinkel 2006a). For example, a greater moored amplitude at MP3 may be due to stronger generation during the
time of the IWAP measurements owing to variations in thermocline depth. Alternately, refraction could have shifted the antinode in the interference pattern closer to MP3 during this period. At MP6, the moored and altimetric phases differ by 70°, which is too large to be explained by the 43-km spatial offset (Table 1), which is parallel to wave crests (Fig. 1). We speculate that refractive phase shifts during the time of the IWAP deployment relative to the longer-term altimetric mean may be responsible.

b. Spatial propagation

The northeastward progression of the mode-1 semidiurnal internal tide at phase and group speeds is now examined. Theoretical propagation times are first computed for a linear free wave propagating toward the northeast by integrating Eqs. (4) and (5) along the mooring array starting from FFS (Fig. 1), following previous work (Rainville and Pinkel 2006b; Alford and Zhao 2007a). Travel time is plotted versus latitude with dashed lines in Figs. 12a–c, with the associated speeds given in Fig. 12d. Phase travel times for $M_2$ (Fig. 12a) and $S_2$ (Fig. 12b) are nearly indistinguishable from one another. However, the slight difference between their implied wavelengths leads to dispersion and group propagation (Fig. 12c).

For the altimetry, the phase propagation time is calculated by unwrapping phases of the northward $M_2$ and $S_2$ SSHAs, each of which increases northward along the track (Figs. 11b and 11d, red line). To focus on the signals emanating from Hawaii, only the northward component is used.

Our separate resolution of the $M_2$ and $S_2$ enables an altimetric estimate of group propagation (Fig. 13), which to our knowledge has not been attempted. Group propagation of the coherent signals arises from the phase difference between $M_2$ and $S_2$. Note that the phase propagation times of $M_2$ and $S_2$ have very slight differences (Figs. 12a and 12b, gray line). The $M_2$ and $S_2$ phases (Fig. 13a) diverge owing to their different wavelengths, as is better seen in the unwrapped quantity (Fig. 13b). Their phase difference and a linear fit are shown in gray and
black, respectively, in Fig. 13c. The spring and neap tides appear at the spatial locations where $M_2$ and $S_2$ are in phase ($\Delta \phi = 0^\circ$) and out of phase ($\Delta \phi = 180^\circ$), respectively. For the chosen reference time 0000 UTC 1 January 1992, the spring tide occurs at 40.4°N, and the neap tide at 26.3°N. The distance ($\approx 1740$ km) divided by the time ($\approx 7.4$ days, half of a spring–neap tidal cycle) gives a mean group velocity of $\approx 2.7$ m s$^{-1}$, which is in line with theoretical values (Fig. 12d).

For the moorings, the $M_2$ and $S_2$ phase propagation times are computed (Figs. 12a and 12b, dots) by integrating phases of the $M_2$ and $S_2$ SSHAs, respectively (Figs. 11b and 11d, purple dots). Because the moored phase estimates may have an integral number of tidal periods added to them, the altimetric phase (which agrees well with the moored phase; Figs. 11b and 11d) is used to resolve this uncertainty. A group propagation time (Fig. 12c, dots) is estimated from the time lag of the observed spring–neap maxima in the energy flux at each mooring (Figs. 4–7), relative to those in the barotropic forcing at FFS, following Alford and Zhao (2007a). At MP5, SSHA rather than flux is used, because the flux time series is too short to have spring–neap cycles (section 2e).

In summary, the phase and group estimates computed from moorings and altimetry agree well with each other, and with expectations from linear theory. The moored group speed estimate for MP2 is an exception, probably because of a noisy and less well pronounced spring–neap cycle (not shown).

5. Interference patterns

Several observations reported in the previous sections imply an interference pattern that results from multiple superposed waves: 1) eastward deflection of the moored time-mean energy flux at MP2–MP4 rather than at MP1 (Fig. 1); 2) variation of the HKE/APE ratio,
which varies in contrast to the constant value for a single progressive wave (Fig. 8); and 3) half-wavelength amplitude modulations in altimetric SSHA, and their separate resolution into north and south components (Fig. 11). Here, we employ the altimetry and a simple two-wave model to further examine the internal tide field in a two-dimensional view, instead of the above one-dimensional picture along the mooring array. We will show that the IWAP moorings lie in the overlapping region of north-eastward Hawaiian and southeastward Aleutian fluxes. Our focus is on the dominant $M_2$ internal tide, though $S_2$ also shows standing-wave features.

As described in section 2h, 2D snapshots of mode-1, $M_2$ SSHA are plotted for the northward component (Fig. 14a), the southward component (Fig. 14b), and their total (Fig. 14c). Regions where no clear north–south separation is possible are left white. As discussed in the work of Zhao and Alford (2009), the north-eastward Hawaiian beam (Fig. 14a) originates from FFS and Nihoa Island (NI), two strong generation sites at the Hawaiian Ridge (Merrifield and Holloway 2002; Lee et al. 2006). The low spatial resolution (250 km × 250 km) of the altimetric estimates precludes the separation of the individual FFS and NI contributions. [Zhao and Alford (2009) show that the FFS and NI beams cross near 32°N.] Southward signals (Fig. 14b) can be tracked over 2000 km back to Amukta Pass (about 100 km in width, centered at 52.5°N, 188°E) in the Aleutian Islands. The combined signal (Fig. 14c) shows a phase discontinuity near 198°E, and a maximum aligned in the north–south direction to the west of that, between the dashed lines.

To examine the interference pattern, two plane waves are next constructed (Figs. 14d and 14e) by matching the observed amplitude, wavelength, phase, and propagation direction (values in Table 3) of the northbound and southbound signals. Their superposition (Fig. 14f) reproduces both the phase discontinuity and the “apparent beam” seen in the observed combined signals (Fig. 14c). This feature is similar to that to the south of the Hawaiian Ridge, which is caused by multiple sources at the ridge (Rainville et al. 2010). The parameters used are chosen...
to match the altimetric observations (Table 3). We will next examine the amplitude, phase, energy, and energy flux in a subregion (Fig. 14f, box) and compare these results with the moored observations.

The small-scale patterns in the amplitude, phase, energy, and energy flux of the interference pattern are next examined in the subregion 27°–32°N, 194°–199°E, and compared with the moored observations at MP2–MP4. For each wave, each of these quantities is simply related to SSHA (Chiswell 2006; Zhao and Alford 2009). The corresponding expressions for the superposed wave fields are more complicated owing to the cross terms in the velocity and pressure (Nash et al. 2004; Martini et al. 2007; Rainville et al. 2010). Here, a constant inertial frequency $f$ at 28.9°N is taken, neglecting its small latitudinal change. The results for the amplitude, phase, energy, and flux are shown in Fig. 15: interference patterns are presented in Figs. 15a–d, and comparisons with moored observations are given in Figs. 15e–h.

The amplitude and phase of the superposed signal shown in the snapshot in Fig. 14f are plotted in Figs. 15a and 15b. The amplitude displays a banded half-wavelength periodic structure, while the phase shows a twisted pattern. Plotted along the mooring array versus latitude, half-wavelength

![Figure 14](https://example.com/figure14.png)

**Fig. 14.** Snapshots of SSHA from (top) altimetry and (bottom) the superposition of two plane waves. Altimetric $M_2$ SSHAs for (a) the northward constituents, (b) the southward constituents, and (c) their total. FFS and NI are labeled. Green dots and line indicate the IWAP moorings and T/P track 249, respectively. Snapshots of SSHA for a mode-1 wave traveling toward (d) the northeast, (e) the southeast, and (f) their sum. Amplitude, wavelength, direction, and phase values are given in Table 3. In (c),(f), dashed lines indicate an “apparent beam” resulting from their constructive interference.

![Table 3](https://example.com/table3.png)

**Table 3.** Parameters of plane waves that are used to simulate the Hawaiian and Aleutian beams.

<table>
<thead>
<tr>
<th>Plane wave</th>
<th>Wavelength (km)</th>
<th>Amplitude (mm)</th>
<th>Direction (°)</th>
<th>Phase (°)</th>
</tr>
</thead>
<tbody>
<tr>
<td>Hawaiian beam</td>
<td>162</td>
<td>12</td>
<td>68</td>
<td>39</td>
</tr>
<tr>
<td>Aleutian beam</td>
<td>162</td>
<td>6</td>
<td>-77</td>
<td>297</td>
</tr>
</tbody>
</table>

modulations are present in both quantities (Figs. 15e and 15f, black). Both are reminiscent of modulations observed in the along-track altimetric observations (Figs. 11a and 11b). For comparison, the individual northward and southward waves (purple) show constant amplitude and linearly increasing phase. The moored observations (green dots) agree very well with the two-wave simulated results, with the exception of the amplitude at MP3.

The simulated energy (Fig. 15c), HKE/APE ratio (Fig. 15g), and flux (Figs. 15d and 15h) all show banded interference patterns, demonstrating that their values are spatially variable rather than constant values for plane waves. HKE/APE ranges from 0.25 to 13 in this case (Fig. 15g, black), compared with a constant ratio ($\approx 1.69$ for $M_2$ at 28.9$^\circ$) for free progressive plane waves (purple). Generally, the observed HKE/APE ratios at MP2–MP4 (Fig. 15g, green) agree with the two-wave simulation. Importantly, the superposed flux vector shows alternate bands of flux in the transverse direction that do not point away from either source (Nash et al. 2004; Alford and Zhao 2007b; Martini et al. 2007). The observed veering and modulation of the moored flux vectors (Fig. 15h, green arrows) are generally consistent with the simulated fluxes (red arrows).

6. Summary

Observations of the semidiurnal internal tide from six highly resolved profiling moorings spanning 1400 km have been presented, with the aim of characterizing the internal tide’s synoptic spatial structure and progression over a long distance. The collocation of the mooring array and one T/P track has enabled a direct comparison of the moored and altimetric SSHAs. We have thus obtained a qualitative picture of the semidiurnal internal tide from the mooring array that is well resolved in time and depth, and from the satellite altimeter that is well resolved in space along one track. The two complementary methods reveal complicated interference patterns caused by two internal tidal beams. Major findings include the following:
7. Discussion

A major motivation for the IWAP experiment was to determine the factors governing the attenuation of internal tides as they propagate away from their sources. A key long-term goal is determining the distribution of the dissipation of the internal tide. Toward this end, this study represents some progress, but also highlights the difficulty of the problem. Progress stems from the demonstration of the utility of combined moored-altimetric studies in diagnosing some of the aspects of the internal tide radiation field that would not be possible with either technique alone. The remarkable agreement between detection techniques as disparate as (a) an array of moorings deployed for less than 2 months, (b) a synoptic shipboard snapshot (Alford and Zhao 2007a), and (c) a harmonic fit from multiple years of remotely sense data offers hope that the spatial context offered by the altimetry will be of great value in making future progress.

However, the complexity of the interference patterns, which are present even close to a known strong conversion region such as Hawaii, should give pause in simple interpretations. For example, in an ocean with a single point source, the internal tide’s attenuation can be measured as the range derivative of flux, accounting for cylindrical spreading. In the presence of multiple sources, the resulting interference pattern has a nonmonotonic decay of flux even along “apparent beams” (Rainville et al. 2010). In certain regions, such as this one, where two sources dominate, sense can be made of the patterns. In other more complicated regions extreme care should be used in interpreting point measurements.

Additional complications arise because of the limitations of the altimetric technique (Zhao and Alford 2009). As discussed, the primary issues are the coarse spatial resolution, and the inability to resolve the temporally incoherent portion of the signals, which are expected to grow with increasing range from the source owing to mesoscale refraction by a time-variant mesoscale field. While this study makes it clear that direct comparisons can be made, a quantitative transfer function between the altimeter-observed internal tides and the in situ, time-variant internal tides is sorely needed. Only then can observed attenuation be associated with “true” dissipation and radial spreading. In the interim, the processes and degree of internal tide dissipation remain open questions, and altimetric signals remain a valuable tool for interpretation.

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APPENDIX

Errors

Errors arise from moorings’ imperfect vertical resolutions, the adequacy of which depends on the mooring geometry and the modal partitioning of the internal tide signals (Nash et al. 2005). Rainville and Pinkel (2006a) successfully determine mode-1 internal tide solutions from well-resolved measurements in the upper 800-m
water column alone. In IWAP, higher-mode signals render upper-column-only solutions extremely unreliable and even of the incorrect sign, as determined by comparison of the energy flux computed from full-depth lowered ADCP–CTD time series (Fig. 1, blue dots) with the moored quantities (not shown). Following Nash et al. (2005), the magnitude of the effect is assessed with Monte Carlo simulations by running a synthetically

FIG. A1. Mean (circles) and standard error (bars) vs mode number (left to right) 1–5 and weighting of the 3000-m data for (top) displacement, (middle) velocity, and (bottom) flux, simulated using Monte Carlo methods following Nash et al. (2005). A synthetic internal tide is created by superposing 30 modes with a red spectrum in the vertical wavenumber \( m - m^{-1} \). The horizontal lines represent the “true” values in the synthetic wave field.

FIG. A2. Comparisons of barotropic tidal current ellipses from mooring measurements (black) and the TPXO6.2 tidal model (gray) for (top) \( M_2 \) and (bottom) \( S_2 \); (left to right) MP1 to MP6. The lines indicate current vectors at 0000 UTC 1 Jan 2006. MP5 is excluded due to its short profiling time. Note the scale difference between the top and bottom panels.
generated wave field past our mooring geometry (Fig. A1). As found by Rainville and Pinkel (2006a), mode-1 displacement is reliably estimated without deep data (see top-left panel of Fig. A1; weight = 0). However, our measurements at 3000 m are critical to constraining the fits for displacement modes ≥2 (top-right panels in Fig. A1), and all velocity modes (middle row of panels in Fig. A1). Here, a weight is added to the 3000-m data to increase their contributions in modal decomposition. Mode-1 fluxes (bottom row of panels Fig. A1) are hence unreliable (showing high bias and large error bars) when the deep data are excluded from calculation (weight = 0) but are otherwise insensitive to weighting. Corresponding fractional standard errors with a weight of 10 (taken in this paper) for each uncorrelated estimate are 8%, 30%, and 50% for the first three modes.

Building on the synthetic calculations, one means of assessing the reliability of the modal projection is to compare our estimates of barotropic tidal velocity [Eq. (6), mode = 0] with predictions from the TPXO6.2 tidal model (Egbert and Erofeeva 2002). The mooring-obtained $M_2$ and $S_2$ currents are ≈2 and ≈1 cm s$^{-1}$ (Fig. A2, black). Agreement with TPXO6.2 (gray) is better for the $M_2$ constituent, presumably owing to the stronger signals. For $M_2$, the amplitude, elliptical elongation, and phase are in good agreement with the model predictions. For $S_2$, the moored amplitudes appear to be biased slightly high, and the phases fall slightly behind, at all locations (except for MP6 in amplitude). We interpret the general good agreement as evidence that the modal fits are partitioning energy reasonably into the appropriate modes.

REFERENCES


