Restratification of Abyssal Mixing Layers by Submesoscale Baroclinic Eddies

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(Manuscript received 20 April 2018, in final form 12 July 2018)

ABSTRACT

For small-scale turbulence to achieve water mass transformation and thus affect the large-scale overturning circulation, it must occur in stratified water. Observations show that abyssal turbulence is strongly enhanced in the bottom few hundred meters in regions with rough topography, and it is thought that these abyssal mixing layers are crucial for closing and shaping the overturning circulation. If it were left unopposed, however, bottom-intensified turbulence would mix away the observed mixing-layer stratification over the course of a few years. It is proposed here that the homogenizing tendency of mixing may be balanced by baroclinic restratification. It is shown that bottom-intensified mixing, if it occurs on a large-scale topographic slope such as a midocean ridge flank, not only erodes stratification but also tilts isopycnals in the bottom few hundred meters. This tilting of isopycnals generates a reservoir of potential energy that can be tapped into by submesoscale baroclinic eddies. The eddies slide dense water under light water and thus restratify the mixing layer, similar to what happens in the surface mixed layer. This restratification is shown to be effective enough to balance the homogenizing tendency of mixing and to maintain the observed mixing-layer stratification. This suggests that submesoscale baroclinic eddies may play a crucial role in providing the stratification mixing can act on, thus allowing sustained water mass transformation. Through their restratification of abyssal mixing layers, submesoscale eddies may therefore directly affect the strength and structure of the abyssal overturning circulation.

1. Introduction

The return of Antarctic Bottom Water from the abyss back to the surface requires water mass transformation by diapycnal mixing (e.g., Lumpkin and Speer 2007; Talley 2013). While small-scale turbulence in the bulk of the abyssal ocean is relatively weak, turbulence levels are elevated where tidal or geostrophic flows pass over a rugged seafloor (e.g., Polzin et al. 1997; Ledwell et al. 2000; St. Laurent et al. 2012; Waterhouse et al. 2014). Internal waves are excited, become nonlinear, and break, typically within a few hundred meters of the ocean bottom (e.g., Legg and Huijts 2006; Garrett and Kunze 2007; Muller and Bühler 2009; Nikurashin and Ferrari 2010; Nikurashin and Legg 2011). This topographically induced turbulence can achieve substantial water mass transformation and thus help close the large-scale overturning circulation of the abyssal ocean, producing a few tens of Sverdrups (1 Sv = 10^6 m^3 s^-1) of both diapycnal up- and downwelling (e.g., St. Laurent et al. 2001; Jayne 2009; de Lavierne et al. 2016; Ferrari et al. 2016; McDougall and Ferrari 2017; de Lavierne et al. 2017; Holmes et al. 2018; Callies and Ferrari 2018b).

For the topographically generated turbulence to have an effect on the overturning, however, the mixing must occur in stratified water. Otherwise, the turbulence would mix mixed fluid, and there would be no water mass transformation (cf. Armi 1978, 1979; Garrett 1979, 1990). Observations do suggest that abyssal mixing layers are stratified: hydrographic sections show isopycnals that only gently bend down within the bottom 500–1000 m, for example, on the flanks of the Mid-Atlantic Ridge (Fig. 1). Vertical profiles show a stratification that is reduced in the bottom few hundred meters by only about a factor of 3 (Fig. 2a, blue curve). There must be a process that maintains the stratification in abyssal mixing layers against the homogenizing effect of the turbulence.

When turbulence occurs on large-scale topographic slopes that intersect the ambient stratification, like on the flanks of the Mid-Atlantic Ridge, cross-slope mean flows develop in response to the mixing (e.g., Phillips 1970; Wunsch 1970; Garrett et al. 1993). These mean flows have a restratifying tendency: they advect dense water up the slope at the bottom and light water down the slope above. As in Munk’s (1966) balance, the turbulent buoyancy flux divergence is balanced by mean advection, except that the mean flow is in the cross-slope direction instead of the vertical. A dipolar cross-slope
flow emerges because buoyancy fluxes have a pattern of divergence overlying convergence when mixing is bottom intensified.

While these cross-slope mean flows tend to increase the stratification in abyssal mixing layers, they are unlikely to be strong enough to maintain the observed stratification. As shown in section 3, the one-dimensional theory that describes this mechanism of mean-flow restratification, given realistically bottom-intensified mixing, produces mixing layers that are much more weakly stratified than what is observed (Fig. 2a, orange curve; Figs. 3a,b). One-dimensional theory is missing a crucial piece of physics.

Section 4 demonstrates that abyssal mixing-layer stratification can instead effectively be maintained by submesoscale baroclinic eddies. In Wenegrat et al. (2018), we showed that the steady state produced by one-dimensional theory is linearly unstable to baroclinic perturbations. The steady state’s weak stratification implies strongly sloping isopycnals, a configuration that is associated with a large reservoir of potential energy available for release (Fig. 3b). Baroclinic instability taps into this potential energy reservoir by sliding dense water under light water in the cross-slope direction at an instability scale of order 10 km and a growth rate of order (10 days)$^{-1}$. The dynamics are similar to those of submesoscale baroclinic instabilities in the surface mixed layer (e.g., Boccaletti et al. 2007; Fox-Kemper et al. 2008; Callies et al. 2016).

Using three-dimensional calculations, I show in section 4 that this submesoscale baroclinic instability disrupts the one-dimensional balance of mixing layers on slopes by producing finite-amplitude eddies and substantial restratification. Over the course of a few months, baroclinic eddies eliminate the unrealistically weak mixing-layer stratification of the one-dimensional steady state (Fig. 2a, green curve; Fig. 3c). A new balance is achieved between the homogenizing effect of mixing and the restratifying effect of baroclinic eddies, in which the observed stratification is maintained (Fig. 2a, red curve; Fig. 3d). The restratification by submesoscale baroclinic eddies is much more powerful than that by mean flows and may be crucial in enabling topographically generated small-scale turbulence to mix stratified fluid, to incur substantial water mass transformation, and thus to affect the abyssal overturning circulation.

2. Observational context

To set the stage for the physical arguments presented in subsequent sections, I begin by reviewing observations from the microstructure component of the Brazil Basin Tracer Release Experiment (Polzin et al. 1997; St. Laurent et al. 2001). I focus on the profiles east of 27°W that are on the western flank of the Mid-Atlantic Ridge (Fig. 1a), where the bottom is rough and small-scale turbulence is enhanced. The bottom-intensified turbulence there is similar in magnitude and vertical structure to the turbulence observed over a rough seafloor elsewhere in the World Ocean (Waterhouse et al. 2014).

From the 1996 and 1997 cruises, the University of California, San Diego (UCSD), microstructure database (microstructure.ucsd.edu) holds 126 profiles for which a valid bottom depth is reported. For each profile, I calculate the stratification using the TEOS-10 seawater toolbox (McDougall and Barker 2011), and I average the stratification and the reported dissipation rates over 25-m bins of height above the bottom. I dismiss data within 100 m of the bottom because few profiles reach this deep, and averages are unreliable. I average across profiles and obtain 95% confidence intervals using the bootstrap percentile method (e.g., Efron and Hastie 2016). This yields a crude description of the bottom 2000 m on the western flank of the Mid-Atlantic Ridge, averaged over horizontal and temporal variations. This crudeness is acceptable here because the goal is not to match the observations exactly with the dynamical models discussed below, but to understand whether the proposed
dynamics can predict the order of magnitudes of the observed profiles. See St. Laurent et al. (2001) for a comprehensive description of the data and their variability.

In a 1000-m-thick layer overlying the bottom, the stratification and dissipation rates show a notable change from relatively constant far-field values above (Fig. 2a, blue curve; Fig. 2b). The stratification decreases from a value of about $1.6 \times 10^{-6}$ s$^{-2}$ at 1000–2000 m to about $5 \times 10^{-7}$ s$^{-2}$ at 100 m (Fig. 2a, blue curve). The kinetic energy dissipation rate increases sharply from a value of about $4 \times 10^{-10}$ m$^2$ s$^{-3}$ at 1000–2000 m to about $3 \times 10^{-9}$ m$^2$ s$^{-3}$ at 100 m (Fig. 2b).

The observed dissipation rate $\epsilon$ is commonly related to the vertical buoyancy flux by $F' = -\Gamma \epsilon$, where $\Gamma$ is an empirical nondimensional flux ratio typically assumed constant and equal to 0.2 (Osborn 1980). The buoyancy flux thus inferred from the dissipation rates increases in magnitude toward the bottom. Together with the decrease in stratification $B_z$, this gives rise to a sharp increase in the turbulent diffusivity $\kappa = -F'/B_z$ (Fig. 2c). Fitting the functional form

$$\kappa = \kappa_0 + \kappa_1 e^{-z/h}$$

(1)

yields a far-field diffusivity of $\kappa_0 = 5.2 \times 10^{-5}$ m$^2$ s$^{-1}$, a bottom enhancement of $\kappa_1 = 1.8 \times 10^{-3}$ m$^2$ s$^{-1}$, and a decay scale of $h = 230$ m. I will use these values, together with the far-field buoyancy frequency $N = 1.3 \times 10^{-3}$ s$^{-1}$, throughout the paper (Table 1).

3. One-dimensional dynamics: Mean-flow restratification

In this section, I demonstrate that the observed stratification in the mixing layer cannot be reproduced by one-dimensional dynamics. I consider the dynamics of a rotating fluid adjacent to a planar slope that is inclined at an angle $\theta$ with respect to the horizontal, with the flow depending on the slope-normal coordinate only (e.g., Wunsch 1970; Phillips 1970; Thorpe 1987; Garrett et al. 1993). I assume the fluid to have a constant far-field buoyancy frequency $N$, subject it to the bottom-intensified mixing inferred from the Brazil Basin observations (Fig. 2c), and consider the local response. If

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1 Variations in the flux ratio $\Gamma$ with stratification and dissipation rate have been detected in numerical simulations and laboratory experiments (e.g., Gregg et al. 2018). These effects are too small to affect the qualitative results of this paper and are neglected for simplicity.
mean-flow restratification were dominant, such a one-dimensional model would be able to predict the observed stratification change over the bottom few hundred meters (Fig. 2a, blue curve). But while the thickness of the layer over which the stratification is modified from its far-field value is accurately predicted, the mixing-layer stratification is predicted to be much weaker than what is observed (Fig. 2a, orange curve).

A number of solutions to the one-dimensional dynamics on slopes are available, but no full steady solution has been derived for the realistic case at hand. There is a well-known analytical solution for constant mixing coefficients (Wunsch 1970; Thorpe 1987), but it is quite different in character from the case with bottom-intensified mixing. Thompson and Johnson (1996) considered bottom-intensified mixing coefficients, but their transient solution did not reach steady state. Garrett (1991) derived an approximate analytical solution for the steady state, but his solution is valid only outside an inner layer that carries the upwelling part of the dipolar mean flow. I here extend Garrett’s solution into the inner layer and thus obtain a uniformly valid approximate analytical solution.2,3

The dynamics are best described in a coordinate system aligned with the slope, with $x$ being the cross-slope coordinate (increasing upslope), $y$ being the along-slope coordinate, and $z$ being the slope-normal coordinate.4 The local dynamics can be isolated by splitting the buoyancy field into a far-field stratification and an anomaly, $B = N^2(z \cos \theta + x \sin \theta + b)$. Assuming that neither the buoyancy anomaly $b$ nor the flow varies in the plane of the slope, and that the flow is in steady state, reduces the Boussinesq equations to

\[
A = f \kappa, \quad \cot \theta = 2 \times 10^{-3}
\]

\[
B = N^2(z \cos \theta + x \sin \theta) + b
\]

TABLE 1. Parameters appropriate for the western flank of the Mid-Atlantic Ridge in the Brazil Basin. These parameters are used in all calculations presented in this paper.

<table>
<thead>
<tr>
<th>Parameter</th>
<th>Value</th>
</tr>
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<tbody>
<tr>
<td>Inertial frequency $f$</td>
<td>$-5.5 \times 10^{-5}$ s$^{-1}$</td>
</tr>
<tr>
<td>Far-field buoyancy frequency $N$</td>
<td>$1.3 \times 10^{-3}$ s$^{-1}$</td>
</tr>
<tr>
<td>Slope angle $\theta$</td>
<td>$2 \times 10^{-3}$</td>
</tr>
<tr>
<td>Far-field diffusivity $\kappa_0$</td>
<td>$5.2 \times 10^{-5}$ m$^2$ s$^{-1}$</td>
</tr>
<tr>
<td>Bottom-enhancement of diffusivity $\kappa_1$</td>
<td>$1.8 \times 10^{-3}$ m$^2$ s$^{-1}$</td>
</tr>
<tr>
<td>Decay scale of diffusivity $h$</td>
<td>$230$ m</td>
</tr>
<tr>
<td>Prandtl number $\sigma$</td>
<td>$1$</td>
</tr>
<tr>
<td>Hyperviscosity/hyperdiffusivity $\mu$</td>
<td>$6.25 \times 10^{4}$ m$^4$ s$^{-1}$</td>
</tr>
</tbody>
</table>

2 Garrett (1990) presented numerical solutions to the full bottom-intensified case, but his setting $A = f \kappa, \cot \theta$ (his notation) with a variable $\kappa$, was not justified and renders the solutions invalid.

3 In Callies and Ferrari (2018b), we obtained steady solutions for bottom-intensified mixing, but these solutions depended on our choice of Rayleigh drag as a parameterization for friction. These drag-controlled boundary layers were nothing more than stand-ins for the insufficiently understood dynamics of these layers, in order to study what implications the up- and downwelling in these layers have for the large-scale circulation. Here, I resort to a more conventional diffusive parameterization of turbulent momentum transport. While it is not clear whether the diffusive closure is appropriate for turbulence produced by internal wave breaking, its form as a flux divergence at least excludes momentum sinks in the interior of the fluid.

4 For the small slope angles typical of the abyssal ocean, the slope-normal coordinate nearly coincides with the vertical, so there is no need to distinguish between the two coordinates in the analysis of observations (section 2).
\[-f \nu \cos \theta = b \sin \theta + (\nu u_z) z, \quad (2)\]
\[fu \cos \theta = (\nu u_z) z, \quad (3)\]
\[u N^2 \sin \theta = [\kappa (N^2 \cos \theta + b_z)] z, \quad (4)\]

where \(u\) is the cross-slope velocity, \(v\) is the along-slope velocity, and \(f\) is the inertial frequency (assumed constant). Turbulent momentum transfer is parameterized by a diffusive closure, and the turbulent viscosity \(\nu = \sigma \kappa\) is assumed to have the same vertical profile as the turbulent diffusivity (i.e., the turbulent Prandtl number \(\sigma\) is assumed constant).

The slope-normal velocity vanishes, as required by continuity, and pressure can be inferred from the slope-normal momentum balance. Boundary conditions are that the flow and buoyancy flux vanish at the slope (i.e., \(u = 0, \nu = 0,\) and \(N^2 \cos \theta + b_z = 0\) at \(z = 0\)); decay conditions \(u_z \to 0, \nu_z \to 0,\) and \(b_z \to 0\) are enforced for \(z \to \infty\). As detailed by Garrett (1991), the system in (2)–(4) can be reduced to

\[
(\nu \psi_{zz})_{zz} + \left(\frac{f^2 \cos^2 \theta}{\nu} + \frac{N^2 \sin^2 \theta}{\kappa}\right) \psi = \frac{N^2}{\nu} \sin \theta \psi + \frac{f^2 \cos^2 \theta}{\nu} - \kappa \psi + \nu \psi_{zz}, \quad (5)
\]

where the streamfunction \(\psi\) is defined by \(\psi_z = u\). The boundary conditions become \(\psi = 0\) and \(\psi_z = 0\) at \(z = 0\), and \(\psi \to \kappa \psi + \psi_{zz}\) as \(z \to \infty\) (Thorpe 1987).

This system is easily solved numerically by limiting the domain to a large but finite height \(H = 4000\) m, projecting the original set in (2)–(4) onto 2048 Chebyshev polynomials, and solving the resulting linear system (Burns et al. 2016). The solutions are insensitive to imposing the upper-boundary conditions at a finite height as long as the domain is tall enough to fit the layer over which the background stratification is significantly modified.

An illuminating approximate analytical solution can be found if the decay scale of mixing \(h\) is much larger than the thickness of a frictional inner layer (cf. Garrett 1991; Garrett et al. 1993). The thickness of the inner layer is given by (cf. Wunsch 1970; Thorpe 1987)

\[q^4 = \frac{1}{4(\nu_0 + \nu_1)} \left(\frac{f^2}{\nu_0 + \nu_1} + \frac{N^2 \sin^2 \theta}{\kappa_0 + \kappa_1}\right) = \frac{f^2 \cos^2 \theta (1 + S \sigma)}{4(\nu_0 + \nu_1)^2}, \quad (6)\]

where \(S = N^2 \tan^2 \theta / f^2\) is the slope Burger number, \(\nu_0 = \sigma \kappa_0,\) and \(\nu_1 = \sigma \kappa_1.\) The scale \(q^{-1}\) is the boundary layer thickness that would emerge if the mixing coefficients were constant and equal to their bottom values. This scale is equal to the thickness of a conventional Ekman layer modified by a factor \((1 + S \sigma)^{-1/4}\), which arises from a combination of the effects of the slope and stratification.

For parameters typical of the abyssal ocean, the scale \(q^{-1}\) is much smaller than the decay scale \(h.\) For the western flank of the Mid-Atlantic Ridge sampled by the Brazil Basin Tracer Release Experiment, \(f = -5.5 \times 10^{-5}\) s\(^{-1}\) for a latitude of 22°S, \(N = 1.3 \times 10^{-3}\) s\(^{-1}\), and the mean slope is approximately \(\theta = 2 \times 10^{-3}\) (Fig. 1, Table 1), which gives a slope Burger number of \(S = 2.2 \times 10^{-3}\). The Ekman layer thickness is thus only marginally modified if the Prandtl number is not very large. For \(\sigma = 1\) and mixing coefficients obtained from the observations (Fig. 2c, Table 1), the inner-layer scale is \(q^{-1} = 8.3\) m. With \(h = 230\) m, it is thus safe to assume that \(qh \gg 1\) and to apply boundary layer theory (e.g., Bender and Orszag 1999).

Outside the inner layer, the solution varies on a scale \(z \sim h,\) and the first term in (5) can be neglected (Garrett 1991). A solution valid in the outer layer is thus

\[\psi^{(0)} = \cot \theta \left(\kappa_0 + \kappa_1 e^{-zh} \frac{S \sigma}{1 + S \sigma}\right). \quad (7)\]

The outer-layer solution does not satisfy the boundary condition at \(z = 0,\) however, so an inner layer is present there. The inner-layer adjustment \(\psi^{(1)}\) is determined by the homogeneous version of (5), but since \(\psi^{(1)}\) varies on a scale \(z \sim q^{-1}\), the first term in (5) is now of leading order. Variations in mixing coefficients, on the other hand, are slight over the inner-layer scale and can be neglected, such that the coefficients become constant:

\[\psi^{(1)}_{zz} + 4q^4 \psi^{(1)} = 0. \quad (8)\]

The inner-layer correction thus has the same structure as the solution for constant mixing coefficients (Wunsch 1970; Thorpe 1987), but the boundary conditions must be matched to the outer-layer solution. Enforcing \(\psi^{(0)} + \psi^{(1)} = 0\) and \(\psi^{(1)}_{zz} = 0\) at \(z = 0\) (neglecting much smaller outer-layer contributions to \(\psi^{(1)}\)), as well as decay conditions as \(z \to \infty\), yields

\[\psi^{(1)} = -\left(\kappa_0 + \kappa_1 \frac{S \sigma}{1 + S \sigma}\right) \cot \theta e^{-q^{-1} z} (\cos qz + \sin qz). \quad (9)\]

The sum of (7) and (9) is a uniformly valid approximate solution to (5). Figure 4 illustrates this solution and compares it to the numerical solution. The approximation is excellent.

The analytical solution shows that the cross-slope transport integrated over the outer layer is \(-\kappa_1 \cot \theta \sigma \sigma (1 + S \sigma)\) and that the inner-layer transport

\[\frac{S \sigma}{1 + S \sigma}\]

Garrett (1991) noted that the outer-layer solution (7) allows the Prandtl number \(\sigma\) to be a function of height. If \(\sigma\) variations are much slower than \(q^{-1}\), the boundary layer approach still works, and \(\sigma\) in (9) is simply replaced by its value at \(z = 0,\)
is $\kappa_0 \cot \theta + \kappa_1 \cot \theta S \sigma/(1 + S \sigma)$. The solution thus exhibits compensating cross-slope transports of magnitude $\kappa_1 \cot \theta S \sigma/(1 + S \sigma)$ in the inner and outer layers and a net upslope transport $\kappa_0 \cot \theta$ that is carried by the inner layer (cf. Thorpe 1987). The compensated part of the transport is the anticipated dipole of upwelling in the inner layer and downwelling in the outer layer that is due to the bottom intensification of mixing. Its magnitude, however, is reduced by a factor of $S \sigma/(1 + S \sigma)$ compared to what it would be if the stratification in the outer layer was equal to the background. This factor is small for abyssal values if the Prandtl number is not large.

The stratification (in the slope-normal direction) follows directly from (7) and (9) through $\psi N^2 \sin \theta = \kappa (N^2 \cos \theta + b_z)$ and is given by (Figs. 4b,f)

\[
N^2 \cos \theta + b_z = N^2 \cos \theta \left[ \frac{\kappa_0}{\kappa_0 + \kappa_1 e^{-\gamma h}} + \frac{\kappa_1 e^{-\gamma h} S \sigma}{\kappa_0 + \kappa_1 e^{-\gamma h}} \frac{1 + S \sigma}{1 + S \sigma} - \left( \frac{\kappa_0}{\kappa_0 + \kappa_1} + \frac{\kappa_1}{\kappa_0 + \kappa_1} \frac{S \sigma}{1 + S \sigma} \right) e^{-\gamma z} (\cos qz + \sin qz) \right].
\]

(10)

The stratification is reduced in the outer layer up to $z \sim h \log(\kappa_1/\kappa_0)$, the level at which the diffusivity has decayed to approximately its background value. This reduction in the outer layer is strong if $\kappa_0 \ll \kappa_1$ and

FIG. 4. One-dimensional steady state in the (top) outer layer and far field and (bottom) inner layer. Shown are the (a),(e) streamfunction, (b),(f) stratification, (c),(g) cross-slope flow, and (d),(h) along-slope flow as produced by the analytical solutions for the outer layer and far field (blue) and by the analytical solution including the inner-layer contribution (green). The approximate analytical solution matches the numerical solution (orange). Note the different scales of the abscissae in the upper and lower panels; the gray boxes in the upper panels show the plot range that is expanded in the lower layers, with the exception of (g), which extends to much larger positive values than indicated in (c).
\[
\sigma r/(1 + \sigma r) \ll 1.
\]
In the inner layer, the stratification decays from its reduced value at the inner edge of the outer layer to zero, as required to satisfy the boundary condition. The change in stratification in the inner layer has the same magnitude as the stratification value at the inner edge of the outer layer. There is a slight enhancement in stratification at the outer edge of the inner layer (around \(z = 25\) m in Fig. 4f). For \(\kappa_0 \ll \kappa_1\) and \(\sigma r/(1 + \sigma r) \ll 1\), however, stratification changes over the inner layer are small compared to those in the outer layer (Figs. 4b,f), and the enhancement at the upper edge of the inner layer is too slight to reverse the horizontal buoyancy gradient \(-b_z \sin \theta\) (cf. Phillips et al. 1986; Garrett et al. 1993).

The cross-slope flow implied by (7) and (9) is (Figs. 4c,g)

\[
u = -\frac{k_1 \cot \theta e^{-z/h} \sigma r}{h(1 + \sigma r)} + 2g \cot \theta \left(\frac{\kappa_0 + \kappa_1}{1 + \sigma r}e^{-qz}\sin qz\right),
\]  

This flow consists of a dipole of downslope flow in the outer layer and compensating upslope flow in the inner layer, plus an additional upslope flow in the inner layer that is proportional to \(\kappa_0\). This uncompensated component of the inner-layer flow has the same functional form as if one had set \(\kappa = \kappa_0\) uniformly (cf. Thorpe 1987), except that \(q\) is now based on the enhanced bottom-mixing coefficients. The downslope flow in the outer layer decays with the same scale \(h\) as the diffusivity, and its magnitude is proportional to \(\kappa_1 \cot \theta \sigma r/(1 + \sigma r)\).

The along-slope shear implied by (7) and (9) is

\[
u_z = \frac{f \cot \theta \cos \theta}{\sigma} \left[\frac{k_1 e^{-z/h} \sigma r}{\kappa_0 + \kappa_1 e^{-z/h} (1 + \sigma r)} - \left(\frac{\kappa_0}{\kappa_0 + \kappa_1} + \frac{\kappa_1}{\kappa_0 + \kappa_1 + 1/\sigma r}e^{-qz}(\cos qz + \sin qz)\right)\right],
\]  

which can easily be integrated upward from \(v = 0\) at \(z = 0\) to infer the flow (Figs. 4d,h). There is significant shear in the outer layer, in the direction of Kelvin wave propagation, up to the height \(z \sim h \log(\kappa_1/\kappa_0)\). This shear is in thermal wind balance with the horizontal buoyancy gradient \(-b_z \sin \theta\) associated with the sloping isopycnals.

It should be noted that the solution derived above remains valid as \(\kappa_0 \to 0\), but it loses its confinement to the bottom: the reduced stratification and the along-slope shear extend to \(z \to \infty\) (cf. Garrett 1990). This emphasizes that it is not the decay scale of mixing \(h\) that sets the thickness of the layer over which the hydrography is modified, but the height at which the mixing coefficients stop changing, which occurs at \(\kappa_0 \sim \kappa_1 e^{-z/h}\) or \(z \sim h \log(\kappa_1/\kappa_0)\).

For Brazil Basin parameters, the one-dimensional solution produces a stratification that is much weaker than what is observed (Fig. 2a, blue and orange curves). The thickness of the outer layer \(h \log(\kappa_1/\kappa_0) = 800\) m roughly matches the observed thickness of the layer with reduced stratification, but the reduction is much too severe. This shows that the cross-slope mean flow in (11) produced by one-dimensional dynamics, while displaying the anticipated dipolar structure, is much too weak to maintain realistic stratification. Instead of generating a strong dipole of cross-slope mean flow that balances the large diabatic buoyancy flux divergence and convergence produced by turbulence in a stratified fluid, one-dimensional dynamics predict a near homogenization of the mixing layer, in which the little water mass transformation that remains is balanced by weak cross-slope mean flows. I will return to this point in section 5.

4. Three-dimensional dynamics: Baroclinic restratification

The inability of the one-dimensional dynamics in (2)–(4) to reproduce the observed mixing-layer stratification (Fig. 2a, blue and orange curves) suggests that a crucial piece of physics is missing. In this section, I demonstrate that the one-dimensional balances are upset by a baroclinic instability and the resulting nonlinear baroclinic eddies. I consider two experiments in the same setup on a planar slope as in section 3, but I now allow three-dimensional and fully nonlinear flow. Experiment 1 starts from the one-dimensional steady state. This steady state is linearly unstable to baroclinic perturbations. The unstable modes grow, attain finite amplitude, and rapidly restratify the mixing layer, suggesting that we should not expect to see the one-dimensional steady state in observations. While experiment 1 clearly demonstrates the breakdown of the one-dimensional balances, it is a bit academic because the initial state has unrealistically weak stratification in the mixing layer, so the baroclinic restratification is unrealistically vigorous. But restratifying baroclinic eddies also emerge in the more realistic experiment 2, which starts from a stratification close to what is observed in the Brazil Basin. Baroclinic eddies again develop, and their restratifying tendency effectively balances the homogenizing tendency of mixing.

a. Baroclinic instability and restoration of stratification

In Wenegrat et al. (2018), we showed that the one-dimensional steady state obtained in section 3 is linearly
unstable to baroclinic perturbations for parameters typical of abyssal mixing layers ($k_1 \gg k_0$ and $qh \gg 1$). The reduced stratification of the outer layer is associated with strongly sloping isopycns and along-slope shear (Fig. 2d, orange curve; Fig. 3b), a configuration that allows counterpropagating Rossby waves to reinforce each other and grow. For completeness, I repeat the linear stability calculation here, using a domain height of $H = 2500$ m, projecting the slope-normal structure onto 64 Chebyshev modes, and solving the resulting truncated eigenvalue problem (Burns et al. 2016). For the Brazil Basin parameters (Table 1), instability occurs over a range of along-slope wavenumbers, with a maximum growth rate of (9 days)$^{-1}$ located at the along-slope wavenumber $l = 5.2 \times 10^{-4}$ m$^{-1} = 2\pi/(12$ km) (Fig. 5a). This most unstable perturbation grows through the release of available potential energy by buoyancy production in the bottom 400 m, while the bottom Ekman layer exerts a significant drag (Fig. 5b). The most unstable mode displays the classic tilt of the cross-slope velocity into the shear (Fig. 5c; e.g., Pedlosky 1987; Vallis 2017) and the in-phase relationship between slope-normal velocity and buoyancy that allows potential energy release (Figs. 5d,e). The topographic slope does not stabilize the flow because it is much smaller than the isopycnic slope (Wenegrat et al. 2018; Fig. 3b).

To assess the restratification that is induced by the baroclinic instability and the finite-amplitude eddies that it produces, I now consider a fully nonlinear three-dimensional numerical simulation of an abyssal mixing layer initialized with the one-dimensional steady state, using the same parameters as above (Table 1). Now allowing the flow to be time dependent and to vary in the plane of the slope (i.e., in $x$ and $y$), the Boussinesq equations in the slope-aligned coordinate system and with the hydrostatic approximation read

$$u_t + u \cdot \nabla u - f v \cos \theta = -p_z + b \sin \theta$$

$$v_t + u \cdot \nabla v + f u \cos \theta = -p_z + (nu)_z + Dv,$$  

$$p_z = b \cos \theta,$$  

$$\nabla \cdot u = 0,$$  

$$b_t + u \cdot \nabla b + uN^2 \sin \theta + wN^2 \cos \theta = [\kappa(N^2 \cos \theta + b_z)]_z + Db.$$  

The hydrostatic approximation is justified if the aspect ratio of the flow $\delta$ is small (i.e., $\delta^2 \ll 1$) and if $\delta \tan \theta \ll 1$, both of which are satisfied for the flows of interest here. In the equations above, $u$ denotes the three-dimensional velocity vector, $w$ is the slope-normal velocity component, $p$ is the pressure divided by a reference density, and $D = -\mu(\partial^2 + 2\alpha_2^2 + \delta^2)$ is a hyperdiffusion operator that is added for numerical stability. I find the hyperviscosity/hyperdiffusivity $\mu = 6.25 \times 10^4$ m$^2$ s$^{-1}$ to be large enough to suppress gridscale oscillations, albeit small enough to render negligible the energy loss due to hyperviscosity. The vertical boundary conditions are the same as for the one-dimensional equations (i.e., $u = 0$, $v = 0$, and $N^2 \cos \theta + b_z = 0$ at $z = 0$ and $u_z = 0$, $v_z = 0$, and $b_z = 0$ at $z = H$), augmented by normal-flow conditions (i.e., $w = 0$ at $z = 0$ and $z = H$), where $H = 2500$ m is again large enough to have a negligible effect on the flow in the mixing layer. In the plane of the slope, the domain is square with a side length of $L = 256$ km, and periodic boundary conditions are imposed on velocities and buoyancy perturbations. This setup is similar to the widely used frontal-zone setup (e.g., Bretherton and Karweit 1975; Salmon 1980; Haidvogel and Held 1980; Taylor and Ferrari 2010; Callies and Ferrari 2018a) because it imposes a fixed domain-average cross-slope buoyancy gradient $N^2 \sin \theta$. Note that in this setup on a slope, however, the mean buoyancy gradient in the true horizontal can and does change during restratification. Fourier bases with 256 modes each are used in $x$ and $y$, the $z$ direction is again projected onto 64 Chebyshev modes, and the system is fully dealiased. Time stepping is performed with a semi-implicit third-order four-stage Runge–Kutta scheme and a time step of 20 min (Burns et al. 2016).

I initialize the flow with the one-dimensional steady state (section 3), with small random perturbations (normally distributed with standard deviation $2 \times 10^{-7}$ m s$^{-2}$) added to the buoyancy field to initiate the instability. As predicted by the linear calculation (Fig. 5a), the dominant baroclinic mode that emerges after a few $e$-folding times has an along-slope wavelength of about 12 km. Around $t = 50$ days, the instability has grown to an amplitude at which nonlinear terms become important. Secondary instabilities occur, and the flow becomes fully turbulent (Fig. 6). The dominant eddies grow in size because nonlinear-scale interactions preferentially transfer energy to larger scales and because restratification stabilizes the smaller scales. The eddies retain a scale somewhat smaller than the size of the domain, suggesting they are not constrained by the finite domain size.

The evolution of the flow can be tracked in the eddy kinetic energy budget:

$$K_e = \langle u^2 \rangle \sin \theta + \langle w^2 \rangle \cos \theta - \langle u \nabla u + v \nabla v \rangle - \langle (u^2 + v^2) \rangle + \langle u D u + v D v \rangle.$$  

(18)
The eddy kinetic energy is

\[ K = \frac{1}{2} (\overline{u'^2} + \overline{v'^2}). \]  

(19)

The overline denotes an average in the plane of the slope, \( A' = A - A \) denotes perturbations in the plane of the slope, and the angle brackets denote a slope-normal integral over the domain. Eddy kinetic energy can be created by buoyancy production (first term on the right) and shear production (second term on the right), and it is lost to dissipation (third term on the right) and to hyperviscosity (fourth term on the right). The buoyancy production has two terms because the coordinate system is rotated with respect to gravity.

As expected from the linear stability calculation (Fig. 5b), eddy kinetic energy is created by a slight dominance of buoyancy production in the outer mixing layer over dissipation in the bottom Ekman layer (Figs. 7a,b). Both buoyancy production and dissipation keep increasing in the nonlinear phase of the evolution, but they level out and slightly decrease after \( t = 150 \) days. The eddy kinetic energy itself also levels out and starts decreasing shortly after that. Neither shear production nor hyperviscosity plays a significant role in the energetics.

Once the baroclinic perturbations have reached significant amplitude around \( t = 75 \) days, they start modifying the mean state (Figs. 7c,d). The sliding of dense water under light water that energizes the instability also restratifies the mean state (Figs. 3c and 7c). Once the perturbations have reached significant amplitude, it takes just 50 days for the eddies to substantially enhance the stratification in the mixing layer. The restratification is largest just above the bottom Ekman layer, but the entire outer layer restratifies, reaching stratification values of the same order as in the far field around \( t = 150 \) days.

The restratification is attended by a reduction in the along-slope shear (Fig. 2d, green curve; Fig. 7d). This is to be expected because throughout the evolution of the flow the along-slope shear remains in approximate thermal wind balance \( f \bar{v} \cos \theta = -\bar{\nu}_z \sin \theta \), where the
right side is the mean buoyancy gradient in the true horizontal. Initially, the perturbation $b_z$ is strongly negative and opposes the background contribution $N^2 \cos \theta$, so $\tau_z$ is strongly negative as well ($f < 0$). As the mixing layer restratifies, $b_z$ and thus $\tau_z$ become less negative. The reduction in shear is induced by the familiar vertical momentum flux effected by baroclinic eddies (e.g., Rhines and Young 1982).

b. Maintenance of stratification by baroclinic eddies

To show that baroclinic eddies emerge and enter the leading-order balances also in an abyssal mixing layer that is more realistically stratified, I consider a second experiment that is initialized with the observed stratification. To obtain smooth initial conditions, I fit the functional form of the outer-layer stratification profile of the one-dimensional solution—that is, the first two terms on the right side of (10)—to the observed stratification, considering $S\sigma/(1 + S\sigma)$ a free parameter. The best fit is obtained with $S\sigma/(1 + S\sigma) = 0.35$ (Fig. 2a, black curve; Fig. 3a). I assume the along-slope flow to be in thermal wind balance and to vanish at the bottom (Fig. 2d, black curve), and I set the cross-slope flow to zero everywhere. Small random noise is again added to the initial buoyancy field, and the setup, discretization, and parameters are the same as in experiment 1 above, except that now a larger time step of 1 h is used.

To isolate the effect of baroclinic eddies, I compare this three-dimensional experiment 2 to the one-dimensional evolution obtained from (2)–(4), with the tendency terms reinstated:
In these one-dimensional dynamics, the stratification in the mixing layer quickly erodes (Fig. 8). After 6000 days, the bottom 500 m have become nearly unstratified. While 6000 days is too short for this solution to reach the steady state of the previous section, which requires diffusion into the far field (cf. MacCready and Rhines 1991; Garrett et al. 1993), it is long enough for the vigorous near-bottom mixing to effectively homogenize the fluid in the bottom few hundred meters. The diffusive time scale based on the bottom diffusivity and a vertical distance of 500 m is on the order of a year (cf. Thompson and Johnson 1996).

The evolution is markedly different in the three-dimensional case (Fig. 9). The erosion of stratification lasts only 500 days, after which baroclinic eddies have reached large-enough amplitude to restratify the mixing layer (Fig. 9c). Like in experiment 1, the eddies are energized by a slight imbalance between buoyancy production and dissipation in the Ekman layer (Figs. 9a,b). Both the eddy kinetic energy and the terms in the energy budget are an order of magnitude smaller than in the experiment started from the one-dimensional steady state, but the baroclinic restratification is still strong enough to maintain the mixing-layer stratification close to the observed value (Fig. 9c). The average stratification over the time interval 3000–6000 days is in remarkable agreement with the observations (Fig. 2a, red curve; Figs. 3a,d).

It is interesting to note that the stratification observed in Brazil Basin mixing layers corresponds to isopycnal slopes that are of the same order as the large-scale topographic slope (Figs. 1b and 3a). This may not be a coincidence. Baroclinic eddies cause strong restratification of abyssal mixing layers, but the instability that gives rise to the eddies can be stabilized by the topographic slope if the topographic and isopycnal slopes are comparable (e.g., Blumsack and Gierasch 1972; Wenegrat et al. 2018). The observed hydrography may thus be the result of baroclinic adjustment to a marginally stable state. More simulations and analysis of observations, however, are needed to test this conjecture.

5. Implications for the abyssal overturning circulation

The presence of strong stratification in abyssal mixing layers is crucial for bottom-enhanced turbulence to achieve water mass transformation and thus to affect the overturning circulation. It is the stratification that allows the turbulence to generate large diabatic buoyancy fluxes, whose convergence and divergence causes the dipole of up- and downwelling that has emerged as a key feature of the diabatic circulation of the abyssal ocean (de Lavergne et al. 2016; Ferrari et al. 2016; McDougall and Ferrari 2017; de Lavergne et al. 2017; Holmes et al. 2018; Callies and Ferrari 2018b). The magnitude of this dipole is substantial: an estimate based on the observed stratification and a semiempirical mixing map yields

\[
\begin{align*}
    u_t & - f v \cos \theta = b \sin \theta + (nu_z)_z, \\
    v_t & + f u \cos \theta = (nu_z)_z, \\
    b_t & + u N^2 \sin \theta = [\kappa (N^2 \cos \theta + b_z)]_z.
\end{align*}
\]

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40 Sv of upwelling and 30 Sv of downwelling across the isopycnal bounding Antarctic Bottom Water (Ferrari et al. 2016).

A similar estimate for the dipole strength can be made by extrapolating the Brazil Basin observations: assuming a planar slope, the diabatic transports can be inferred by taking the differences of $\psi N^2 \sin \theta = -F^2 = \Gamma \delta$ across the up- and downwelling layers (cf. Garrett 1991; McDougall and Ferrari 2017). The difference between the far-field dissipation rate and the bottommost observed value gives an estimate for the downwelling over the observed layer; the bottommost observed value itself gives an estimate for the upwelling below. Multiplying the resulting numbers by an along-slope distance of 100,000 km, about the length of the global oceanic ridge system, yields 16 Sv of downwelling and 18 Sv of upwelling. These numbers are not intended to be serious global estimates, which would have to take variations in dissipation rate, stratification, and slope into account; the multiplication by an along-slope length simply allows me to reason about more familiar numbers. The fact that the result has the same order of magnitude as Ferrari et al.’s estimate, however, suggests that the diabatic transports of the Brazil Basin may not be atypical.

Such a strong dipole in the diabatic transport would be impossible if the dynamics in abyssal mixing layers were one-dimensional. The analytical solution (7) shows that the downwelling transport in the outer layer would be $\kappa_1 \cot \theta \delta \sigma/(1 + \sigma \delta)$. Using the Brazil Basin parameters and again multiplying by an along-slope length of 100,000 km gives 0.2 Sv. The upwelling transport in the inner layer would be $\kappa_0 \cot \theta \delta \sigma/(1 + \sigma \delta) + \kappa_0 \cot \theta$, giving 3 Sv for Brazil Basin parameters. The dipole strength of 0.2 Sv is drastically weaker than the 16 Sv inferred from the observed dissipation rates. This is a direct consequence of the insufficient restratification by mean flows in one-dimensional dynamics: if there is only weak stratification that mixing can act on, the resulting diabatic buoyancy fluxes are reduced in magnitude, and the up- and downwelling dipole is weak.

In the three-dimensional dynamics, on the other hand, baroclinic restratification restores the estimate of the up- and downwelling dipole to what is inferred from observations. By laterally sliding dense water under light water, baroclinic eddies allow small-scale turbulence to act on stratified water. Given the same turbulent diffusivity, the enhanced stratification produces diabatic fluxes and thus water mass transformation that is much

**Fig. 8.** Evolution of the one-dimensional solution initialized with the observed stratification. The vigorous near-bottom mixing quickly erodes the stratification in the bottom few hundred meters if unopposed by baroclinic eddies.

**Fig. 9.** Evolution of the three-dimensional solution initialized with the observed stratification (experiment 2). (a) The time series of eddy kinetic energy (vertically integrated) show that finite-amplitude baroclinic eddies emerge just after $t = 500$ days. (b) The terms in the eddy kinetic energy budget show that kinetic energy is generated again by a slight dominance of buoyancy production over dissipation; shear production and hyperviscosity are small. (c) The stratification starts eroding initially, but once baroclinic eddies have emerged, the stratification is restored to and maintained near the realistic initial value. (d) The along-slope flow shows little change over the course of the simulation, remaining in approximate thermal wind balance.
larger than in the one-dimensional case. This suggests that an up- and downwelling dipole of a few tens of Sverdrups can be achieved if submesoscale baroclinic eddies counter the homogenizing effect of bottom-intensified mixing in abyssal mixing layers.

This result raises questions about the representation of the abyssal overturning circulation in global ocean models. These models have coarse resolution in both the horizontal and vertical (especially in the abyssal ocean) and thus heavily distort the dynamics of abyssal mixing layers. The above discussion suggests furthermore that using a realistically bottom-intensified mixing profile and improving the vertical resolution may not be sufficient to capture the strong up- and downwelling dipole. One may speculate that if topography (which is necessary smoothed, far-field stratification, and mixing coefficients vary only slowly in the horizontal (in a WKB sense), the dynamics will locally be similar to the one-dimensional dynamics discussed above (cf. Callies and Ferrari 2018b). Global models with high vertical but low horizontal resolution would thus suffer from the same problems as the one-dimensional solution. The stratification in abyssal mixing layers would be eroded by the bottom-intensified mixing, restratification by cross-slope mean flows would be weak, and no substantial up- and downwelling dipole would emerge. If baroclinic restratification is the dominant mechanism that maintains stratification, one will have to resolve or parameterize submesoscale baroclinic eddies in abyssal mixing layers in order to produce a realistic abyssal overturning circulation.

6. Discussion

The weak mean-flow restratification produced by one-dimensional dynamics and the strong baroclinic restratification produced by three-dimensional dynamics suggest a crucial role of submesoscale baroclinic eddies in the maintenance of mixing-layer stratification. It should be kept in mind, however, that I made a number of assumptions in the simplified cases considered above.

First, the turbulent momentum and buoyancy fluxes produced by internal wave breaking were represented with simple diffusive closures, and the turbulent Prandtl number was set to unity. It is neither clear that momentum fluxes generated by internal wave breaking are generally downgradient, nor that the turbulent viscosity should equal the turbulent diffusivity. The one-dimensional solution can be brought into much closer agreement with the observations if the Prandtl number is increased to order 100—an optimal fit is obtained with \( \tau = 230 \)—but it is unclear whether internal wave breaking can achieve such large momentum fluxes. It seems more likely that the momentum fluxes are enhanced by baroclinic eddies and that an increased Prandtl number amounts to a parameterization of these eddies (e.g., Rhines and Young 1982; Young 2012). But more work is needed to clarify this issue, for example, by explicitly representing the waves and their breaking (cf. Nikurashin and Legg 2011).

The diffusive dynamics also do not capture any logarithmic turbulent layers adjacent to the bottom (e.g., Townsend 1976). Since the focus here is on the turbulence generated aloft by internal wave breaking, this should not affect the conclusions. It should be kept in mind, however, that the diffusive parameterization glosses over our ignorance about the turbulence in the bottom few tens of meters.

Another consequence of the diffusive parameterization is that feedbacks between the subinertial dynamics considered here and the waves and turbulence that produce the mixing are excluded. In reality, the wave generation and breaking would itself be strongly modulated by the stratification in the mixing layer. The above cases with fixed mixing coefficients are thus best thought of as consistency tests: Are the subinertial dynamics induced by the turbulence capable of maintaining the observed mean state that the turbulence generation depends on? That this test is passed only if submesoscale eddies are present is strongly suggestive of their importance, but the full dynamics with interactive waves, turbulence, and baroclinic eddies should be considered in future work.

Second, the solutions considered above take into account the large-scale topographic slope only. The small-scale topography—that is, the abyssal hills and fracture zones that are responsible for the bottom-enhanced turbulence in the first place—entered the calculations only indirectly through the bottom-enhanced mixing coefficients. These topographic features may steer mean flows and thus change their ability to maintain stratification. For example, the walls of fracture zone canyons may block flows along the large-scale slope of the Mid-Atlantic Ridge (e.g., Thurnherr and Speer 2003). A rough estimate of this blocking effect can be obtained by setting \( v = 0 \) or equivalently \( f = 0 \); that is, by reverting to nonrotating dynamics (Dell 2013). This would amount to \( S\alpha(1 + S\alpha) \to 1 \) and thus yield a dramatic increase in stratification and water mass transformation (cf. Clément and Thurnherr 2018). It remains unclear, however, how strong a stratification can be maintained in the presence of realistically bottom-intensified mixing both within and outside of canyons and over more general irregular topography. The

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\[^{6}\] This fit was obtained using numerical solutions because the assumption \( qh \gg 1 \) breaks down for large Prandtl numbers.
topography may also modulate baroclinic instability (e.g., de Szoerce 1983) and the resulting baroclinic eddies, effects that should also be addressed by future work.

Third, the setup in a sloping coordinate system with a prescribed far-field stratification restricts attention to the local dynamics. Nonlocal effects, such as changes in the large-scale topographic slope, varying far-field stratification, or the change of the inertial frequency with latitude, are all neglected. Such variations can cause convergences and divergences of the cross-slope mean flow and thus give rise to exchange with the far field (e.g., Phillips et al. 1986; McDougall 1989; Kunze et al. 2012; Dell and Pratt 2015; Holmes et al. 2018). Leading-order changes to the mean-flow restratification, however, are not expected: if the variations are slow, a WKMB approach suggests that the leading-order dynamics are still one-dimensional locally, and mean-flow restratification should remain weak.

Beyond exploring the above questions in more realistic numerical simulations, one should look for evidence of submesoscale baroclinic eddies in observations. Moored current meters deployed in a Brazil Basin fracture zone canyon do show substantial subinertial variability that is bottom intensified (Thurnherr et al. 2005; Toole 2007; Clément et al. 2017), as expected for baroclinic eddies generated in mixing layers. But a careful analysis and possibly more observations are needed to discriminate against other sources of subinertial variability.

If the local dynamics on a planar large-scale slope capture the essence of abyssal mixing layers, effective baroclinic restratification should generally appear in abyssal mixing layers. The problem is determined completely by the following set of nondimensional parameters: the slope Burger number $S$, the Ekman number $E = \nu \ell / f h^2$ based on the decay scale of mixing $h$, the bottom-enhancement ratio $k_1 / k_0$, and the turbulent Prandtl number $\sigma$. In the abyssal ocean, where stratification is relatively weak and large-scale topographic slopes are gentle, $S \ll 1$ is typical (except very close to the equator). When turbulence is produced by the breaking of topographically generated internal waves, it typically extends far beyond the Ekman layer $(|E|^{1/2} \ll 1)$ and is much stronger than in the far field $(k_1 \gg k_0)$; e.g., Waterhouse et al. 2014. The calculations presented above are in this typical regime: $S = 2.2 \times 10^{-3}$, $|E| = 6.5 \times 10^{-4}$, $k_1 / k_0 = 35$, and $\sigma = 1$. For baroclinic instabilities to exist, it is crucial that the outer layer predicted by one-dimensional dynamics is much thicker than the inner layer, which requires $\log(k_1 / k_0) \gg |E|^{1/2} (1 + S \sigma)$, and that the outer-layer stratification predicted by one-dimensional theory is much weaker than in the far field, which is the case if $S \sigma \ll 1$ and $k_1 / k_0 \gg 1$. These conditions are typically met where there is strong topographically induced mixing, so baroclinic restratification should be widespread in abyssal mixing layers.

7. Conclusions

The results of this paper raise the possibility that submesoscale baroclinic eddies are instrumental in maintaining stratification in abyssal mixing layers against the homogenizing tendency of bottom-intensified mixing. The restratification of abyssal mixing layers is crucial because it allows the topographically induced turbulence to continually achieve water mass transformation and produce a diapycnal up- and downwelling dipole of a few tens of Sverdrups globally. Baroclinic restratification may thus help shape the overturning circulation of the abyssal ocean.

Considering the local dynamics of abyssal mixing layers on slopes, this paper revealed a stark difference between one- and three-dimensional dynamics. If the flow is one-dimensional, that is, invariant in the plane of the slope, the homogenizing tendency of mixing is countered by cross-slope mean flows only. With parameters appropriate for the Brazil Basin, one-dimensional solutions produce mean flows that are much too weak to maintain the observed stratification. But the weakly stratified mixing layers produced by such one-dimensional dynamics are unstable to baroclinic perturbations. In three-dimensional dynamics, submesoscale baroclinic eddies emerge and restratify mixing layers by laterally sliding dense water under light water. These eddies can effectively maintain the observed stratification.

While baroclinic restratification is largely adiabatic and does not cause water mass transformation directly, it does modify the stratification that is mixed up by small-scale turbulence. Given the same mixing coefficients, the enhanced stratification allows for much stronger diabatic fluxes and thus water mass transformation. The up- and downwelling dipole thought to be generated by bottom-intensified mixing would be much weaker if the mixing-layer dynamics were one-dimensional and mixing layers were largely unstratified. Baroclinic restratification in abyssal mixing layers could therefore directly affect the large-scale overturning circulation of the abyssal ocean.

Global models typically do not resolve these submesoscale instabilities in the abyss because of a lack of both horizontal and vertical resolution. If the abyssal overturning circulation is indeed affected by submesoscale baroclinic restratification, incorporating its effects into global models is imperative. This could be done either by drastically increasing the resolution or by devising a parameterization similar to that of submesoscale baroclinic restratification in the surface mixed layer (Fox-Kemper et al. 2011).
Acknowledgments. I would like to thank the developers of Dedalus (dedalus-project.org), which was used for all computations presented in this paper. In particular, I would like to thank Keaton Burns, who also assisted with the setup and optimization of the computations. Chris Garrett and Carl Wunsch gave valuable feedback on the manuscript, and anonymous reviewers provided useful suggestions.

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