

Skills and Limitations of the Adiabatic Omega Equation: How Effective Is It to Retrieve Oceanic Vertical Circulation at Mesoscale and Submesoscale?

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ABSTRACT: The quasigeostrophic and the generalized omega equations are the most widely used methods to reconstruct vertical velocity w from in situ data. As observational networks with much higher spatial and temporal resolutions are being designed, the question arises of identifying the approximations and scales at which an accurate estimation of w through the omega equation can be achieved and what critical scales and observables are needed. In this paper we test different adiabatic omega reconstructions of w over several regions representative of main oceanic regimes of the global ocean in a fully eddy-resolving numerical simulation with a $1/60^\circ$ horizontal resolution. We find that the best reconstructions are observed in conditions characterized by energetic turbulence and/or weak stratification where near-surface frontal processes are felt deep into the ocean interior. The quasigeostrophic omega equation gives satisfactory results for scales larger than ~ 10 km horizontally while the improvements using a generalized formulation are substantial only in conditions where frontal turbulent processes are important (providing improvements with satisfactory reconstruction skill down to ~ 5 km in scale). The main sources of uncertainties that could be identified are related to processes responsible for ocean thermal wind imbalance (TWI), which is particularly difficult to account for (especially in observation-based studies) and to the deep flow that is generally improperly accounted for in omega reconstructions through the bottom boundary condition. Nevertheless, the reconstruction of mesoscale vertical velocities may be sufficient to estimate vertical fluxes of oceanic properties in many cases of practical interest.

KEYWORDS: Ageostrophic circulations; Fluxes; Mesoscale processes; Vertical motion

1. Introduction

In geophysical fluids, the combined effect of stratification and rotation strongly inhibits vertical velocities over a broad range of horizontal scales L_h , near and above the so-called deformation radius (R_d), typically 30 km in the ocean. Vertical velocities remain small even for L_h much below R_d (Pollard and Regier 1992; Giordani et al. 2006), typically a few meters to several tens of meters per day for submesoscale motions, $L_h \sim O(1)$ km, although their intensity tends to increase somewhat at finer scales. Much weaker than horizontal advection, vertical transport of oceanic properties such as heat and biogeochemical tracers is nevertheless of crucial importance for the overall functioning of the World Ocean (Lévy et al. 2012a). In some regions, atmospheric forcings are such that the large-scale flow (100 km and larger) has a vertical component, e.g., under the influence of coastal upwelling favorable winds or positive wind stress curl. At finer scale (1–100 km) intermittent vertical

velocities are associated with the mesoscale and submesoscale turbulence and can be generated by forced and unforced motions such as frontogenesis, baroclinic instabilities or air–sea interactions, with possible coupling between them (e.g., Thomas and Lee 2005). They are responsible for vertical fluxes that have proved difficult to quantify but are widely believed to play a major role in the heat (Ferrari 2011; Su et al. 2020; Siegelman et al. 2020) and salt budgets (Lien et al. 2014), in the carbon and nutrient cycles (Ledwell et al. 2008; Balwada et al. 2018), and in shaping oceanic biodiversity (Lévy et al. 2010, 2012b, 2014, 2018; Siegelman et al. 2020). It should be noted that vertical velocities also vary at the scales smaller than the submesoscale in particular vertical fluxes associated with 3D turbulence (Whitt et al. 2019). However, the submesoscale–mesoscale (hereafter SMS) range on which this study focuses is important because (i) it presumably contains a large fraction of covariance between vertical velocity w and many key tracer fields (Ledwell et al. 2008; Chenillat et al. 2015; Balwada et al. 2018) and (ii) SMS vertical motions can be coherent over relatively long/large time/space scales (including vertically) such that they produce long range vertical displacements and fluxes, e.g., organic and inorganic carbon from the euphotic layer into the dark ocean (Boyd et al. 2019).

Direct measurements of vertical velocities with SMS spatiotemporal resolution would thus be highly desirable. However, the magnitude of w at these scales is typically $O(1)$ mm s⁻¹ (~ 100 m day⁻¹) or less (e.g., up to 2–3 mm s⁻¹ in the highly turbulent Gulf Stream, Lindstrom and Watts 1994)

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which places them below the noise level of any existing current meter for the typical time scales of a few days to weeks over which they vary. Different ingenious ways to circumvent this difficulty have been developed over time. For instance, direct integrated measurements of vertical displacements $\int w dt$ have been made using Lagrangian drifters (Bower and Rossby 1989; D'Asaro et al. 2004; Steffen and D'Asaro 2002). Alternatively, indirect reconstruction methods have been proposed based on the heat/density conservation equation (Strass 1994; Lindstrom et al. 1997; Yu et al. 2019), the vorticity equation (Strass 1994; Giordani et al. 2005) or the 3D nondivergence of the flow (Helber and Weisberg 2001; Horii et al. 2011). A different framework has also emerged to infer vertical velocities from the theory of surface quasigeostrophy (SQG; Held et al. 1995; Lapeyre and Klein 2006a; LaCasce and Mahadevan 2006), where the 3D flow structure can be essentially determined from the knowledge of the surface buoyancy field (Isern-Fontanet et al. 2006; Klein et al. 2009; Ponte and Klein 2013; Qiu et al. 2020). Overall, the most commonly used method to infer w is based on frontogenetic theories and the so-called omega equation, which is the subject of the present study.

Quasigeostrophic and semigeostrophic versions of the omega equation (Hoskins et al. 1978) have been applied for decades (Leach 1987; Pollard and Regier 1992). The former and to a lesser extent the latter are by nature suited to low Rossby number environments, i.e., a priori away from regions of intense vertical velocities. A generalized version of the omega equation was first introduced in the atmospheric community (Davies-Jones 1991; Pauley and Nieman 1992; Giordani and Caniaux 2001) and subsequently applied to the ocean (Viúdez et al. 2002; Giordani et al. 2006). Vertical velocity forcing processes present in this generalized omega equation are kinetic deformation that arises in shear and confluence situations, mixing and momentum diffusion which can also disrupt the thermal wind balance and an additional prognostic term due to the rate of change of unbalanced motions. This latter term shall be zero under the quasigeostrophic approximation and is systematically neglected in ocean applications, which makes the omega equation diagnostic for w [see Qiu et al. (2020) for the only evaluation of this term that we know of]. Despite this simplification, the generalized formulations of the omega equation are expected to hold even in regions where the flow exhibits high Rossby number (Viúdez et al. 2002; Viúdez and Dritschel 2004; Shearman et al. 2000). Since the omega equation is most frequently used to infer w from observations the chosen formulation usually depends on the data available. The adiabatic quasigeostrophic version of the equation only requires density observations and a reference level to derive geostrophic currents from the thermal wind balance. Knowledge of the absolute horizontal velocity field allows to take into account the effect of the ageostrophic deformation and advection. Solving more elaborate forms of the omega equation require additional data, such as atmospheric forcing (Giordani et al. 2006). Statistical methods based on multivariate empirical orthogonal functions can be used to determine the SMS ocean state from surface satellite information and sparse in situ data (e.g., ARMOR3D high-resolution

operational product; Guinehut et al. 2012). Those methods have been used regularly to infer the forcing terms for omega inversions (Buongiorno Nardelli and Santoleri 2005; Buongiorno Nardelli et al. 2012, 2018; Barceló-Llull et al. 2018; Buongiorno Nardelli 2020).

Numerous investigations of oceanic vertical velocities based on the omega equation have been carried out since the 1980s, most frequently in the context of mesoscale-resolving observational efforts. A few of them have used independent ways to estimate w and evaluate the skills of the omega reconstructions. Conclusions are generally that the omega equation has reasonable skills at the mesoscale. A few studies have also applied this equation to numerical model outputs to test its reliability and degree of accuracy (Pinot et al. 1996; Allen et al. 2001; Rixen et al. 2003; Uchida et al. 2019). The domain of validity and typical errors attached to an omega reconstruction for realistic cases are, however, not clearly established. This is particularly true at submesoscale, where the omega equation has been increasingly applied (Pallàs-Sanz et al. 2010). The benefit that one can expect from using more elaborate versions of the omega equation compared to the quasigeostrophic formulation are also not clearly established. An overview of the literature on the subject offers a great diversity of conclusions (from major to no benefits or even degradation of the reconstruction). We provide a review of omega-reconstructions literature in the online supplemental material. Overall, this literature appears to be mainly composed of an accumulation of test cases (Pascual et al. 2017; Buongiorno Nardelli et al. 2012; Rixen et al. 2003; Allen et al. 2001), many of which are in idealized settings (Viúdez and Dritschel 2004). These individual cases can be difficult to compare against one another.

The main novelty of the present study is to apply the same analysis framework to assess the skills of the adiabatic omega equation in several dynamical regimes representative of a broad diversity of ocean conditions. The aim is to propose a more integrated and comprehensive understanding of the skills and limitations of omega reconstructions. Given this ambition, we restrict the scope somewhat by mainly focusing on vertical velocities outside mixed layers, typically at depths ~ 200 – 400 m. Submesoscale velocities in the mixed layer are important too (e.g., Smith et al. 2016) but they more strongly compete and interact with 3D turbulent processes responsible with intense mixing in ways that are just beginning to be clarified (Thomas and Lee 2005; Hamlington et al. 2014; Suzuki et al. 2016; Sullivan and McWilliams 2018; Callies and Ferrari 2018). On the other hand, we expect SMS vertical velocities obtained by means of omega reconstruction to be more readily useful to estimate tracer fluxes below the mixed layer, typically upward fluxes of nutrient into the euphotic layer (Pascual et al. 2015) or downward fluxes of oxygen into subsurface hypoxic layers (Thomsen et al. 2016).

Exploring the ability of the adiabatic omega equation to reconstruct w in our realistic, eddy-resolving, circulation model reveals that, in this depth range, reconstruction skills turn out to be strongly sensitive to the dynamical regime under consideration, but also to the physical (length) scale of interest. In particular, it is shown that omega reconstructions perform well for scales down to ~ 10 km while rapid degradation occurs at

smaller scales. Consequently, the best results are obtained in conditions characterized by high vertical velocity variance which are found in regions with intense mesoscale frontogenesis. On the other hand, relaxing the QG assumptions with the generalized formulation is shown to have a modest impact on the reconstruction. Instead, an analysis of the relative importance of the sources of errors reveals that a major impact can be attributed to the choice of the bottom boundary conditions (BBC). The impact of neglecting vertical mixing terms, though presently deemed limited to the upper layers, would also require dedicated investigations.

The paper is organized as follows. Data and methods are presented in section 2. The vertical velocity regimes are described in section 3. In section 4, vertical velocity reconstructions in the different dynamical regimes retained for this study are evaluated. A series of sensitivities allows us to test the impact of the level of complexity of the omega equation, the formulation of the discrete problem, and several possible choices related to the boundary conditions. Some reflections on the sources of errors and the dynamical environment are offered in the discussion.

2. Data and methods

Mathematical symbols have their usual meaning: T , S , ρ , and σ_t refer to potential temperature, salinity, density, and potential density anomaly, respectively. Parameters x (resp. y and z) and u (resp. v and w) refer to zonal (resp. meridional and vertical) directions and velocity. More precisely, in the case of vertical velocities we will distinguish true velocities w from reconstructed velocities denoted ω with a subscript that refers to the precise omega equation formulation that is being used (see section 2b). Parameter N^2 is the buoyancy frequency.

a. The NATL model

The outputs from a submesoscale-permitting numerical simulation are used to reconstruct vertical velocity fields using different versions of the omega equation and compare them to the ground truth model vertical velocities w_{model} . The NATL60 MJM155 simulation is run using the NEMO v3.5 code. It has a horizontal resolution of $1/60^\circ$ (dx comprised between 0.8 and 1.6 km depending on the latitude), yielding an effective resolution of $\sim 10dx$ (~ 10 – 15 km) (Soufflet et al. 2016). This means that a turbulent feature associated with a typical wavelength around ~ 10 – 15 km (typical scale ~ 1.5 – 2 km) are energetically subjected to a negligible influence of the dissipation operator. Below those scales numerical errors can arise that will affect the representation of physical processes although the misrepresented dynamics in the model solutions may still share important properties with the dynamics of the system obtained after resolution convergence (Le Sommer et al. 2018). On the vertical, the simulation has 300 levels (dz increases with depth and ranges from 1 to 30 m) and in the range of depths of analysis (200–400-m depth), the vertical resolution is ~ 10 m. The simulation is forced by realistic atmospheric forcings (DFS5.2) and boundary conditions (GLORYS2V3). It was integrated over a 5-yr period (2004–08) (Amores et al. 2018; Ducouso et al. 2017; Fresnay et al. 2018).

The model domain encompasses the whole of North Atlantic. For this study, four contrasted sectors were selected. For each region, 11 consecutive daily averaged outputs are analyzed for two opposite seasons of the year, in June and December. Near-inertial motions produced by the high-frequency variability of the atmospheric forcings are partly filtered out and partly aliased in model u , v , σ_t , and w daily fields that are used in this study. A cleaner separation between balanced and unbalanced motions would be useful to untangle their respective roles (e.g., see Qiu et al. 2020). Lacking the high-frequency model outputs needed to do better, we note that, in most situations, omega estimations made using observations contain an unknown but presumably larger inertia–gravity (IG) wave contribution. Our study can thus be considered as a favorable evaluation in which this source of errors is minimized by the daily averaging. Absence of tidal forcing in NATL60 will also tend to underestimate the energy level of unbalanced motion (Qiu et al. 2018) and thus has a similar consequence. The relevance of the study to the real ocean context is, however, justified by the fact that model vertical velocities in the submesoscale-permitting class of simulations are known to capture the patterns expected from the theory and observation based studies.

The four regions we focus on are as follows:

- The Gulf Stream/LatMix region (LMX; Shcherbina et al. 2013) centered around 38°N and 67°W . It encompasses the very energetic Gulf Stream Current. It is characterized by intense mesoscale activity composed of meanders and eddies. Vertical motions in the Gulf Stream region have been the subject of many studies aimed at describing water exchanges between the jet core and its vicinities (Bower and Rossby 1989; Lindstrom and Watts 1994; Joyce et al. 2013). This environment and the large-amplitude meanders of the Gulf Stream in particular are well known to produce intense vertical velocities (Fig. 1a).
- The region of the Azores Current (AZO) centered at 33°N , 34°W , which is characteristic of a subtropical regime with limited atmospheric forcing and modest mesoscale activity. It is the least energetic of the four regions in terms of vertical velocities (Fig. 1b). We relate this to the modest role played by surface density contrasts and upper-ocean frontogenesis in the regional dynamics (Lapeyre 2009; Volkov and Fu 2010, 2011).
- An area of the subpolar gyre southwest of Iceland encompassing part of the Reykjanes Ridge (REK) and centered around 54°N , 31°W . The regional dynamics is characterized by weak mean flow and a moderately intense mesoscale activity made of deep-reaching isolated structures (Fig. 1c).
- A sector of the Atlantic northeastern margin near 49°N and 15°W , where the Osmosis experiment took place (OSM; Buckingham et al. 2016). The regional dynamics is characterized by weak large-scale circulation, weak mesoscale activity, and a marked seasonality of the submesoscale activity which is very intense in the wintertime (Fig. 1d; Thompson et al. 2016).

To limit the computational cost of inverting the omega equation, a 3D subdomain of each region is retained with dimensions ~ 360 km \times ~ 270 km on the horizontal and 1600 m in

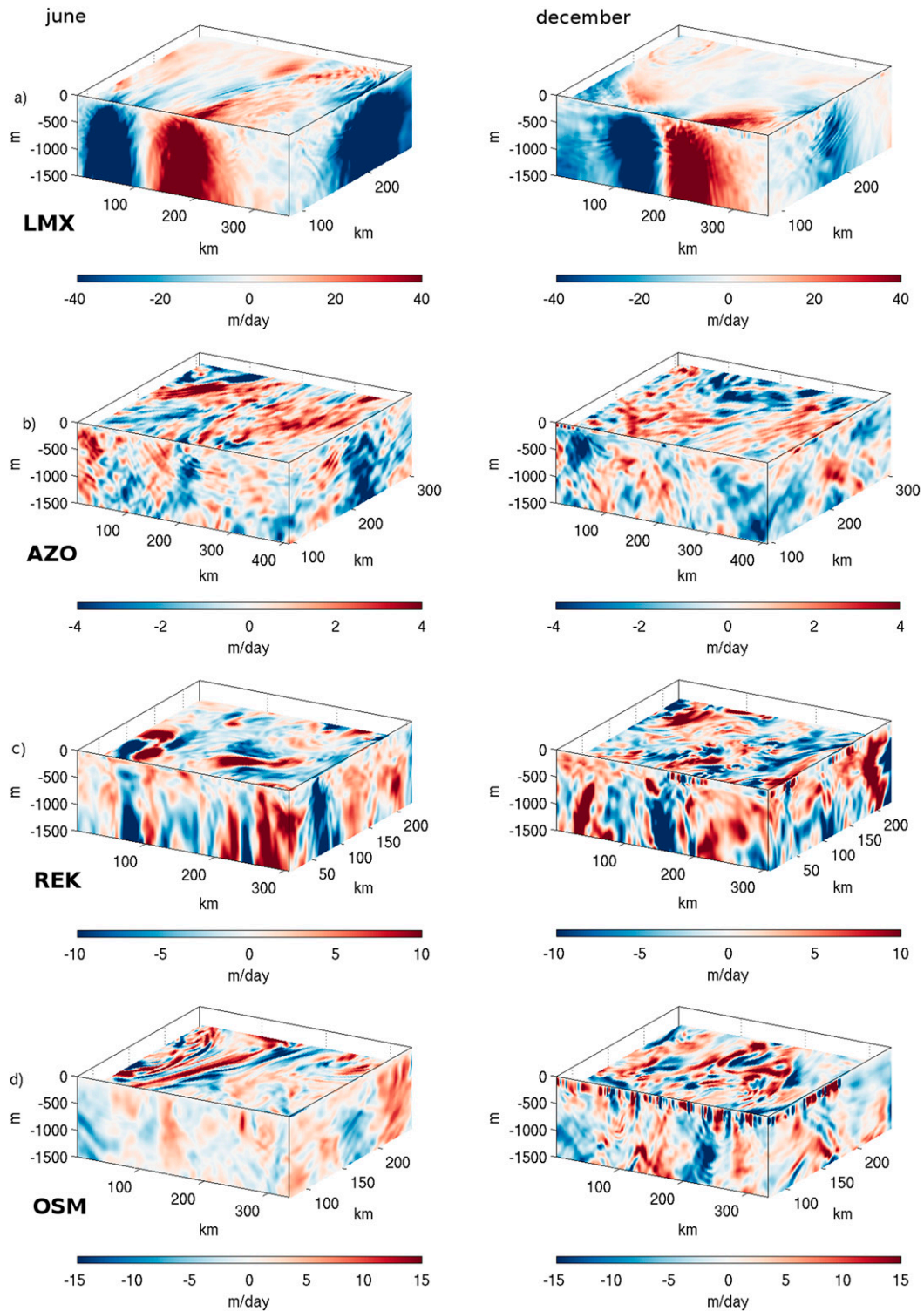


FIG. 1. Model vertical velocity on (left) 10 Jun 2008 and (right) 10 Dec 2008 in the (a) LMX region, (b) AZO region, (c) REK region, and (d) OSM region. The upper horizontal plan corresponds to the depth of analysis ($z_a = 220$ m in LMX, 250 m in AZO, and 380 m in REK and OSM).

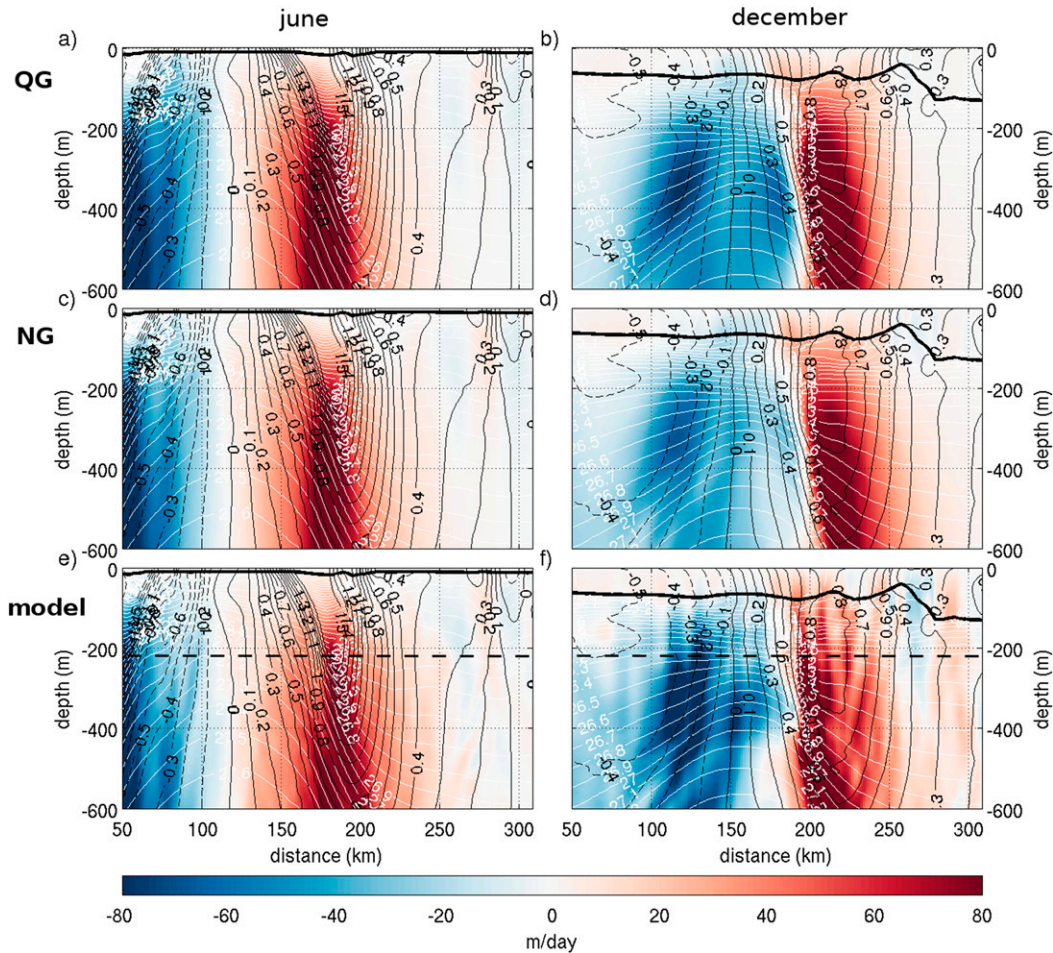


FIG. 2. Vertical velocity in the LMX region on (left) 10 Jun 2008 and (right) 10 Dec 2008 for the (a),(b) QG inversion; (c),(d) NG inversion; and (e),(f) model data. The thin white (resp. black) lines represent the isopycnics (resp. horizontal velocity; m s^{-1}). The bold solid (resp. dashed) black line indicates mixed layer base (resp. the analysis depth z_a).

the vertical (reduced depth ranges will also be used in sensitivity tests in section 4d). Horizontal (native) resolution is ~ 1.5 km while the data are linearly interpolated on a regular vertical grid with a 5-m resolution. For each region, a particular depth is selected to perform some of our analyses. As presented in the introduction, the spirit of this investigation led us to choose depth levels 50 m below the deepest winter mixed layers encountered in that region, so near-surface frontal processes are attenuated at this depth and our study is useful to make progress on biogeochemical fluxes between the mixed layer and the ocean interior. Specifically, selected depths of analysis (referred to as z_a) are 220 m for LMX, 250 m for AZO, and 380 m for REK and OSM.

b. Formulations of the omega equation

The underlying principles of the omega equation combine two lines of argument (Hoskins 1982):

- Kinematical: a turbulent flow stirring a heterogeneous surface buoyancy field produces regions of gradient intensifica-

tion. Specifically, frontal intensification is promoted in the confluence and shear situations that are frequently encountered in mesoscale turbulent conditions (e.g., Fig. 2).

- Dynamical: the thermal wind balance, which should approximately hold at meso/submesoscale, is disrupted in situations of frontal intensification where density gradients are enhanced while the vertical velocity shear is being reduced. This leads to the development of ageostrophic secondary circulations (ASC) that attempt to thwart thermal imbalance by restoring shear and slumping isopycnics. In the ocean interior this process is very efficient at preventing frontal intensification, which is a further justification for the weakness of vertical velocities away from the surface. At the air-sea interface, the upper boundary condition $w \approx 0$ limits the efficiency of ageostrophic circulations so frontal intensification can generally proceed further and yield intense vertical velocities.

Mathematically, the frontogenesis process is captured by a relationship between the spatial derivatives of w and a

forcing term expressed as the divergence of a vector of forcings \mathbf{Q} :

$$\mathcal{L}(w) = \nabla \cdot \mathbf{Q}, \tag{1}$$

where \mathcal{L} is a second-order differential operator that can also take different forms depending on which processes responsible for thermal wind disruptions are being considered in \mathbf{Q} .

A starting point to derive the omega equation for the general case is the set of equations governing hydrostatic and Boussinesq flows:

$$\frac{du}{dt} - fv = -\frac{1}{\rho_0} \frac{\partial p}{\partial x} + F_x, \tag{2}$$

$$\frac{dv}{dt} + fu = -\frac{1}{\rho_0} \frac{\partial p}{\partial y} + F_y, \tag{3}$$

$$\frac{dp}{dt} = F_p, \tag{4}$$

$$\nabla \cdot \mathbf{v} = 0, \tag{5}$$

where (F_x, F_y) and F_p are source/sink of momentum and buoyancy caused by turbulent mixing.

The flow is then decomposed into a geostrophic (\mathbf{v}_g) and ageostrophic (\mathbf{v}_{ag}) component,

$$\mathbf{v} = \mathbf{v}_g + \mathbf{v}_{ag}, \tag{6}$$

where the geostrophic velocity satisfies the thermal wind balance:

$$\begin{cases} f \frac{\partial u_g}{\partial z} = \frac{g}{\rho} \frac{\partial \rho}{\partial y} \\ f \frac{\partial v_g}{\partial z} = -\frac{g}{\rho} \frac{\partial \rho}{\partial x} \end{cases}, \tag{7}$$

and the residual ageostrophic flow component departing from this balance [the so-called thermal wind imbalance (TWI)] satisfies

$$\begin{cases} f \frac{\partial u_{ag}}{\partial z} = f \frac{\partial u}{\partial z} - \frac{g}{\rho} \frac{\partial \rho}{\partial y} \\ f \frac{\partial v_{ag}}{\partial z} = f \frac{\partial v}{\partial z} + \frac{g}{\rho} \frac{\partial \rho}{\partial x} \end{cases}. \tag{8}$$

The generalized omega equation (1) is obtained by manipulating the time evolution equation for the TWI and yields (Giordani and Planton 2000)

$$f^2 \frac{\partial^2 w}{\partial z^2} + \nabla_h (N^2 \cdot \nabla_h w) = \nabla \cdot \mathbf{Q}, \tag{9}$$

where the \mathbf{Q} vector involved in the right-hand side (rhs) of the equation can be expressed as a sum of different forcings:

$$\mathbf{Q} = 2 \underbrace{(\mathbf{Q}_{tw_g} + \mathbf{Q}_{tw_{ag}})}_{\mathbf{Q}_{tw}} + \mathbf{Q}_{dag} + \mathbf{Q}_{dr} + \mathbf{Q}_{th} + \mathbf{Q}_{dm} \tag{10}$$

following the notations of Giordani et al. (2006). Parameter \mathbf{Q}_{tw} is the kinematic deformation and can be decomposed

into a geostrophic \mathbf{Q}_{tw_g} and an ageostrophic component $\mathbf{Q}_{tw_{ag}}$. Parameter \mathbf{Q}_{dag} is the deformation of the thermal wind imbalance and \mathbf{Q}_{dr} its material rate of change. In all practical situations, \mathbf{Q}_{dr} cannot be estimated and is therefore unaccounted for in the remainder of the study. Parameters \mathbf{Q}_{th} and \mathbf{Q}_{dm} refer to the diabatic turbulent buoyancy and momentum forcings, respectively.

Being interested in the ability to determine vertical velocities through and below the thermocline we neglect the effects of diffusive momentum and buoyancy fluxes which are mainly active in or immediately below the mixed layer (Giordani et al. 2006; Yoshikawa et al. 2012; Thomas et al. 2010) but have a limited effect on subsurface velocities (Nagai et al. 2006; Yoshikawa et al. 2012). In fact, Xie et al. (2017), using microstructure shear measurements to infer the vertical diffusivity, derived the vertical mixing terms (\mathbf{Q}_{th} and \mathbf{Q}_{dm}) and showed that below the thermocline the vertical velocity associated to those terms is one to two orders of magnitude smaller than the one associated to the deformation of the flow. An exception is the study of (Qiu et al. 2020) in which vertical velocities arising from diffusive terms remain of magnitude comparable to those produced by ageostrophic deformation ($\mathbf{Q}_{tw_{ag}}$) well below the mixed layer. We suspect that this is because vertical velocities due to mixing are not calculated explicitly in Qiu et al. (2020) but instead are obtained as a residual and thus also contain contributions from various sources such as imperfect boundary conditions and other numerical errors (see sections 4c and 4d).

We will limit ourselves to comparing the two ω reconstructions that are most commonly used with real ocean data: the QG version, where ω_{QG} is solely forced by the curvature of the geostrophic flow ($\mathbf{Q}_{QG} = 2\mathbf{Q}_{tw_g}$), as in Hoskins et al. (1978), and a more complete formulation (NG) in which the \mathbf{Q} vector forcing ω_{NG} includes some contribution from the ageostrophic flow ($\mathbf{Q}_{NG} = 2\mathbf{Q}_{tw} + \mathbf{Q}_{dag}$).

The \mathbf{Q}_{tw_g} vector in the QG formulation has the form

$$\mathbf{Q}_{tw_g} = \frac{g}{\rho_0} \left(\frac{\partial u_g}{\partial x} \frac{\partial \rho}{\partial x} + \frac{\partial v_g}{\partial x} \frac{\partial \rho}{\partial y}, \frac{\partial v_g}{\partial y} \frac{\partial \rho}{\partial y} + \frac{\partial u_g}{\partial y} \frac{\partial \rho}{\partial x} \right), \tag{11}$$

where u_g and v_g are the horizontal components of the geostrophic velocity. In this study, these velocities are estimated by applying the thermal wind balance downward starting from the sea level with the reference velocities at that level being derived from the model sea level anomalies. This procedure mimics what can be optimally done with real oceanic data, assuming that sea level elevation is known with good accuracy.

The NG formulation takes into account the curvature of the total flow and the deformation of the ageostrophic flow:

$$\mathbf{Q}_{tw} = \frac{g}{\rho_0} \left(\frac{\partial u}{\partial x} \frac{\partial \rho}{\partial x} + \frac{\partial v}{\partial x} \frac{\partial \rho}{\partial y}, \frac{\partial v}{\partial y} \frac{\partial \rho}{\partial y} + \frac{\partial u}{\partial y} \frac{\partial \rho}{\partial x} \right), \tag{12}$$

$$\mathbf{Q}_{dag} = f \left(\frac{\partial v}{\partial x} \frac{\partial u_{ag}}{\partial z} - \frac{\partial u}{\partial x} \frac{\partial v_{ag}}{\partial z}, \frac{\partial v}{\partial y} \frac{\partial u_{ag}}{\partial z} - \frac{\partial u}{\partial y} \frac{\partial v_{ag}}{\partial z} \right), \tag{13}$$

where u_{ag} and v_{ag} are the ageostrophic horizontal velocities. In practice, u_{ag} and v_{ag} are obtained as the difference between

model horizontal velocities and the calculated geostrophic velocities u_g and v_g .

Two baseline reconstructions ω_{QG} and ω_{NG} are thus computed from the adiabatic QG and NG formulations of the \mathbf{Q} vector. For each of them the computation is made on the horizontal subdomains defined in section 2a. The domain extension in the vertical goes from the surface down to 1600 m. Dirichlet boundary conditions are used at all the frontiers. A tridimensional buoyancy frequency N^2 that varies both in the horizontal and in the vertical is used to solve all the inversions, although tests were run using a horizontally averaged profile and little differences were observed (not shown). Specific reconstructions using different resolution, domain size or boundary conditions are made to explore sensitivities and described in the corresponding sections.

To investigate the relative importance of different sources of errors we will carry out two forms of omega reconstruction with either perfect right-hand side (ω will be denoted ω_\dagger) or perfect boundary conditions (ω will be denoted ω^*). Perfect rhs omega inversions are computed using forcing terms that are derived from model vertical velocities and the lhs of (9):

$$\nabla \cdot \mathbf{Q}_\dagger = f^2 \frac{\partial^2 w_{\text{model}}}{\partial z^2} + \nabla_h(N^2 \cdot \nabla_h w_{\text{model}}). \quad (14)$$

Precisely, (14) is evaluated using second-order centered differences, i.e., in a way that is consistent with how the MUDPACK elliptic solver (Adams 1989) that we use is being implemented. For instance, along the x direction we use

$$\frac{\partial^2 w}{\partial x^2} = \frac{1}{\Delta x} \left[\frac{(w_{i+1} - w_i)}{\Delta x} - \frac{(w_i - w_{i-1}))}{\Delta x} \right]. \quad (15)$$

Alternatively, perfect boundary conditions can be imposed by applying w_{model} at the edges of the inversion domain. Note that this is, however, not precisely possible to do at the ocean surface because it is a moving interface in NEMO while it must be held fixed in the solving of the omega equation.

c. Baroclinic mode decomposition

Under the assumptions of flat bottom and rigid lid at the ocean surface, the linearized primitive (or quasigeostrophic) equations governing the horizontal and vertical components of inviscid fluid motion can be separated (Cushman-Roisin and Beckers 2011). In the vertical, two sets of normal mode eigenfunctions F_n and G_n form complete bases onto which pressure/horizontal velocity and vertical velocities can respectively be projected (see the appendix).

Normal mode decomposition has proved useful even in situations where all above assumptions are not satisfied and in particular when the bottom of the ocean is not flat (e.g., Rocha et al. 2013). To gain insight into the sources of reconstruction errors normal modes decomposition will be used to interpret the differences in reconstruction skills for the different regimes.

In practice the MODES program available at <http://www.d.umn.edu/~smkelly/software.html> (Kelly 2016) is used

to determine the mode structure at each model point. The modal amplitude of the model vertical velocity is then determined locally.

3. Description of the vertical velocity regimes

The four regions and two seasons selected for this study exhibit contrasted dynamical regimes. This will allow us to explore the sensitivity of the omega reconstruction behavior and skills to the nature of the meso/submesoscale turbulence that produces the vertical velocities. The diversity of vertical flow behavior is visually illustrated by snapshots of model vertical velocities (Figs. 1–5) and confirmed by their spectral distribution of variance (Figs. 6a,b).

In the LMX region, the displayed vertical section was chosen so as to cross the Gulf Stream (Fig. 2). In this region the EKE is about 10 times higher than in the other regions (Figs. 7c,d), the root-mean-square (rms) of w is 3–5 times higher (Figs. 7e,f), and intense density fronts are observed both in June and December (Fig. 2). In the frontal region, the vertical velocity structures extend down to 1500 m (Figs. 7e,f) depth or more, with peak $|w|$ in excess of 80 m day^{-1} . Overall, this is the region with the largest vertical velocity variance at all scales and a dominant fraction of this variance is found at the largest scales fitting in the study domain. Also note that LMX interior vertical velocities are only weakly affected by the seasonal cycle of the near-surface submesoscale activity (which is present but hardly visible in Fig. 1).

Compared to the LMX region, the AZO region is a lot less energetic (Figs. 7c,d), the isopycnics are flatter, and the horizontal velocities are slower. The vertical circulation is by far the weakest of all four regions (Figs. 7e,f and 6) and exhibits limited seasonal changes. Vertical velocities are organized into structures whose size is intermediate (smaller than in LMX but larger than in OSM and to a lesser extent REK) as readily apparent from the inspection of fields in the physical space and also from the spectral distribution of w variance (Fig. 6): in the range of scales larger than 10 km the w spectrum is flatter (resp. steeper) than that for LMX (resp. REK and OSM). The w patterns in Fig. 3 also tend to be tilted with respect to the vertical axis. This is a plausible indication that vertically propagating IG waves are involved in the generation of w [slanted phase lines associated with near-inertial waves are found in numerous studies, e.g., Furuichi et al. (2008)].

In the REK region, the average vertical velocity magnitude is $O(10) \text{ m day}^{-1}$ with localized higher values near mesoscale structures that can reach of up to 100 m day^{-1} and tend to have a large vertical extension (Figs. 1 and 4). In the physical space, w structures are frequently tilted with respect to the vertical although less so than in AZO. Again this is presumably the signature of near-inertial wave activity. But contrary to the AZO case, the horizontal patterns visible in w fields are consistent with the structuring role of the (sub) mesoscale activity, particularly during summer: the presence of multipolar w patterns (Lapeyre and Klein 2006b; Viúdez 2018), which could be the signature of eddy-induced Ekman pumping and/or vortex Rossby waves (McWilliams et al. 2003;

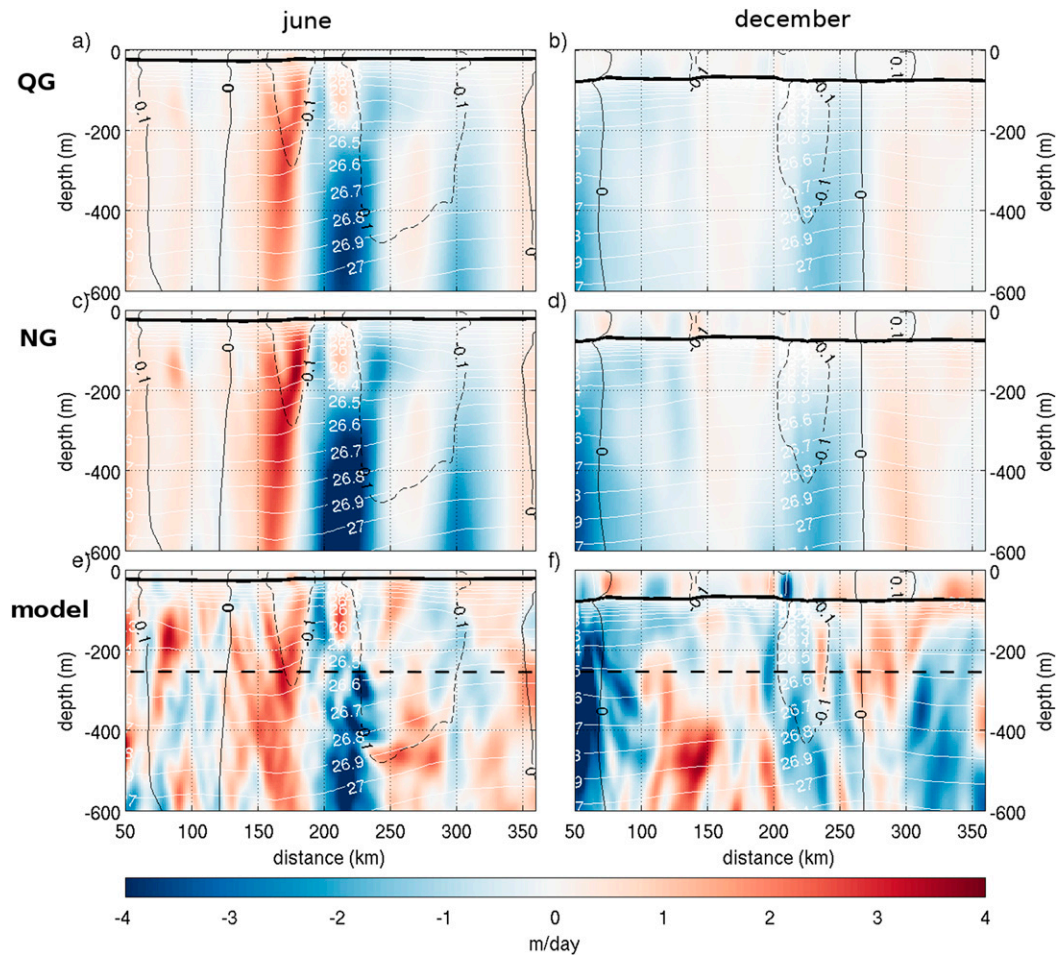


FIG. 3. As in Fig. 2, but for the AZO region.

Buongiorno Nardelli 2013; Barceló-Llull et al. 2017), as in the left corner of Fig. 1c, and elongated filaments of elevated w (Capet et al. 2016). During winter the enhancement of mixed layer submesoscale turbulence is particularly marked (Fig. 1c). Most of the finescale w patterns appear to be confined into the mixed layer and the visual aspect of the w field differs noticeably on either side of the mixed layer base. On the other hand, w increases significantly in magnitude from summer to winter (Figs. 7e,f). In terms of spatial scales, summer w have a clear mesoscale dominance with a peak around 50–100-km wavelength. For winter w , the role of the mesoscale is less prominent while the contribution associated with submesoscales is strongly reinforced (Fig. 6).

Vertical velocities in the OSM region share many similarities with those of REK: magnitude is $O(10)$ m day^{-1} , flatness of the w power spectrum in the mesoscale range 50–100 km, importance of the near-surface submesoscales particularly during the wintertime, and modest indication of vertical tilt. The main difference between the two regions is the greater degree of w continuity across the mixed layer and upper thermocline in winter, presumably as a

consequence of very low upper-ocean stratification in the OSM region (Fig. 7b).

4. Omega reconstruction of the vertical velocity fields

a. Baseline skill assessment

The skills of the omega reconstructions vary greatly depending on the region and season as illustrated qualitatively (as quantified in Figs. 2–5 (Fig. 8)). Figure 8 represents the spectral coherence between the model w and reconstructed ω , i.e., the degree of covariance between them as a function of scale. The scales and wavelengths above which the spectral coherence is larger than 0.6 for the QG and NG inversion are listed in Table 1 along with the fraction of w variance retained at scales larger than that threshold.

In general, omega reconstructions perform best at scales ≥ 10 –20 km. The energetic mesoscale structures are better reproduced and spectral coherence drops rapidly at scales ≤ 2 –10 km, i.e., the omega reconstruction are not well suited in the submesoscale range. This is particularly well illustrated by the LMX region where spectral coherence reaches levels close to 1 above 20 km scale but drops below 0.6 around 7–8 km

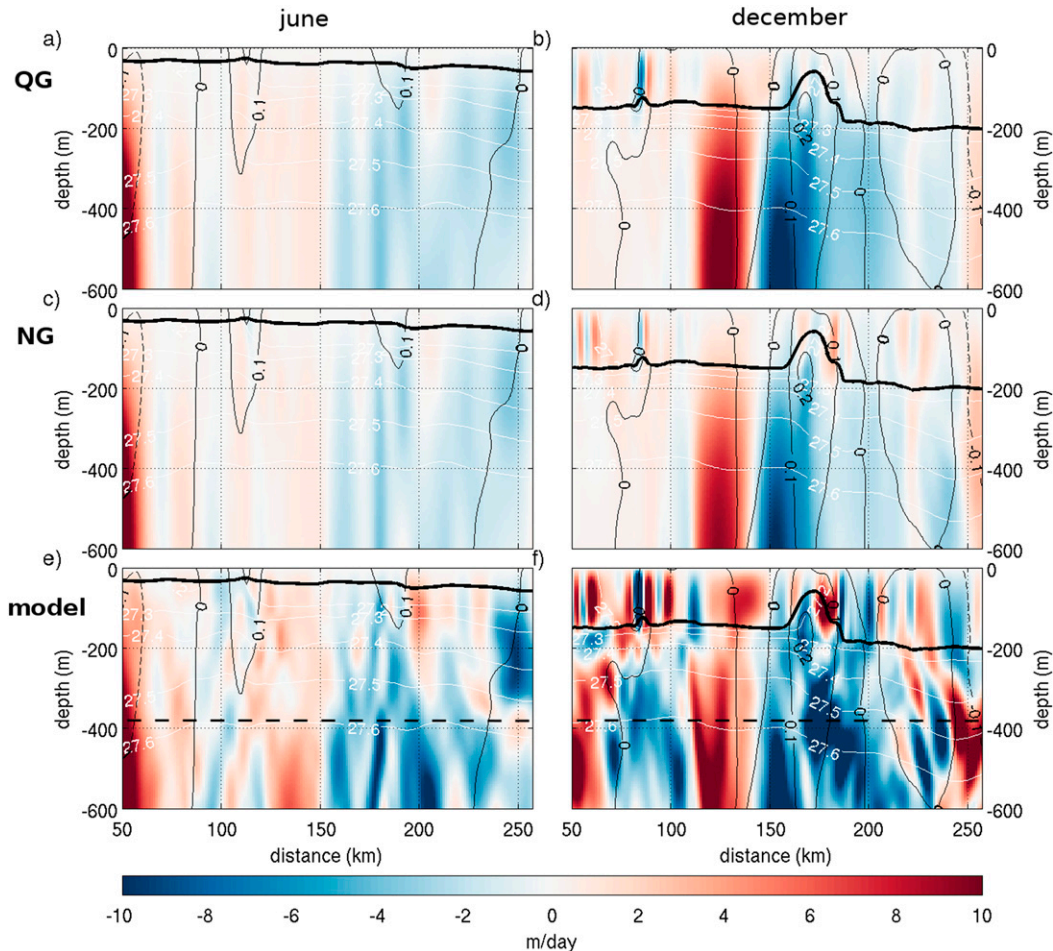


FIG. 4. As in Fig. 2, but for the REK region.

(Table 1). On the other hand, there is relatively limited w variance at finescale. Thus, the overall quality of the reconstructions is manifest for the randomly chosen situation shown in Fig. 2 where both the intensity and lateral/vertical extension of the w poles are well reproduced. Overall, 80%–90% of the w signal is captured with a coherence of 0.6 or higher in LMX.

Wintertime OSM is a notable exception where coherence levels remain elevated down to 3-km scales, particularly when ageostrophic effects are accounted for (see below). As in LMX, upper-ocean turbulent stirring and ageostrophic secondary circulations reach down to depths of 300–500 m in this region but the deformation radius is much smaller than in LMX and the stratification is particularly weak. The w field is thus organized in finescale structures that are frequently several hundred meters thick in the vertical (Fig. 5), many of which are well captured by the omega reconstruction.

In the REK region, the omega equation reproduces adequately some of the stronger and larger vertical velocity structures organized in alternated bands of upward and downward velocity intensity (Fig. 4). As for the finer slanted structures visible in both seasons but more intense in winter they are almost completely missed. The spectral coherence

peaks for scales of ~ 10 km with lower coherence for smaller and also larger scales. This might arise because vertical velocities at large scale are driven by other dynamical processes associated with the terms not retained in our omega reconstructions such as Ekman and inertial pumping.

Finally, in agreement with the visual impression drawn from Fig. 3, omega reconstructions perform very poorly for wintertime AZO with spectral coherence systematically below 0.5. Including ageostrophic effects has virtually no impact on the reconstruction skill for the AZO region as clearly illustrated in Fig. 3.

More generally, including the ageostrophic terms makes a significant difference for the LMX and wintertime OSM reconstruction only (the scale where the 0.6 coherence threshold is found decreases by up to 20% for these cases). We relate this to the fact that high Rossby number dynamics is prevalent in both regions albeit for different reasons. In the LMX sectors, the frontality of the Gulf Stream leads to intense turbulent stirring by the mesoscale field. In the OSM sector deep mixed layers in winter lead to intense submesoscale activity whose influence reaches to great depths owing to the reduced subsurface stratification. Conditions resembling those found for

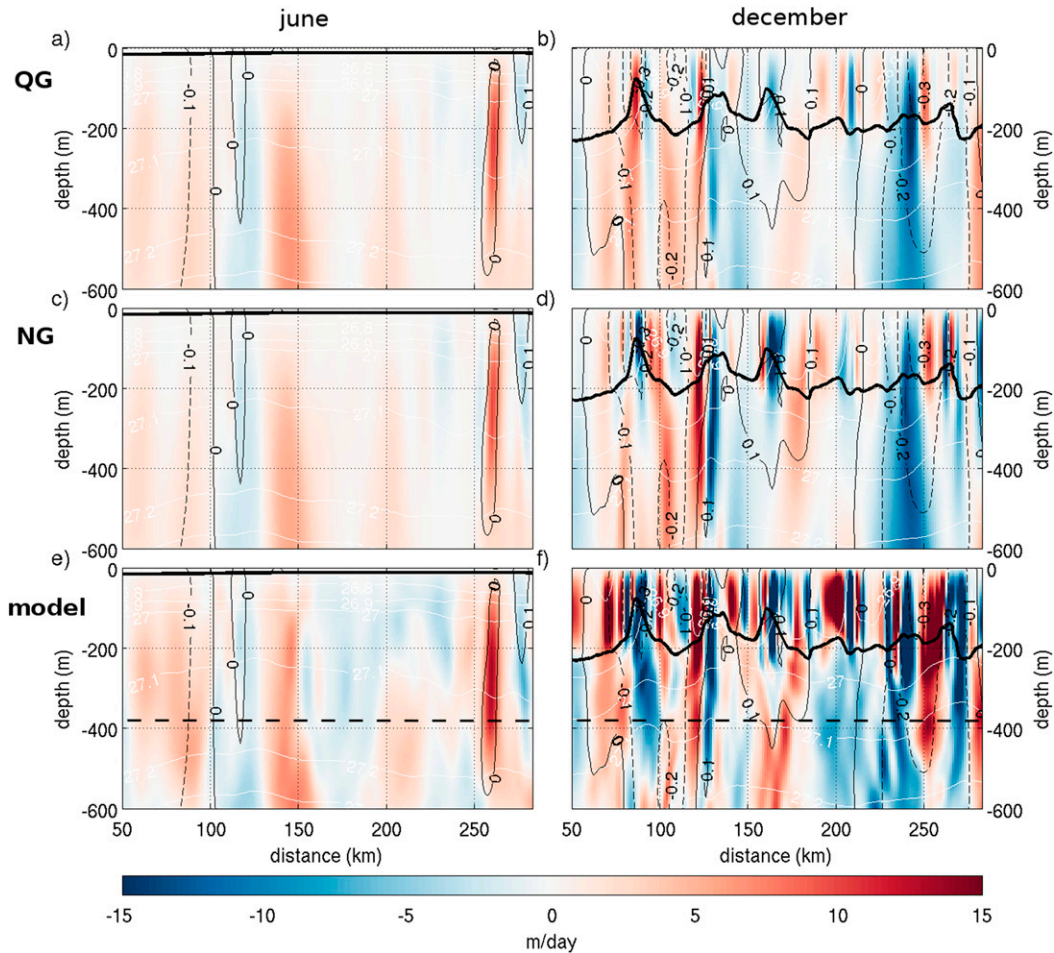


FIG. 5. As in Fig. 2, but for the OSM region.

winter OSM also exist in the REK region in winter but, in this latter case, the modest stratification present around 200-m depth (Fig. 7b) seems sufficient to isolate the interior circulation from the influence of the near-surface dynamics and thwart ageostrophic effects, hence the limited increase in coherence between ω_{QG} and ω_{NG} in Fig. 8.

b. Sensitivity to temporal averaging

An important limitation of the omega reconstructions presented in the previous section may arise from the role played by IG waves. This role cannot be captured by the forcing terms we retained for the expression of \mathbf{Q} in the QG and NG reconstructions, which ignores the material rate of change of the thermal wind imbalance [\mathbf{Q}_{dr} in (10)]. Using daily averages tends to reduce the effects of these processes, particularly where the inertial period is close to 24 h (at 30° latitude), but does not eliminate them. It is unclear whether the IG wave and subinertial quasi-balanced flow contributions to w should be separated (see discussion section) but the behavior of the omega reconstruction as a function of the temporal scale content of the forcing term $\nabla \cdot \mathbf{Q}$ is of practical as well as theoretical interest. This behavior is explored by increasing the

degree of temporal averaging that is performed prior to the computation of the omega reconstruction right-hand side. All simulated fields are averaged similarly, including w_{model} . The expectation is that more temporal averaging will reduce the amount of variance associated with the w field to be reconstructed but improve the skill of the omega reconstruction for that low-passed w . The existence of such a trade-off is confirmed in Fig. 9 for most dynamical regimes. Figure 9 presents the spectral coherence between ω_{NG} and w_{model} for larger (10–25 km) and finer (3–10 km) scales as a function of the remaining variance ratio $\text{Var}(\langle w \rangle_t) / \text{Var}(w)$, where $\langle \cdot \rangle_t$ is the low-pass temporal averaging whose length is varied from 1 day (standard output for that study) and 10 days. Results are shown for ω_{NG} but only minor differences were found with ω_{QG} .

For all regions except LMX an optimum in spectral coherence is found for time-averaging longer than 1 day, both for the large scales (10–25 km) and the fine scales (3–10 km). The averaging period associated with this optimum is between 2 and 5 days. As anticipated, spectral coherence improvement comes at the expense of a loss of reconstructed w variance. In several cases the loss of w variance is less than 25%: REK, AZO and OSM for the large scales and REK for the fine scales during

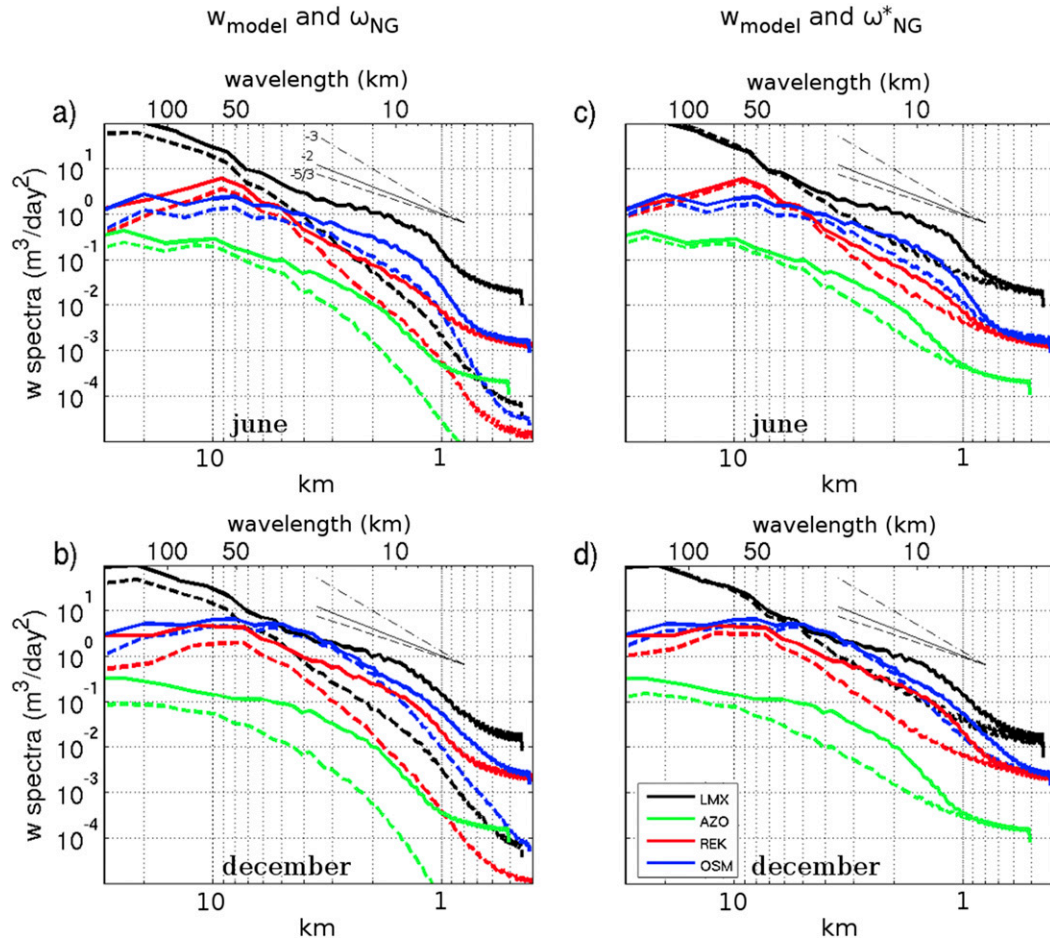


FIG. 6. (a),(b) Vertical velocity variance power spectra for w_{model} (solid) and ω_{NG} (dashed) at depths z_a in the LMX (black, $z_a = 220$ m), AZO (green, $z_a = 250$ m), REK (red, $z_a = 380$ m), and OSM (blue, $z_a = 380$ m) regions in (a) June and (b) December. (c),(d) As in (a) and (b), but for w_{model} (solid) and ω_{NG}^* (dashed, the reconstructed ω_{NG}^* is computed using perfect boundary conditions). Straight (resp. dashed and dot-dashed) lines indicate -2 (resp. $-5/3$ and -3) roll-offs. The bottom (resp. top) horizontal axis displays the length scale (resp. wavelength).

summer; OSM for the large scales during winter. Particularly for this latter case the spectral coherence improvement is substantial and 2–5 day averaging seems justified. Conversely, the most dramatic improvement in (ω_{NG}, w) spectral coherence seen in the AZO region (winter, 10–25-km scales, 5-day averaging) has a major effect on the retained ω_{NG} variance (only 37% of the original w variance) and may be of limited utility in practice.

c. Discretization errors

We also wish to draw the reader’s attention on the delicacies of the omega inversion and the implications these delicacies have on the estimation of vertical velocities. To do so we attempt to minimize reconstruction errors. In addition to imposing perfect boundary conditions by applying w_{model} at the edges of our domain, we also compute a “perfect” right-hand side $\nabla \cdot \mathbf{Q}_i$ to be used in the elliptic inversion (see section 2b). The resulting vertical velocity estimate ω_i^* found as a numerical solution is associated with minute errors, on average 3% with

the solving parameter choices we pass to MUDPACK. These errors result from 1) the imperfect convergence of the elliptic solver and 2) the imperfect boundary condition at the ocean surface.

To gain insight into the sensitivity of the elliptic inversion, variant ω_i^* were also computed by employing a discretization scheme alternative to (15) to calculate $\nabla_i^2 w_{\text{model}}$ in (14):

$$\frac{\partial^2 w}{\partial x^2} \Big|_i = \frac{1}{2\Delta x} \left[\frac{(w_{i+2} - w_i)}{2\Delta x} - \frac{(w_i - w_{i-2})}{2\Delta x} \right] \quad (16)$$

Both schemes are consistent, centered, and second-order in accuracy. They only differ by the length of their stencil (3 or 5 points).

We find that the discretization choice has a moderate but significant impact. In the LMX region in June for instance the use of the scheme (16) to compute the perfect rhs produces relative w errors in excess of 12%, to be compared with the minimal convergence errors of $\sim 3\%$ with scheme (15).

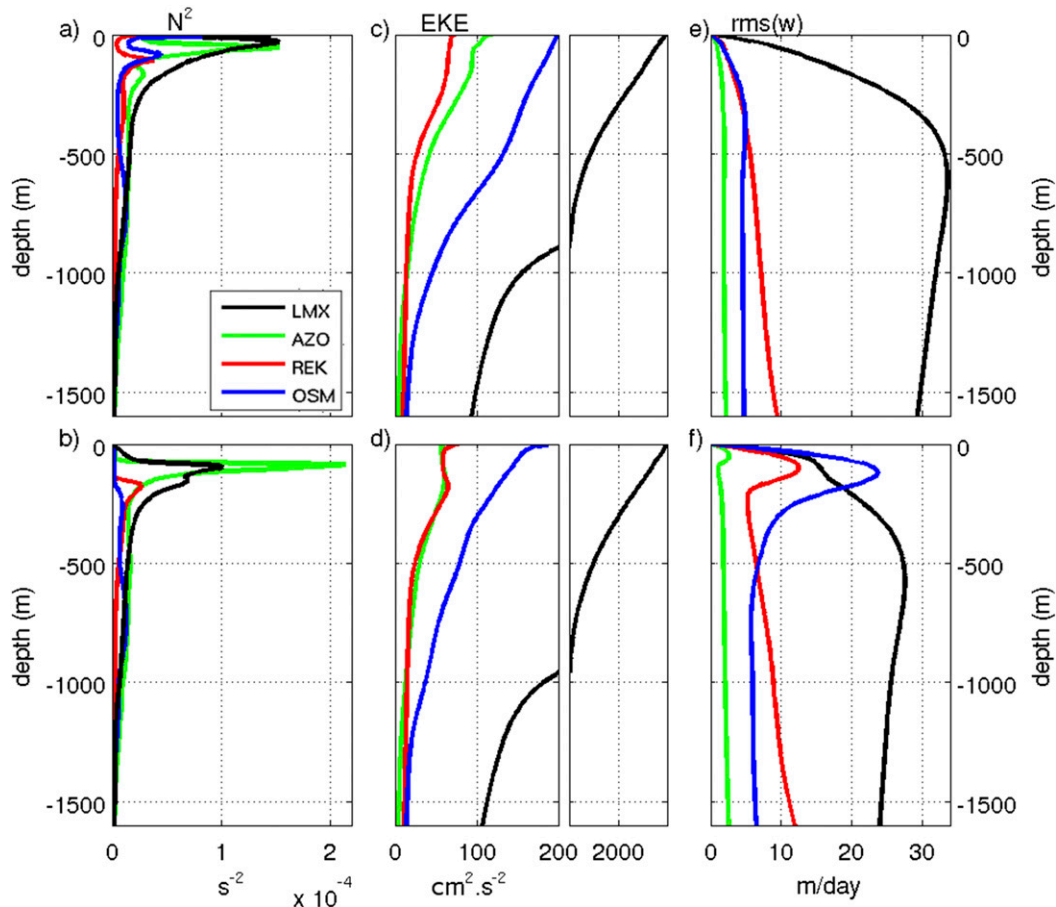


FIG. 7. Averaged vertical profiles of (a),(b) buoyancy frequency; (c),(d) eddy kinetic energy; and (e),(f) vertical velocity root-mean-square for the LMX (black), AZO (green), REK (red), and OSM (blue) regions in (top) June and (bottom) December. Note the subpanels with a change of scale to accommodate the large differences in EKE between LMX and the other regions.

The elevated error figure obtained with the second scheme helps place realistic limits to vertical velocity estimation attempts and draws the attention on the fact that minute errors/uncertainties on the rhs of the omega equation can have significant impacts on the reconstructed vertical velocities.

d. Sensitivity and errors due to boundary conditions

Finally, an important source of error when solving the omega equation is the choice of the bottom boundary condition. As numerous studies have pointed out (Allen and Smeed 1996; Rudnick 1996; Pascual et al. 2004) ω is not very dependent on the lateral boundary conditions but the form of the BBC and the depth where it is applied can have a strong impact on the solution of the equation (Pidcock et al. 2013; Rudnick 1996). Two types of BBCs are generally used, the Dirichlet condition $w = 0$ or the Neumann condition $\partial w / \partial z = 0$. A third one was proposed by Rudnick (1996): to limit the influence of the boundary condition on the solution, the vertical grid was extended below the depth where observations were available. Over the depth range with no observations $\nabla \cdot \mathbf{Q}$ was simply assumed to be zero. In practice, this led Rudnick (1996) to shift

the bottom boundary from 400 m (where the deepest observation was available) to 2524 m depth where a Dirichlet BBC ($w = 0$) was imposed.

To assess the impact of the BBC on the solution we compare the errors obtained for Dirichlet and Neumann conditions and for different domain extensions, with the bottom boundary depth z_{bottom} ranging from 150 to 1600 m (Fig. 10). This is done in the four regions and for the two seasons. We also test the boundary condition proposed by Rudnick (1996) with two different depths (400 and 800 m) below which $\nabla \cdot \mathbf{Q}$ is assumed to be zero down to 800 and 1600 m, respectively. In this section, the omega reconstruction is solved using $\nabla \cdot \mathbf{Q}_+$ forcing computed as described in section 2b so relative errors exceeding a few percent can only be due to boundary conditions.

Figure 10 shows the relative errors of omega reconstructions made at the reference depths z_a when different boundary conditions are applied. The errors at depth z_a are horizontally and temporally averaged over the entire reconstruction domain on the horizontal and over the 11 consecutive daily model outputs. The relative error field can be quite large, generally in excess of 0.5 for domain depths shallower than 500 m, despite a

TABLE 1. Statistics derived from the spectral coherence field (Fig. 8). Note that relating the horizontal scale $L_{0.6}$ to the deformation radius R_d , which is ~ 15 km in the OSM and REK regions and ~ 30 km in the LMX and AZO regions, yields a performance threshold of $0.3R_d$ on average, with a minimum $0.17R_d$ in winter OSM and a maximum $0.49R_d$ in winter REK. In the AZO region in December the squared coherence never exceeds 0.6 thus leading to a fraction of variance null.

	Horizontal scale $L_{0.6}$ (km) [and wavelength (km)] below which the squared coherence between true and reconstructed vertical velocities is below 0.6. Values are provided for QG and NG reconstructions.				Fraction of vertical velocity variance retained by the reconstructed vertical velocity over the scale range $L > L_{0.6}$. Values are only provided for the NG reconstruction.	
	Jun		Dec		Jun	Dec
	QG	NG	QG	NG		
LMX	9.2 (58)	6.7 (42)	9.5 (60)	7.8 (49)	0.90	0.84
AZO	6.4 (40)	6.3 (40)	—	—	0.74	0
REK	4.7 (30)	4.6 (29)	7.3 (46)	6.5 (41)	0.87	0.61
OSM	3.9 (25)	3.9 (25)	3.0 (19)	2.6 (16)	0.72	0.88

good agreement in w and ω patterns. This is generally consistent with the findings of several past studies (Pinot et al. 1996; Allen et al. 2001; Rixen et al. 2003).

Relative errors vary with z_{bottom} in a relatively simple and intuitive way: deeper z_{bottom} systematically translates into weaker errors with Neumann or Dirichlet boundary condition. More precisely, errors tend to stabilize around 20% when z_{bottom} greater than 1000–1500 m are used (Table 2), with the notable exception of REK for which relative errors are ~ 0.3 – 0.4 and still decreasing at $z_{\text{bottom}} = 1600$ m. The most evident manifestation of error saturation behavior is found for LMX and Neumann boundary conditions, with a threshold around $z_{\text{bottom}} = 700$ m beyond which little improvement is observed. More generally we note that: the quality of the reconstruction can vary from low to high depending on the choice being made for the depth of the boundary condition; except in the LMX region the Neumann and Dirichlet boundary conditions give relatively similar results; the use of a buffer region to increase the domain depth following Rudnick (1996) can ameliorate the reconstruction skill and provide an optimum over all possible choices but the improvement is generally marginal. Estimating the relative error as a volume-average over the whole water volume above the depth where the boundary condition is applied paints a more complicated picture. For instance, in a number of situations, reconstruction errors are found to increase when the location of the boundary layer deepens beyond certain thresholds (not shown). We relate this to the fact that deep vertical velocities can be particularly intense owing to processes not properly resolved by our implementations of the omega reconstructions (IG wave activity, flow–topography interaction).

In an attempt to gain further insight, we use vertical mode decomposition to characterize the vertical structure of the w field. Consider a situation where w would project onto a single mode, for instance baroclinic mode 2. In that case, w systematically vanishes at the zero crossing of G_2 and a Dirichlet boundary condition at that depth would not introduce any error. A perfect Neumann boundary condition would similarly exist at the zero crossings of F_2 . There

are several potential obstacles preventing this from happening. Most importantly, vertical velocity tends to be associated with fine horizontal scales and this tends to be also true in the vertical (Figs. 3–5). The vertical mode decomposition of w generally involves many modes, and in most cases, less than half of the w signal is explained by the sum of the first three vertical modes (Fig. 12, appendix). It is therefore impossible to choose a fixed depth where the dominant F_n or G_n are zeros. The LMX regime is an exception where the w vertical structure projects to a large extent onto the first three baroclinic modes (Fig. 12). For LMX the smallest reconstruction errors (volume averaged) are obtained for a Neumann boundary condition placed at approximately 750-m depth (not shown), i.e., the depth below which errors evaluated at a single depth ($z_a = 220$ m) tend to reach a plateau around 0.2–0.25 (Figs. 10a,e). The 750-m value falls in the depth range where F_1 and F_3 are very close to zero (Fig. 11a). The link between the vertical mode’s structure and the effect of z_{bottom} on the reconstruction skill is far less clear for the other three regions. We attribute these inconsistencies to the limited role played by the gravest baroclinic modes in AZO, REK, and OSM (Fig. 12).

To better characterize the reconstruction errors due to imperfect (bottom) boundary conditions, we compute an alternative omega reconstruction ω_{NG}^* using perfect w_{model} information at the boundaries (see section 2b). Differences between ω_{NG}^* and w_{model} can only arise from the simplification made to \mathbf{Q} in the NG inversion (i.e., considering \mathbf{Q}_{dr} , \mathbf{Q}_{th} , and \mathbf{Q}_{dm} as null). They range from ~ 1 to ~ 5 m day $^{-1}$ depending on the region and season. They are typically 1–3 times as high as errors associated with a Dirichlet boundary condition at 1600 m (Table 2) and generally lower than the ones associated with a Dirichlet boundary condition at 500 m (Fig. 10). Discrepancies in spectral distribution between ω_{NG}^* and w_{model} are much reduced compared to those for ω_{NG} . This is particularly true for the scales below 1 km (Figs. 6c,d). A similar skill improvement is noticeable for the spectral coherence (Fig. 8). A significant fraction of the finescale vertical motions in the upper ocean is thus directly linked to unidentified processes active below 1600 m. Consistent with previous studies (e.g., see

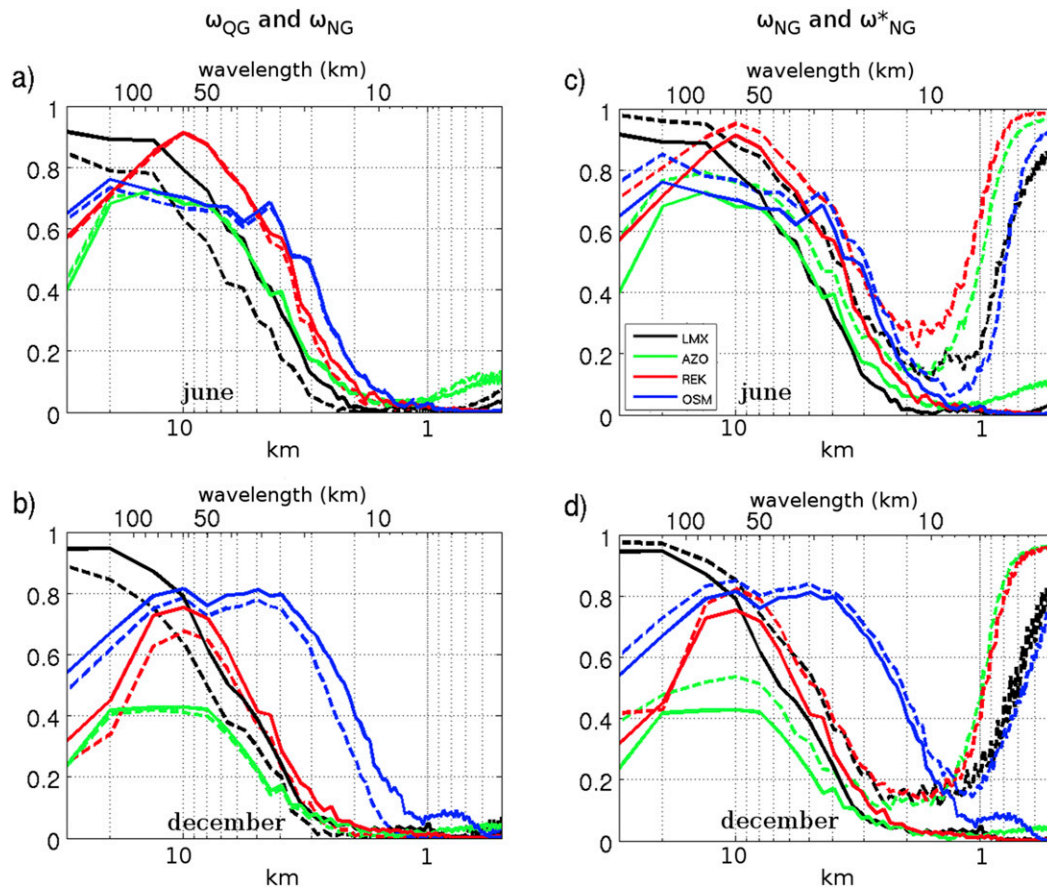


FIG. 8. Spectral coherence as a function of scale between w_{model} and ω_{NG} (resp. ω_{OG}) at depths z_a is represented by a solid (resp. dashed) line in the LMX (black), AZO (green), REK (red), and OSM (blue) regions in (a) June and (b) December. (c), (d) As in (a) and (b), but the dashed line represent the coherence between w_{model} and ω_{NG}^* (the reconstructed ω_{NG}^* is computed using perfect boundary conditions). The bottom (resp. top) horizontal axis displays the length scale (resp. wavelength).

Jouanno et al. 2016) flow–topography interactions is a likely candidate, which would pose serious difficulties if it were to be explicitly incorporated into the omega reconstruction approach (inversion would need to be performed for the whole ocean depth range with the knowledge of the vertical velocity at the ocean floor $w = -\mathbf{u}\nabla h$). We do not see any prospect for this, but this source of errors should be kept in mind.

5. Discussion

Since the 1980s a large number of studies have focused on inferring oceanic vertical velocities through more or less elaborate forms of the omega equation. Most of them have been applied to local in situ data, in which the sources of errors are difficult to identify and quantify (see the supplemental material). In some instances the reconstructed vertical velocity field has been used to qualitatively interpret concomitant tracer distributions (Pollard and Regier 1992; Rudnick 1996; Martin and Richards 2001; Allen et al. 2005; Ruiz et al. 2009; Pallàs-Sanz et al. 2010; Rousselet et al. 2019). The uncertainty on reconstructed w is an important limitation when doing so (also note that the vertical

tracer distribution at any given time reflects the past history of vertical advection—and mixing). The w velocities are now increasingly being used quantitatively, e.g., as inputs to tracer models (Pascual et al. 2015; Barceló-Llull et al. 2016).

Model studies have addressed the various sources of errors involved and the ways to reduce their impact with a general focus on three main issues: (i) the merits of more complete versions of the omega equation; (ii) the sensitivity to particular choices of boundary conditions when solving the omega equation; and (iii) the errors induced by the lack of resolution, homogeneity, and synopticity of in situ sampling when the omega reconstruction is applied to real ocean observations (see the supplemental material). These model-based assessments of ω reconstructions were typically performed in simplified flow conditions, composed of a single coherent eddy or front, with a marked preference for the early stages of destabilization of baroclinic zones (Strass 1994; Pinot et al. 1996; Allen et al. 2001; Rixen et al. 2003; Viúdez 2018). More recently, the omega equation has been used to determine oceanic vertical velocity fields over larger domains and extended time periods from observation-based gridded reconstructions of

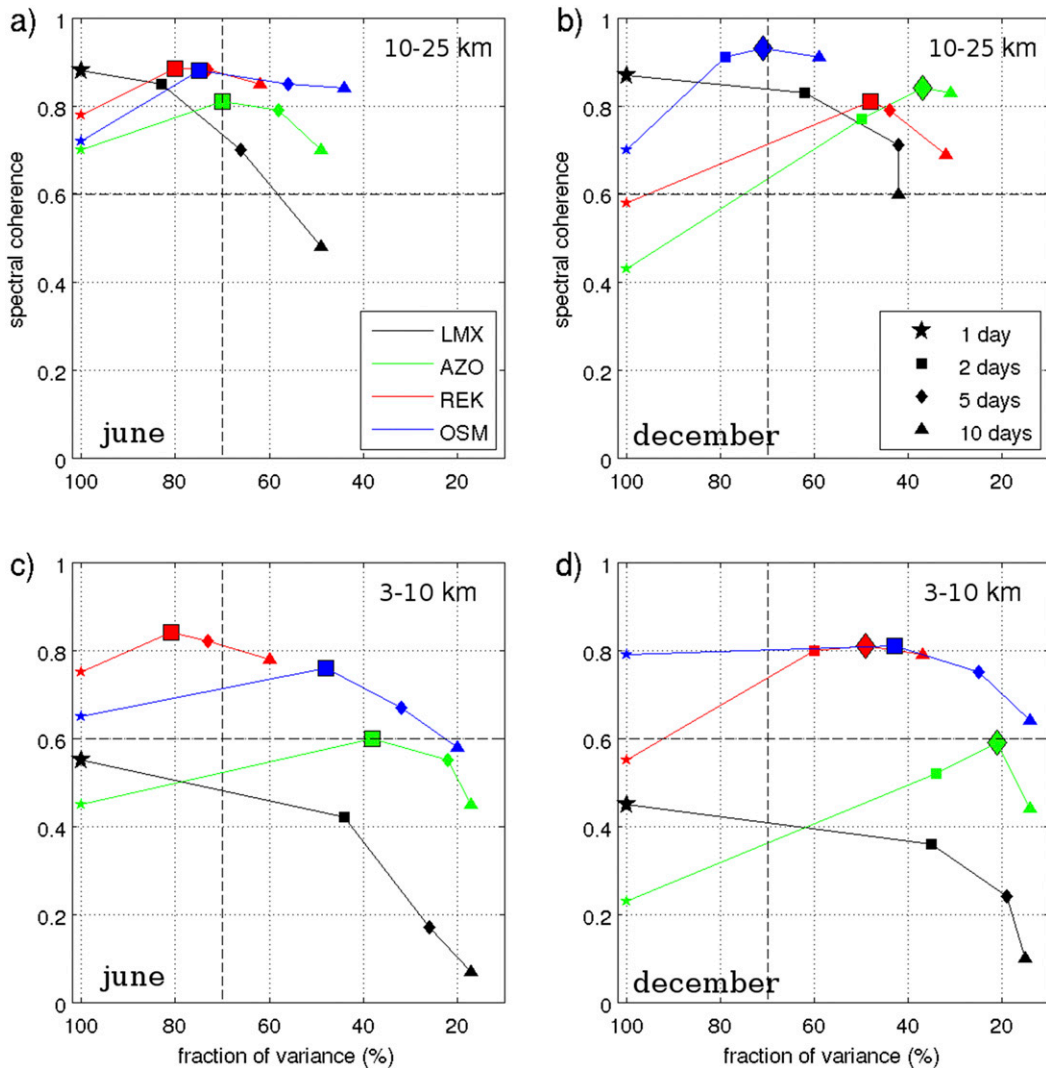


FIG. 9. Spectral coherence between w_{model} and ω_{NG} at depths z_a averaged over the scale ranges (a),(b) 10–25 km and (c),(d) 3–10 km as a function of the fraction of variance retained by each averaging intervals. Results are shown for the LMX (black), AZO (green), REK (red), and OSM (blue) regions in (left) June and (right) December. The markers represent the timespan of the averages: the star is 1 day, square 2 days, diamond 5 days, and triangle 10 days. The highest coherence is indicated by a larger marker with a black contour. The black dashed lines mark a fraction of variance of 75% and a coherence of 0.6.

temperature and salinity fields (Pascual et al. 2015; Barceló-Llull et al. 2016; Buongiorno Nardelli et al. 2018). The resulting ω fields can subsequently be used to estimate vertical fluxes, and this has been attempted for several biogeochemical tracers, over different relevant time scales.

Overall, a general assessment of the skills and limitations of the omega reconstruction is lacking. The present study is an effort in that direction that mainly pertains to errors i and ii, and we deliberately excluded errors of the type iii. Although they pose important and perhaps leading-order limitations to w reconstructions from observations, this type of errors could be significantly reduced by observing the ocean at higher resolution in specific regions of interest and combining these observations with finescale remotely sensed information and model

integrations. Optimal ways to produce such combinations may involve relocation strategies as proposed in Rixen et al. (2001) and Pascual et al. (2004). Errors i and ii pose different type of challenges that have not received much attention. We have followed the steps of Strass (1994) whose analyses of the omega reconstruction skills included a scale-dependent coherence diagnostic. An originality of our study is to estimate the fidelity of w reconstructions for fully turbulent realistic flows. Several findings reported in the previous section deserve further discussion.

a. On the limited skills of omega reconstructions at finescale

First, the important limitations to the omega reconstructions found at relatively finescale, despite the absence of type iii

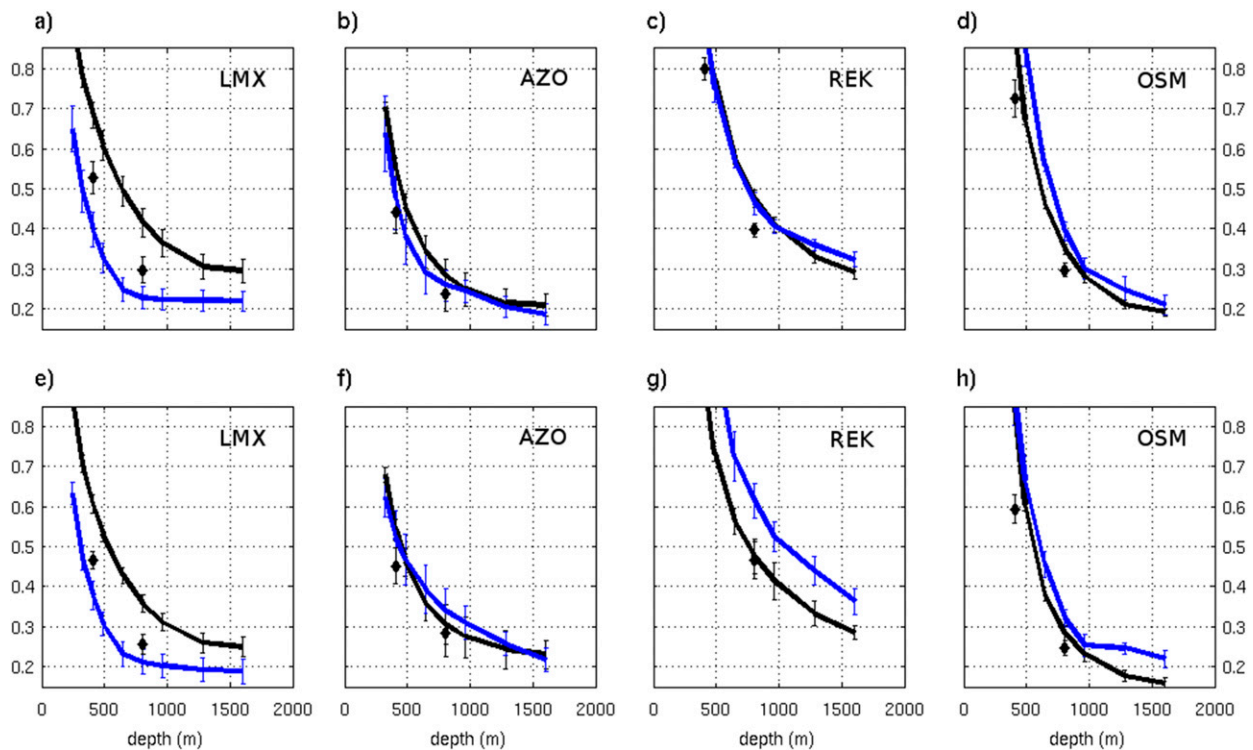


FIG. 10. Relative error, at depths z_a , between w_{model} and ω_{\dagger} in (top) June and (bottom) December in the (a),(e) LMX; (b),(f) AZO; (c),(g) REK; and (d),(h) OSM regions as a function of the depth where the bottom boundary condition (z_{bottom} in the text) is imposed. Errors for a Dirichlet (resp. Neumann) BBC are represented with a black (resp. blue) line. Black diamonds indicate error values for a Dirichlet boundary condition modified as in Rudnick (1996) (see section 4d for details). The vertical bars show the standard deviation over the 11 daily averages used for our analysis.

errors may seem surprising. Some remarks are thus in order. An important motivation for this work is the representation of vertical tracer fluxes between the surface and interior of the ocean. Therefore we focused on vertical velocities tens of meters below the mixed layer. At such depths, vertical velocity spectra are significantly more red than closer to the surface where whiter spectra have been reported (Ponte and Klein 2013; Klein et al. 2008). This means that modest levels of variance remain in the scale range where coherence has dropped down. Conversely, a large fraction of w variance ($>60\%$ for all regions and seasons) is found in the scale range where coherence is above 0.6 (Table 1).

This being said, the coherence drop off at scales below 10 km is interesting in itself and deserves further discussion. The only regime for which significant skill is being obtained into

submesoscale range is winter OSM, which clearly stands out in terms of coherence (Fig. 8). This is also the only regime for which the influence of surface intensified frontal turbulence is found to penetrate down to the depth of analysis ($z_a = 380$ m), owing to extremely weak subsurface stratifications. This regime is also the one for which the NG reconstruction provides the largest improvement over the QG reconstruction. The lack of reconstruction skills generally found at submesoscale is therefore not due to a particular difficulty at capturing vertical velocities associated with frontal turbulence. Instead, we attribute reconstruction errors to the imperfection of the BBC used to invert the elliptic omega equation and to the neglect of fluctuations in thermal wind imbalance due to partly or fully unbalanced fast processes such as inertial gravity waves. The former source of error is identified and quantified explicitly

TABLE 2. Averaged errors (m day^{-1} and percentage of w_{model}), at depths z_a , for ω with perfect right-hand side (ω_{\dagger}), perfect boundaries (ω_{NG}^*), and both (ω^*). See section 2b for the detail on the computation of ω .

	ω_{\dagger}^*		ω_{\dagger}		ω_{NG}^*	
	Jun	Dec	Jun	Dec	Jun	Dec
LMX	0.41 (3%)	0.54 (5%)	4.08 (28%)	2.85 (24%)	4.87 (34%)	5.21 (44%)
AZO	0.04 (3%)	0.04 (3%)	0.25 (20%)	0.27 (23%)	0.80 (65%)	0.95 (79%)
REK	0.13 (4%)	0.11 (3%)	0.96 (29%)	1.19 (29%)	1.47 (45%)	2.85 (69%)
OSM	0.06 (2%)	0.20 (4%)	0.66 (19%)	0.90 (16%)	2.06 (58%)	2.96 (52%)

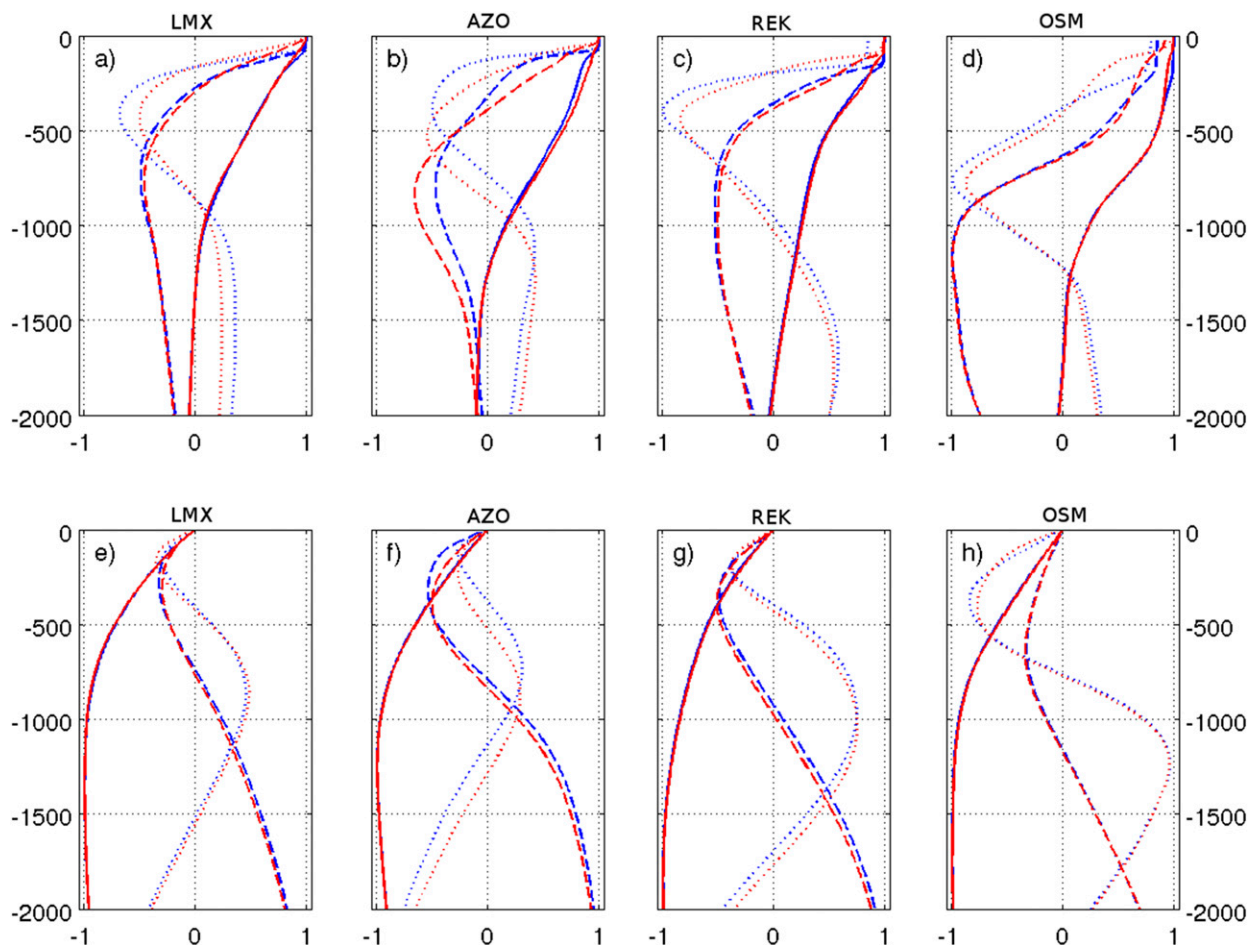


FIG. 11. First three baroclinic pressure modes for the (a) LMX, (b) AZO, (c) REK, and (d) OSM region in June (red) and December (blue). First three vertical velocity modes for the (e) LMX, (f) AZO, (g) REK, and (h) OSM region in June (red) and December (blue). The first mode (resp. second and third modes) is represented with a solid (resp. dashed and dotted) line. The amplitudes of the modes have been scaled so as to vary between -1 and 1 .

while the importance of the latter is inferred indirectly. Further elaboration on these two error sources follow.

The boundary conditions impact omega reconstructions over the entire study domain owing to the elliptic nature of the omega equation. It has been known for a long time that the solution can be improved by properly choosing the type and location of the BBC. We explored the possibility that the BBC strategies could be rationalized based on vertical mode decompositions. Although one exception was found in the Gulf Stream region, vertical velocities tend to project onto a large number of vertical modes (Fig. 12) and no ideal positioning of the BBC can be chosen a priori. Pushing the BBC position z_{bottom} toward greater depths leads to reconstruction improvements even beyond 1500 m. This is not feasible when observations are used to perform omega reconstructions and the typical choices made in this situation ($z_{\text{bottom}} \sim 500$ m; see the supplemental material) fall in the range where errors are most sensitive to z_{bottom} . In the Gulf stream region, vertical velocities associated with upper-ocean frontal dynamics extend deep into the water column and tend to project mainly

onto the first three vertical modes (Fig. 12). These properties are qualitatively consistent with the fact that, for the LMX domain, the Neumann BBC outperforms the Dirichlet BBC and that the relative error curve as a function of BBC depth levels off for z_{bottom} below ~ 700 m. Similar behaviors may presumably be observed in other western boundary regions and the Antarctic Circumpolar Current region. Finally, note that, given a depth down to which data are available to feed an omega reconstruction, the so-called telescopic method developed by Rudnick provides at best a marginal improvement over the classical Dirichlet or Neumann boundary conditions.

The presence of inertia-gravity waves near and beyond the drop-off scale is not accounted for as a source of w (in our case, neglecting the forcing term \mathbf{Q}_{dr} was due to limitations on the model outputs at our disposal but including this term for real ocean situations would pose extremely stringent if not impractical requirements on sampling strategies). The AZO region is the one where inertia-gravity waves are suspected to play the largest role, if only because balanced turbulence is

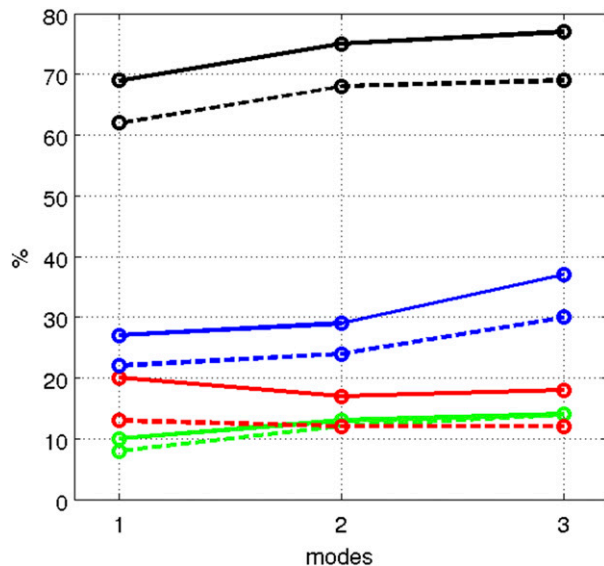


FIG. 12. Percentage of vertical water profiles for which the projection of model w on the subset of the gravest vertical modes (1, 2, or 3 gravest modes) leads to an approximation of w with a relative error that is less than 50%, i.e., percentage of profiles where $\varepsilon/w < 0.5$ [ε being defined in Eq. (A8)]. Percentages are represented for LMX (black), AZO (green), REK (red), and OSM (blue) in summer (plain lines) and winter (dashed lines).

very weak there. There may be other more subtle consequences of the inertia–gravity wave activity on the omega reconstruction we have presented. For instance, REK has more w variance in winter but coherence between ω_{NG} and w_{model} is degraded at that season compared to summer. Although seasonal changes appear comparatively unimportant in the LMX region, a slight coherence degradation occurs in winter: winter coherence for the NG reconstruction is ~ 0.1 lower than in summer over the scale range 2–8 km (gaps are smaller at larger scale but the degradation is noticeable up to 20 km). Again, signs of enhanced NIW activity for the winter analysis period compared to the summer one suggest that near-inertial waves are responsible for this degradation.

For completeness, two other possible sources of omega reconstruction errors are listed here. Daily averaged variables are used to compute the forcing terms which are nonlinear combinations of these variables. This means that “Reynold-type” forcing terms should formally be included due to rapid (intraday) fluctuations in u , v , and ρ . Kinetic energy corresponding to motions with temporal scales below 1 day is very low in simulations with $dx \sim 2$ km and we have assumed that this is a small effect. It would need to be verified that the same applies when in situ observations are low-pass-filtered prior to computing rhs terms for a real ocean omega reconstruction. Forcing terms associated with mixing of momentum and buoyancy have also been ignored. This is generally considered as a valid approximation sufficiently far from the mixed layer while this source of error is expected to increase at shallower depth (see below for more elaboration on this) in the context of ocean tracer dynamics. Although generally consistent with our

findings the latest results of Qiu et al. (2020) cast some doubts on the smallness of w forcing by buoyancy and momentum diffusion. In a western boundary context, this term is found to be comparable to w forcing by ageostrophic kinematic deformation and stronger than w forcing by thermal wind imbalance deformation or thermal wind imbalance tendency, even well into the ocean interior (see their Fig. 4). This issue would, however, need to be scrutinized based on an explicit computation of the mixed effect on w . Qiu et al. (2020) infer this term as a residual which therefore lumps together various contributions including that due to BBC errors. Model output limitations have not allowed us to undertake this.

b. On the consequences of limited reconstruction skills at fine scales for tracer dynamics

Ultimately, what matters in most applications is not vertical velocity per se but the vertical transport of oceanic properties, for instance upward fluxes of dissolved iron enriching the euphotic layer in the Southern Ocean (Nicholson et al. 2019), or downward fluxes of organic material filling the ocean interior and having key implications on O_2 consumption and distribution (Boyd et al. 2019). Vertical tracer fluxes arise from vertical motions because upward and downward motions are, on average, associated with distinct tracer values. We find it useful to discuss three types of limit cases in which this can occur and where the limitations of the omega reconstruction described above may be more or less problematic depending on the scenario. To this end, we consider a nutrient-like tracer τ whose concentration gradient is directed downward. The same reasoning would apply to tracer with an upward gradient (e.g., phytoplankton or dissolved oxygen).

An important distinction between two types of limit cases can be made depending on whether baroclinic instability (BCI) plays an important role or not.

1) CASE 1

When BCI is important, turbulent vertical motions are strongly constrained by the fact that $\langle w'b' \rangle > 0$, where b refers to buoyancy, $\langle \cdot \rangle$ is a low-pass averaging operator and the prime notation refers to deviations from this average. BCI can thus contribute to transport a tracer τ vertically provided that the distributions of τ and b are correlated. Far enough from the surface in the ocean interior available potential energy (APE) release is mainly achieved through mesoscale velocities (Lapeyre et al. 2006; Capet et al. 2016). Unless the correlations of τ and b were weak at mesoscale and strong at submesoscale, the proper description of mesoscale w would thus suffice to capture the dominant part of $\langle w'\tau' \rangle$. Typically, correlations between a tracer τ and buoyancy exist owing to processes acting at large scale, although large decorrelation can ensue from biological reactions (e.g., for oxygen and nitrate). Correlations are then transferred down to smaller scales by turbulent cascades and decorrelation between buoyancy and τ is preferentially found at the finest scales [see discussion in Fox-Kemper et al. (2013, and references therein)]. Therefore, we expect omega reconstructions to be useful in this case despite the limitations we have described.

We now turn to the situation where vertical motions are not constrained by baroclinic instability and result instead from,

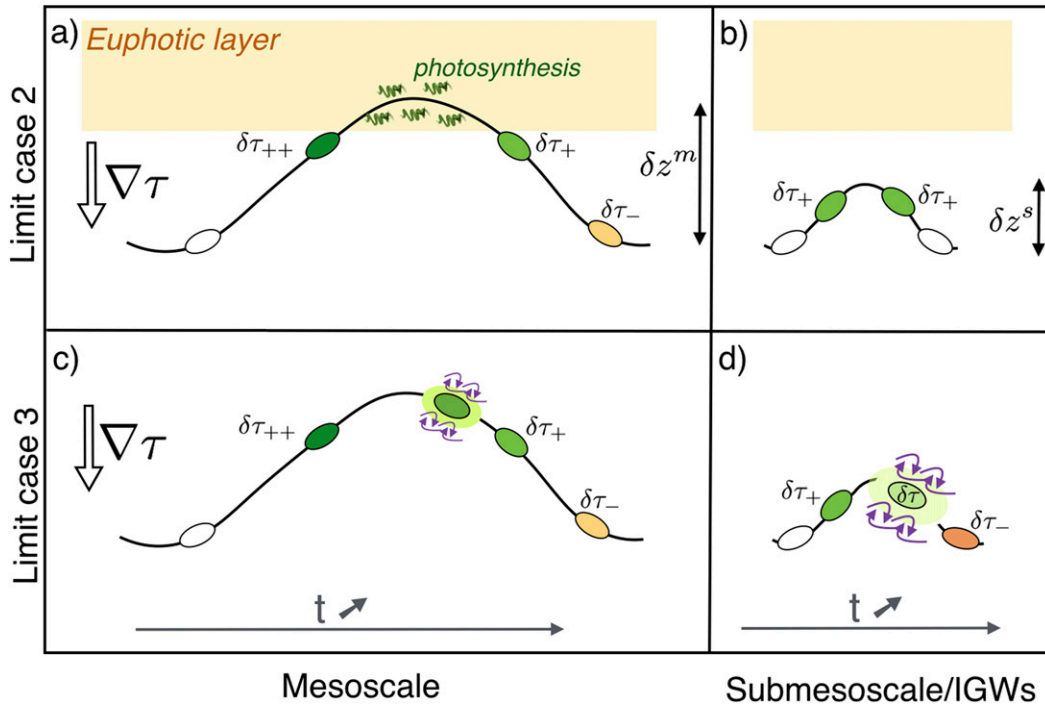


FIG. 13. Schematic representation of the vertical disturbance (y direction) and associated tracer anomaly temporal evolution undergone by a water parcel in four different situations. (top) A nutrient-like tracer τ with a photosynthesis-like sink term ($\nabla\tau$ is positive downward) in the upper ocean is subjected to (a) a mesoscale and (b) a submesoscale vertical oscillation (respectively of amplitude δz^m and δz^s). (bottom) A tracer with no sink-source term is subjected to (c) a mesoscale and (d) a submesoscale vertical oscillation, with a diabatic redistribution of tracer through shear driven mixing (purple arrows). In all panels, the horizontal direction represent time, increasing from left to right. Positive (resp., negative and null) tracer concentration anomalies ($\delta\tau$) relative to the parcel depth are represented with green (resp. orange and white) colors. The + and - symbols also provide indications on tracer anomalies (resp. positive and negative). In limit case 2, submesoscale oscillation have a lesser impact than mesoscale ones while the opposite may be true in limit case 3 if submesoscale oscillations are more effective to produce shear-driven turbulence.

e.g., fossil mesoscale/submesoscale turbulence and/or internal wave activity. In this situation, any water parcel has a well-defined reference depth position z_{ref} (quite independent of horizontal position) around which it may oscillate, and an additional process must be invoked for these oscillations to produce τ vertical/diapycnal fluxes. We take a Lagrangian viewpoint and consider $\delta z = \int_t w dt$, the vertical displacement of a water parcel from its equilibrium position and we distinguish two limit cases illustrated in Fig. 13.

2) CASE 2

Let us consider the situation in which a sink term for τ is located in the euphotic layer where photosynthesis draws τ levels down (Figs. 13a,b). The τ anomalies are schematically represented for a mesoscale (Fig. 13a) and a submesoscale (Fig. 13b) upward oscillation where the amplitude of the oscillations have been chosen so as to reflect the distribution of w variance found in the ocean interior for all regimes ($\delta z^m > \delta z^s$). Because $\delta z^m > \delta z^s$ and also because the mesoscale excursions last on average longer the photosynthesis sink term for τ is more effective at mesoscale than it is at submesoscale. The tracer biogeochemistry acts as a low-pass filter and a representation

of finescale vertical motions is unessential in the determination of tracer vertical fluxes. The deficiencies of the omega equation evidenced in this study should thus have limited consequences on the determination of tracer vertical fluxes, e.g., as performed in Barceló-Llull et al. (2016).

Note that the depth where the mean gradient of τ resides is a parameter that is potentially important to consider. If this mean gradient were located very close below the mixed layer base, the region where the photosynthesis sink term is active would be closer and the relative amplitude of submesoscale vertical motions (compared to that for the mesoscale) would be larger, hence a stronger influence of submesoscale processes.

3) CASE 3

The last limit case is the one where the only possible modification of the tracer concentration attached to water parcels is through (vertical/diapycnal) mixing with surrounding water masses when and where the underlying flow structure produces local shear enhancement. Shear being proportional to spatial derivatives of the velocities, finescale vertical motions can potentially have a major contribution on τ vertical/diapycnal fluxes despite their lesser magnitude relative to those at

mesoscale. Reconstructed vertical velocities have a general bias toward overly weak variance that is particularly pronounced at scales below 30–50 km, depending on the regime. This bias and its impact on the determination of tracer fluxes may be amenable to statistical corrections if omega reconstructions for case 3 type problems turn out to be of importance. The impact of (very) finescale motions on tracer dynamics is presently accounted for by shear-driven mixing parameterizations (Gregg 1989; Polzin et al. 1995; and references therein), which represent the effect of internal gravity wave. There is evidence that different forms of heterogeneities such as submesoscale flows affect internal wave breaking (Whitt and Thomas 2013; Avicola et al. 2007) but the utility of the omega equation to incorporate these effects in our ocean representations is presently unclear.

6. Conclusions

The ever expanding coverage of observational networks supplemented by satellite observations with increasing spatiotemporal resolution like the upcoming SWOT mission (d'Ovidio et al. 2019) was a strong motivation for this work. In the coming years we may be able to estimate vertical fluxes more accurately and at finer scale than ever before. Today, the most common tool to infer vertical velocity is the omega equation. In this paper we explored the ability of the most common configurations adopted to solve this equation to provide information on the vertical circulation at different horizontal length scales. The main novelty of our work is to have used a broad variety of regimes which helped unravel the diversity of reasons underlying the errors in omega reconstructions of w fields, depending on the regional dynamics.

Generally the adiabatic omega reconstruction gives good results for mesoscale vertical circulation, typically for structures with horizontal scales larger than 10 km. These structures tend to be associated with a large part of the w variance, although their relative importance and the precise skills of the omega reconstruction at such scales depend on the local dynamics. Omega reconstruction skills degrade strongly in the submesoscale range. This degradation is manifest both in terms of reduced variance levels and lack of coherence when comparing reconstructed and true vertical velocities. There are two main reasons explaining these findings: the overall weakness of (true) w variance levels below the surface boundary layer in the submesoscale range; the coexistence in that range of several processes contributing to vertical motions, not just frontal turbulence but also inertia-gravity waves which force vertical velocities through terms that are generally not accessible. Overall, the best reconstructions are thus observed in conditions characterized by energetic turbulence and/or weak stratification (such that near-surface frontal processes are felt well into the ocean interior). In particular we observed the best results in regimes for a western boundary (LMX with elevated EKE) and wintertime Porcupine Abyssal Plain (OMS, very weak water column stratification). Conversely the weak vertical flow found in the AZO region where EKE is low and subsurface stratification is relatively strong is poorly captured by the omega approach.

Implications on the role of submesoscale vertical velocities and whether/how to integrate them into our representations of property exchanges in the ocean interior will require further elaboration. Numerous studies have highlighted the strong impact frontal submesoscale turbulence can have on physical and biogeochemical fluxes (Lapeyre et al. 2006; Klein et al. 2008; Thomas and Ferrari 2008; Capet et al. 2008). For a given tracer τ , this is only true to the extent that the mean distribution of τ exhibits vertical contrasts sufficiently close to the surface where frontal activity remains important. How close depends on the background environment and in particular on the upper-ocean stratification. The surface quasigeostrophic (SQG) theory is useful to qualitatively apprehend this dependence (Lapeyre et al. 2006).

In most past studies dealing with vertical velocities a quasigeostrophic version of the omega equation was used, but more recently, several authors have included higher-order terms. For this reason, we also evaluated the possible improvements obtained by using a more elaborate version of the adiabatic equation compared to the simple quasi geostrophic formulation. NG reconstructions that include ageostrophic forcing terms due to flow deformation leads to improved reconstructions. The improvements are substantial only in conditions where near-surface frontal processes (i.e., submesoscales) are important. This finding has important implications in the context of efforts aimed at estimating vertical velocities in the real ocean because the adiabatic QG reconstruction only requires the knowledge of the density field and of the sea surface height (or a reference level of no motion).

The consequences of various numerical choices on the reconstruction skills have been investigated in this work. Choices made for the bottom boundary condition have a major impact. Pushing the boundary condition as deep as possible is the only rule of thumb that can be provided. The relative performance of the Dirichlet versus Neumann boundary conditions cannot be anticipated a priori while the BBC tweak proposed by Rudnick (1996) offers at best marginal improvements. Obviously, omega reconstruction skills will worsen when they are applied to real observations with lesser spatial resolution (for the ocean interior data), synopticity issues, and instrumental errors.

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APPENDIX

Baroclinic Mode Decomposition

To support the results in section 4d, we use a normal mode decomposition to establish possible relations between the

depth of the vertical boundary condition and the error committed when estimated w with ω . The eigenfunctions are solution of the eigenvalues problem:

$$\frac{\partial^2 G_n}{\partial z^2} + \frac{N^2}{c_n^2} G_n = 0, \quad (\text{A1})$$

or

$$\frac{\partial}{\partial z} \left(\frac{1}{N^2} \frac{\partial F_n}{\partial z} \right) + \frac{1}{c_n^2} F_n = 0, \quad (\text{A2})$$

with the boundary conditions:

$$G_n = 0 \quad \text{at} \quad z = 0 \quad \text{and} \quad z = -H, \quad (\text{A3})$$

$$\frac{\partial F_n}{\partial z} = 0 \quad \text{at} \quad z = 0 \quad \text{and} \quad z = -H, \quad (\text{A4})$$

where the vertical velocity modes G_n and the pressure modes F_n are related through the relation,

$$F_n = \frac{\partial G_n}{\partial z}. \quad (\text{A5})$$

Vertical variability can then be projected onto the vertical modes F_n and G_n so that

$$p(x, y, z, t) = \sum_{n=0}^{\infty} \tilde{p}_n(x, y, t) F_n(z), \quad (\text{A6})$$

$$w(x, y, z, t) = \sum_{n=0}^{\infty} \tilde{w}_n(x, y, t) G_n(z), \quad (\text{A7})$$

with \tilde{p}_n and \tilde{w}_n the modal amplitudes of the pressure and vertical velocity, respectively.

To investigate more thoroughly the contribution of the vertical modes to the vertical velocity signal, we expressed the vertical velocity by the projection on the three firsts vertical modes and a residue,

$$w(x, y, z, t) = \sum_{n=1}^N \tilde{w}_n(x, y, t) G_n(z) + \varepsilon. \quad (\text{A8})$$

The term ε was estimated for $N = [1:3]$. The number of profiles where ε/w was lower than 0.5, that is to say that more than half of the w dynamics is explained by the first one, two, or three vertical modes, was estimated (Fig. 12).

Note that in Fig. 11, F_n and G_n are calculated for a mean stratification that is obtained using spatial averaging over the entire region and time averaging over 11 daily model outputs. However, they do not differ substantially from the modes calculated on each point and used to infer modal amplitudes (section 2c).

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